



CMSSM

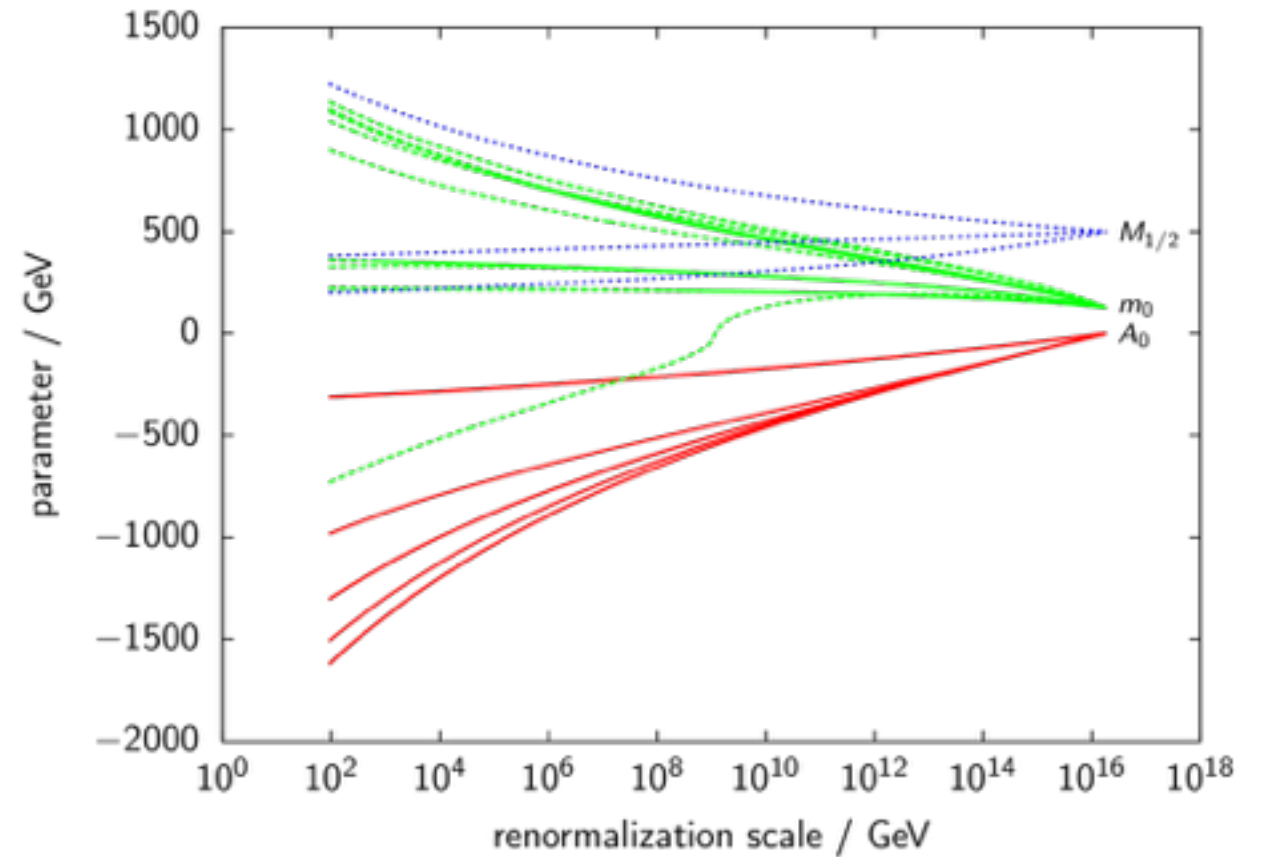
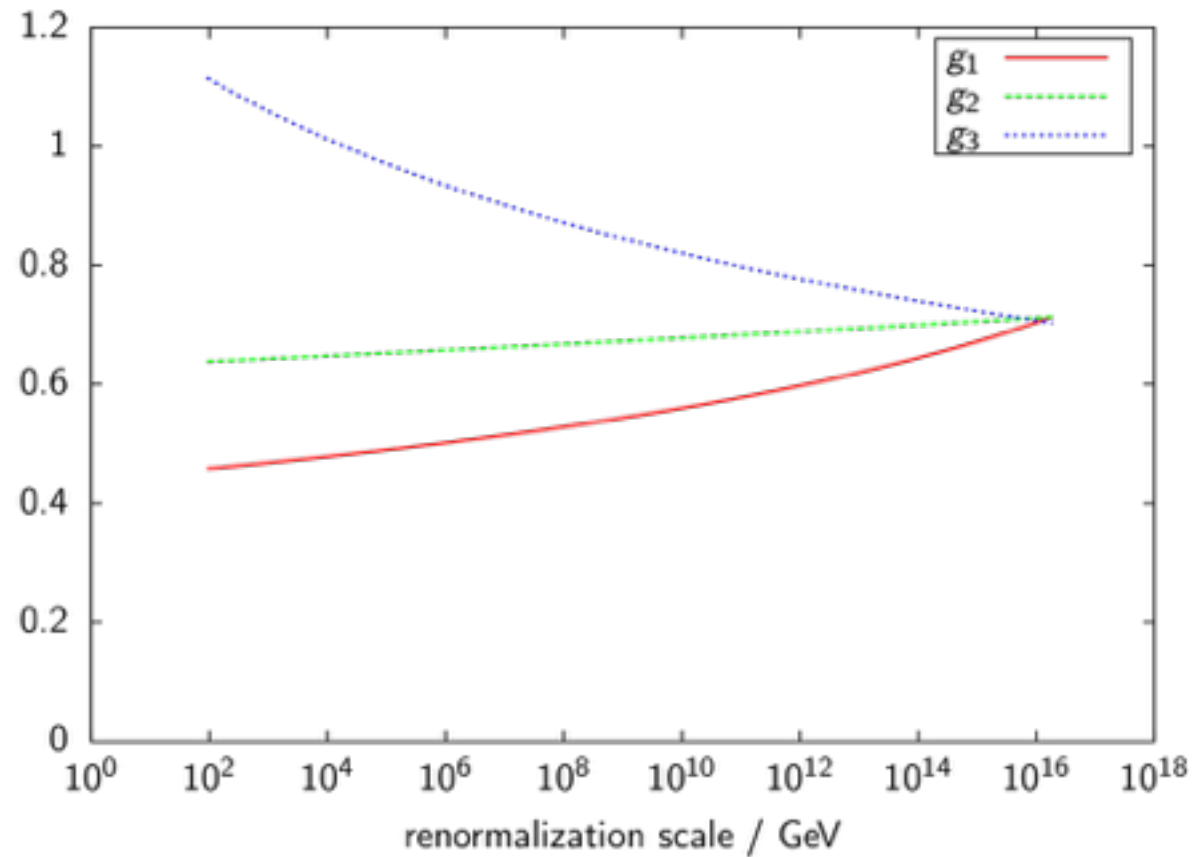
In the light of 13 TeV LHC and LUX WS2014-16

张阳

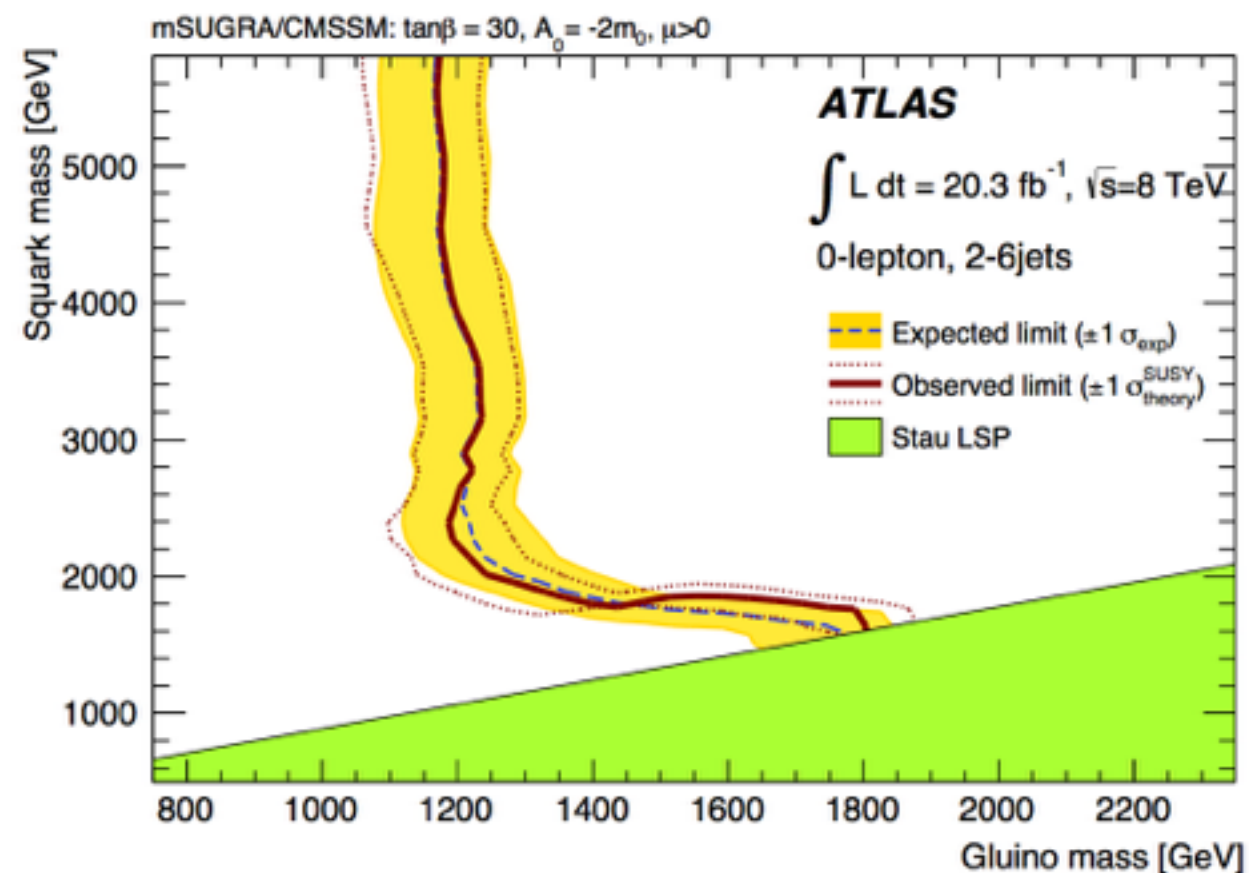
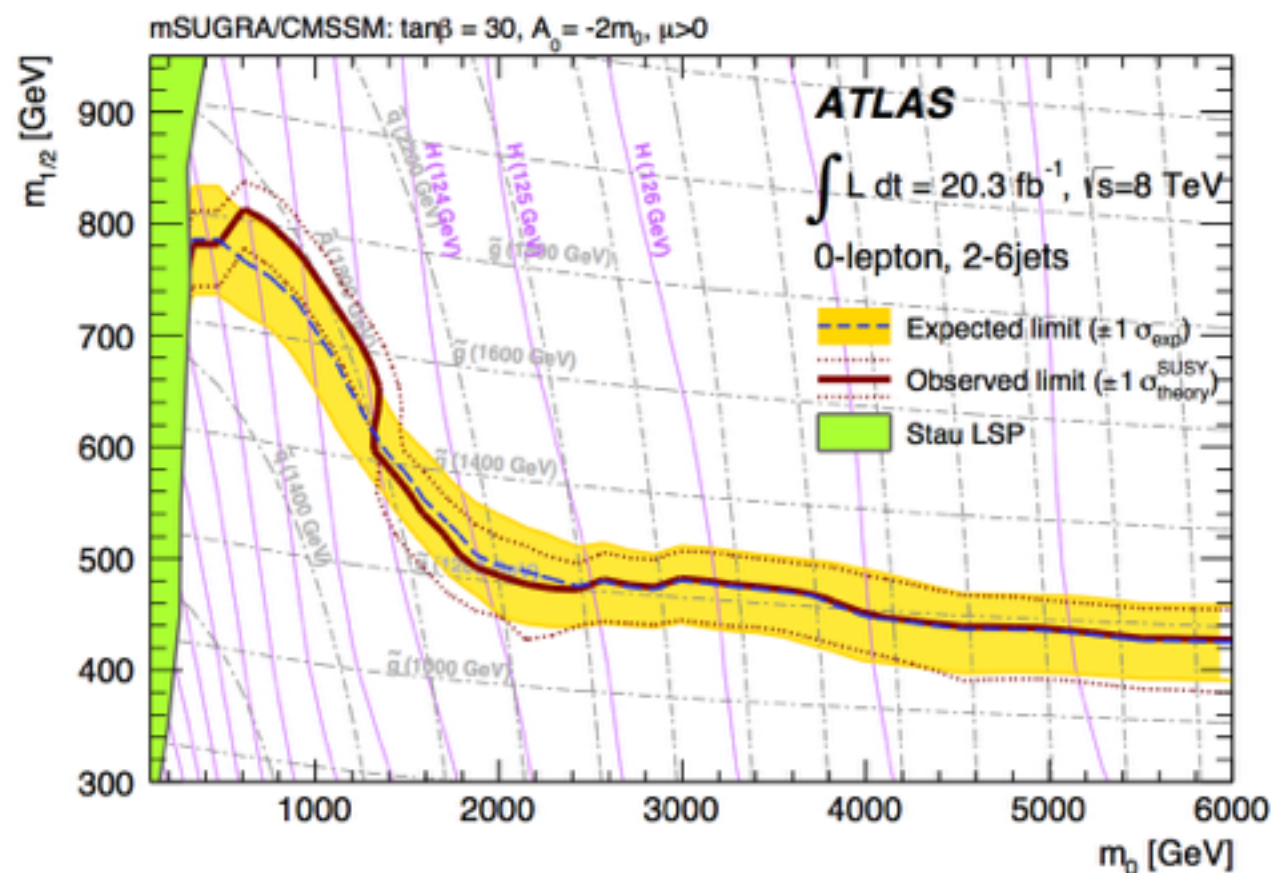
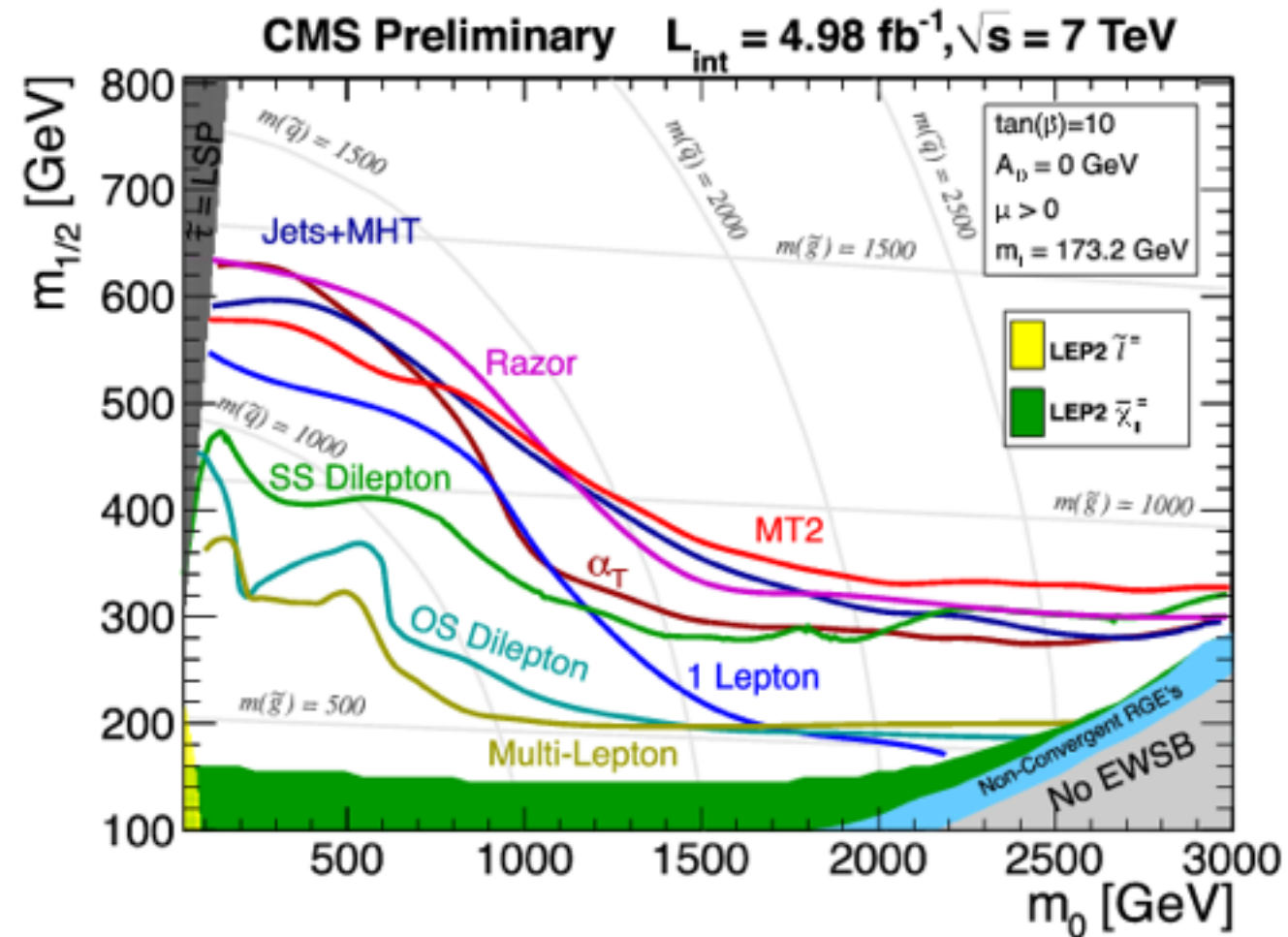
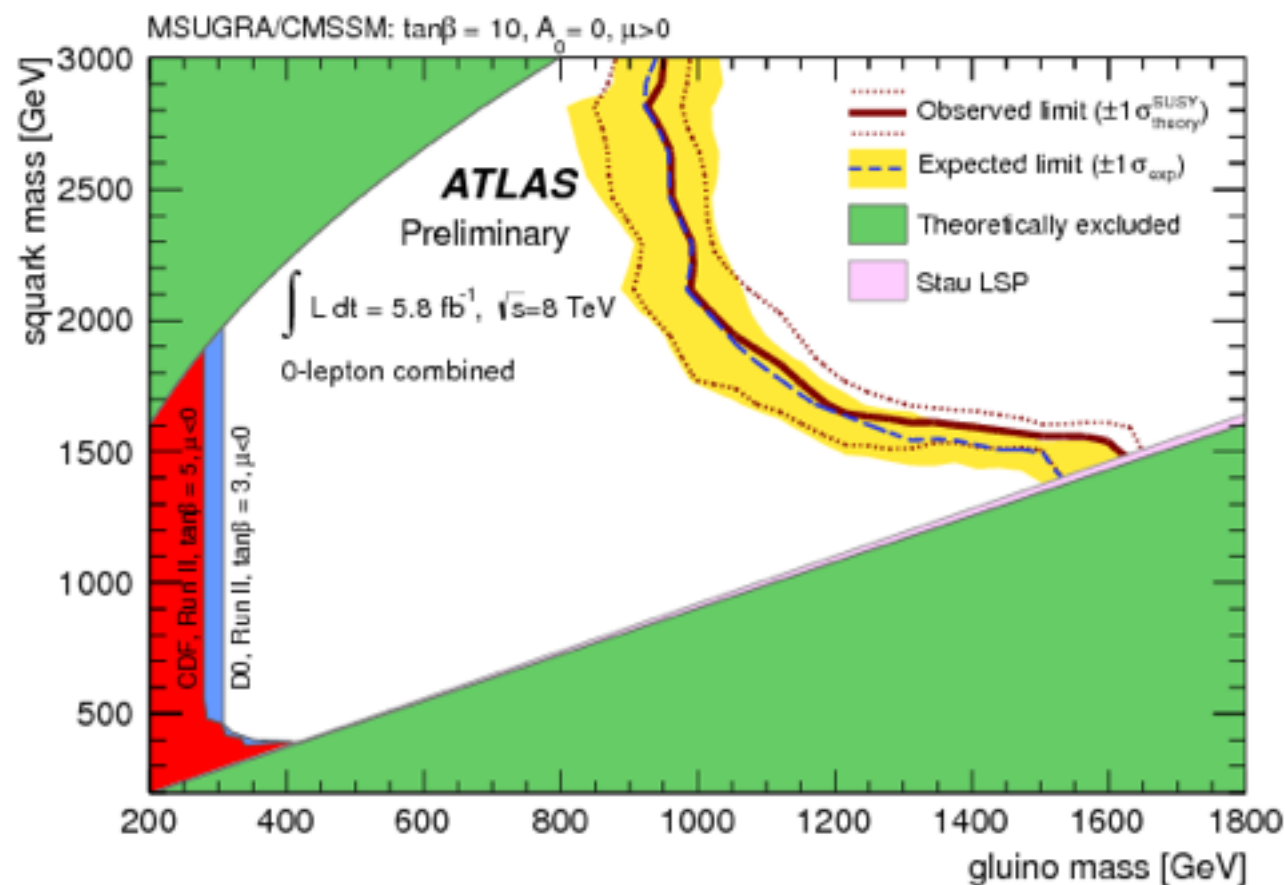
中科院理论物理所

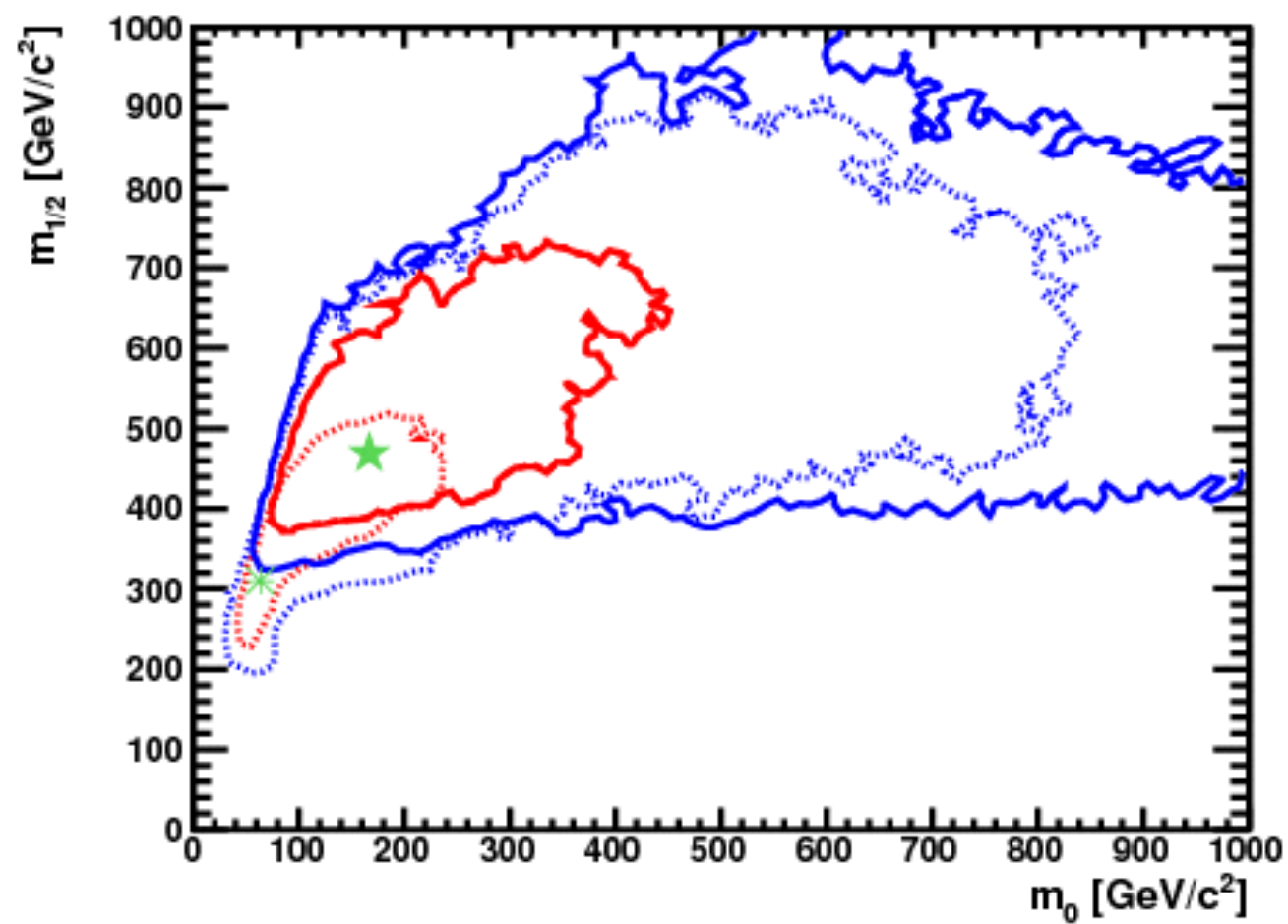
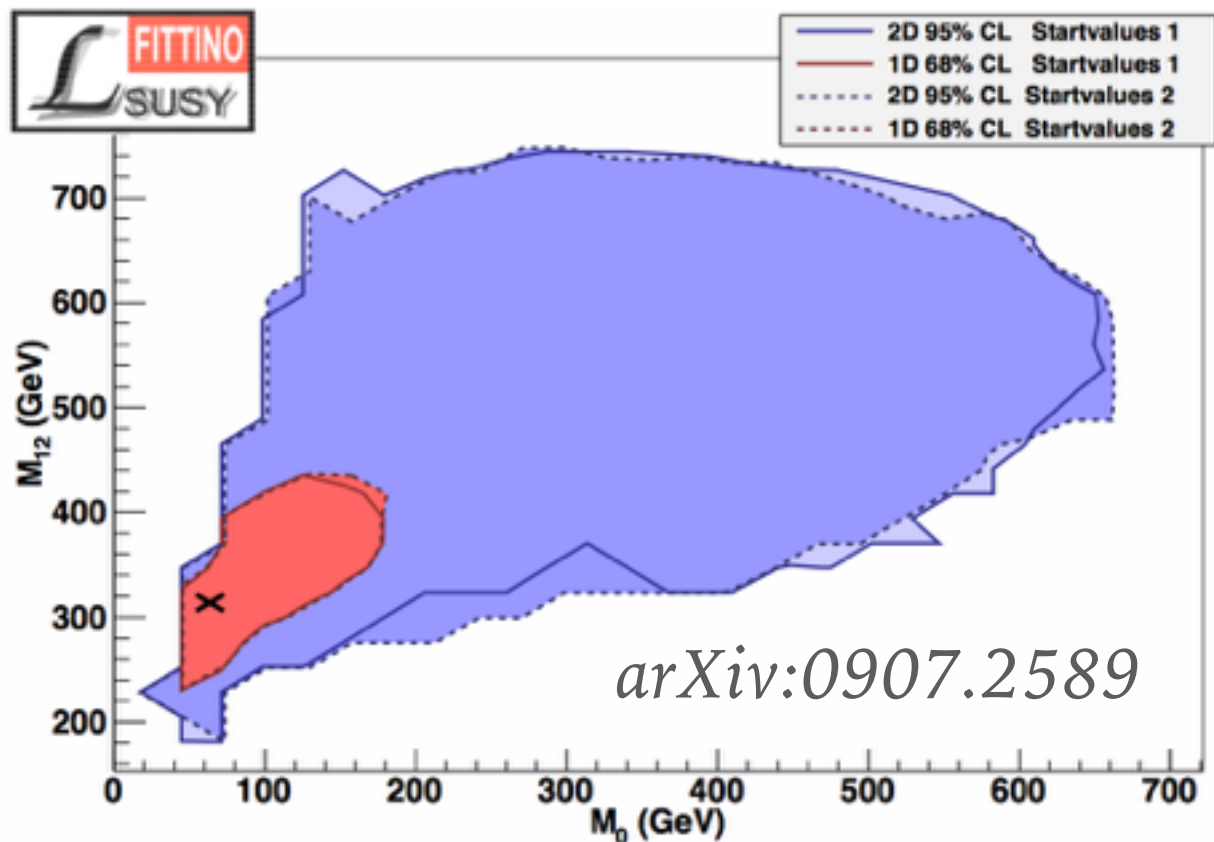
2016.09.23 郑州大学

INTRODUCTION



The CMSSM is famous for the small number of parameters, only including four and 'half' parameters at Grand Unification(GUT) scale, the universal scalar and gaugino mass parameters M_0 and $M_{1/2}$, the universal trilinear coupling A_0 , the ratio of two Higgs doublet vacuum expectation value $\tan\beta$, and the sign of the Higgs/Higgsino mass parameter $\text{sign}(\mu)$.



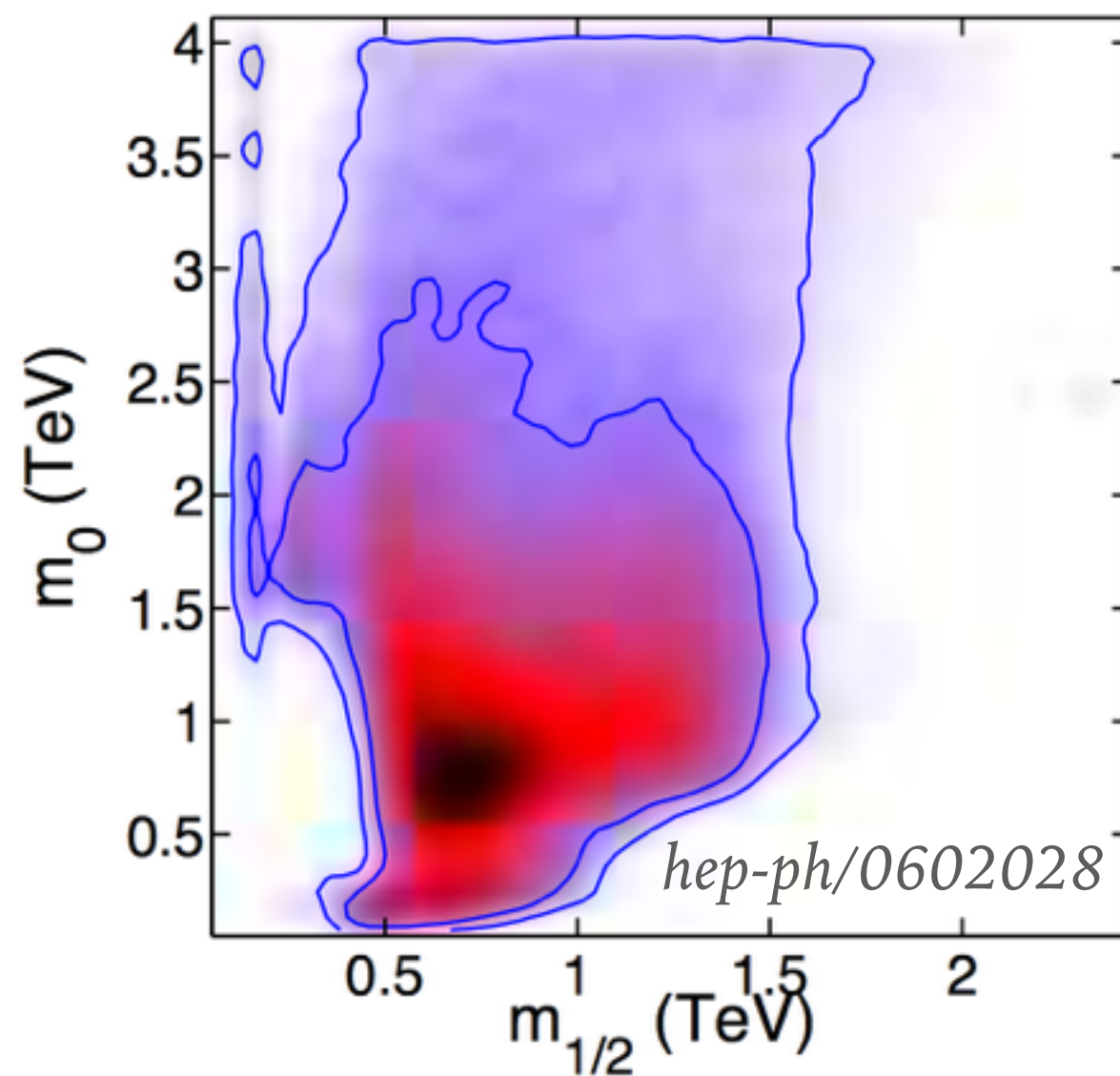


SuperBayes/BayesFit

MasterCode

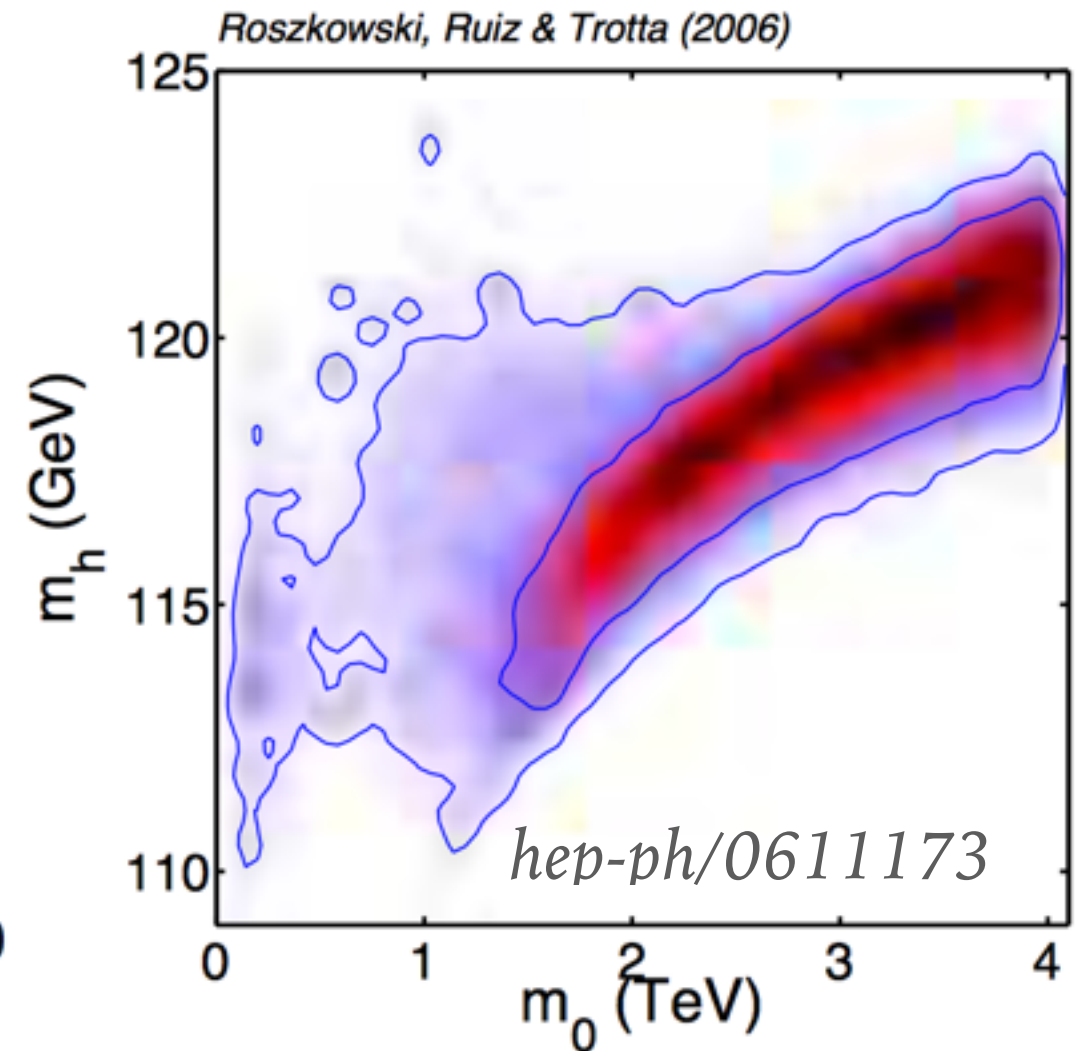
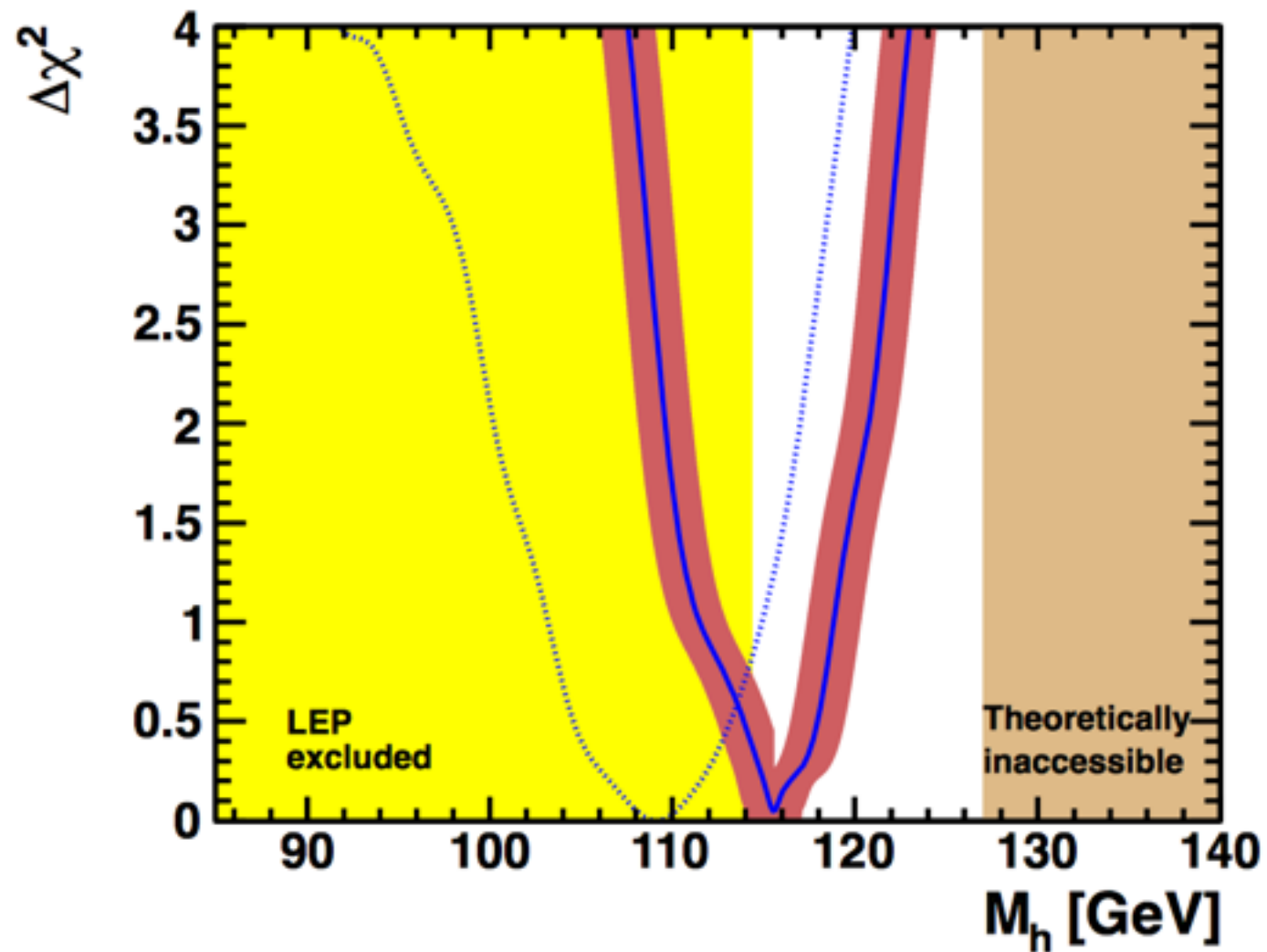
Fittino

Ruiz, Trotta & Roszkowski (2006)

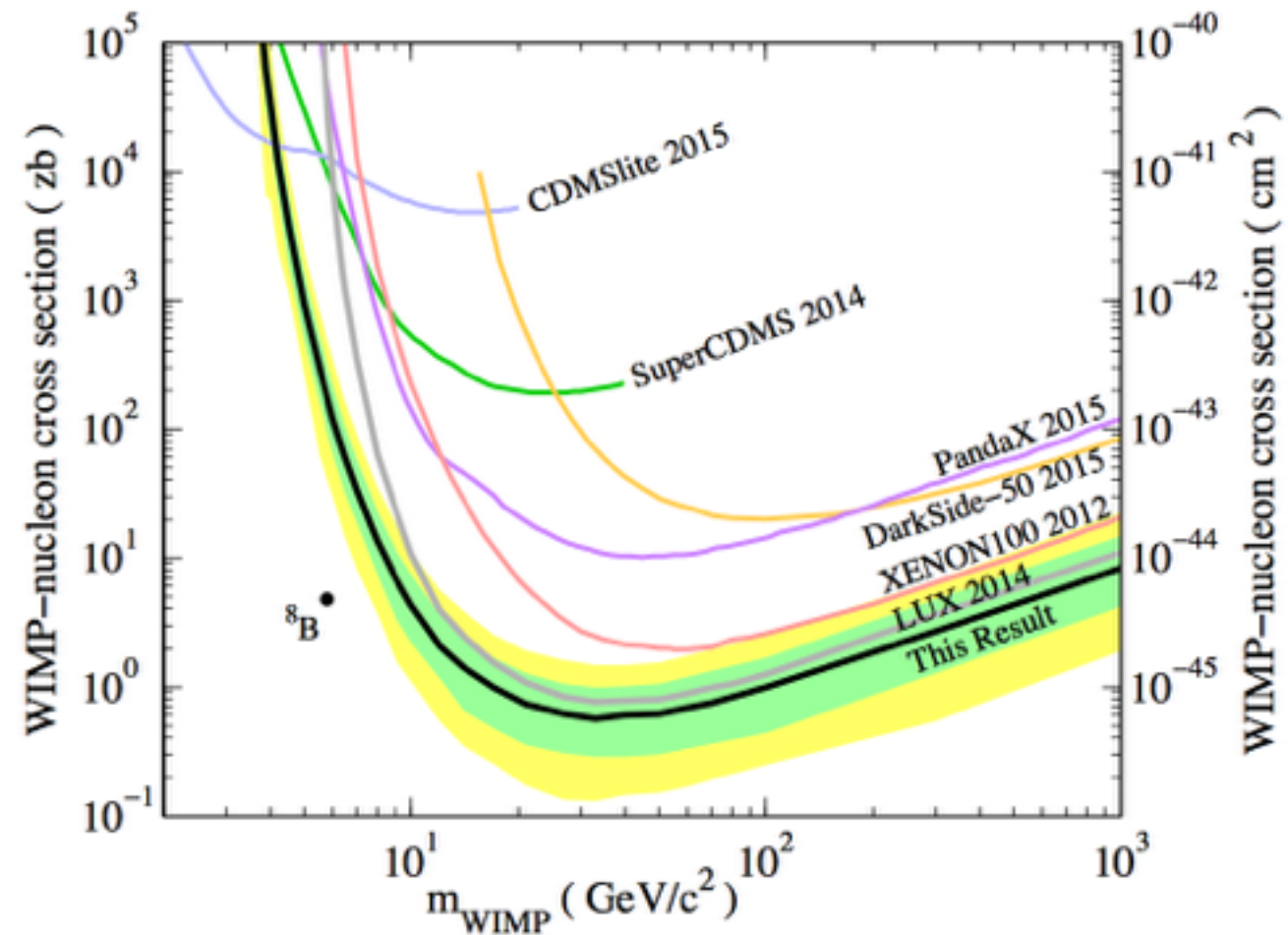
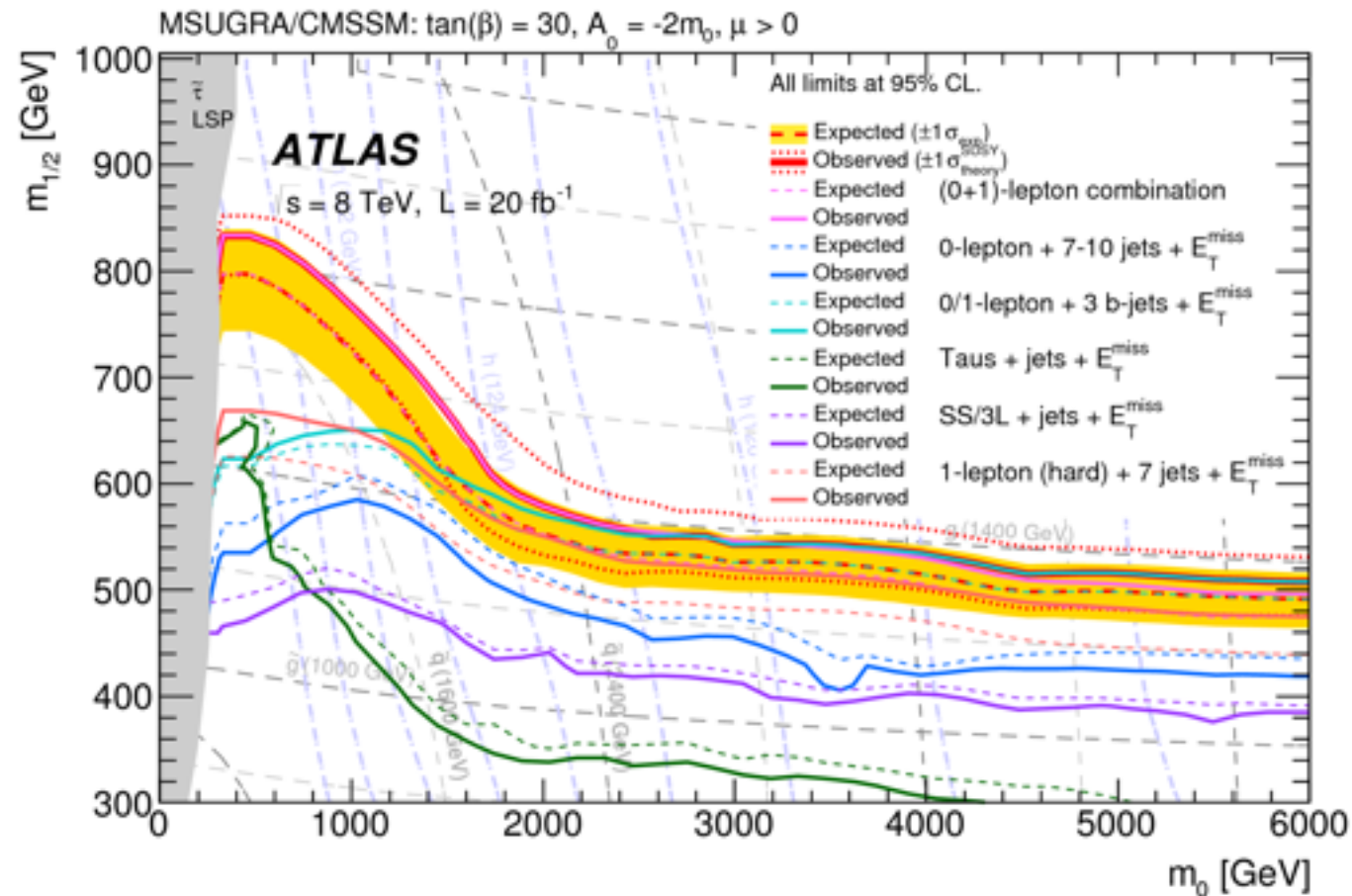


Then the 125 GeV Higgs is discovered at LHC ...

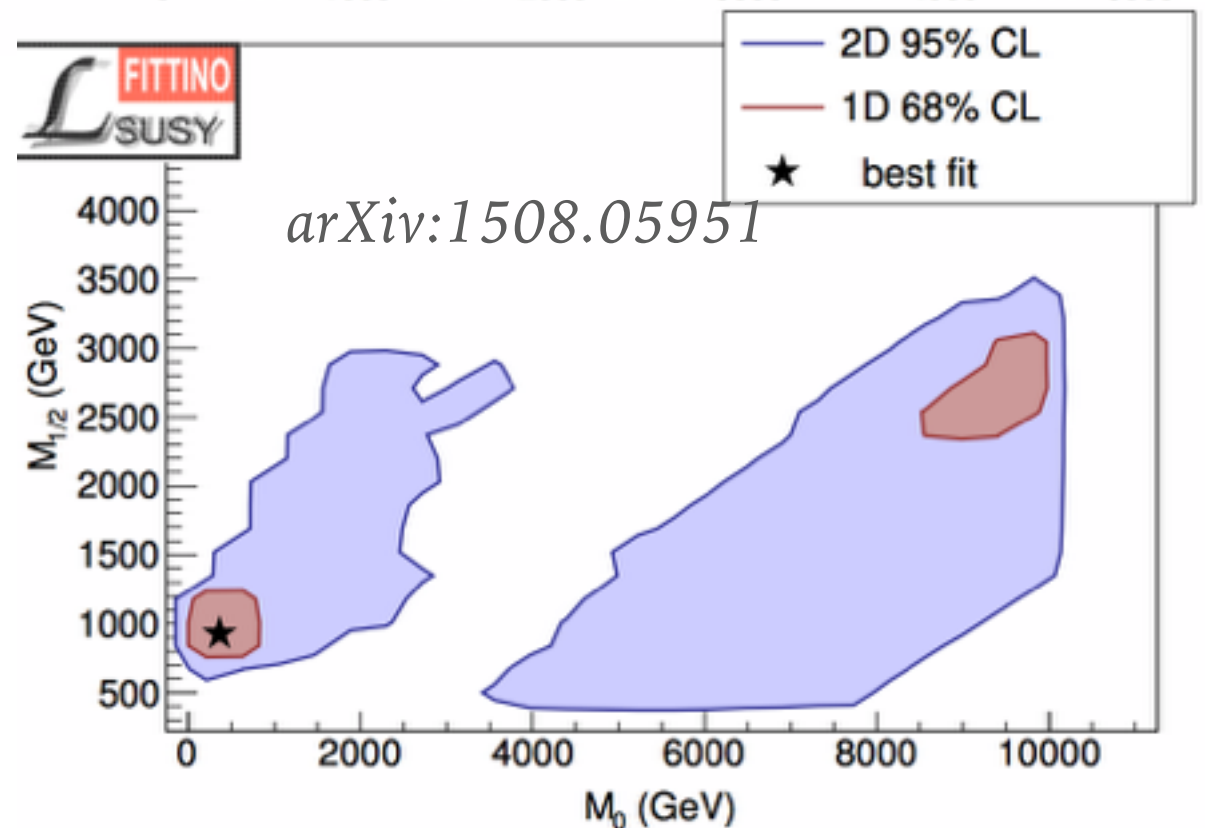
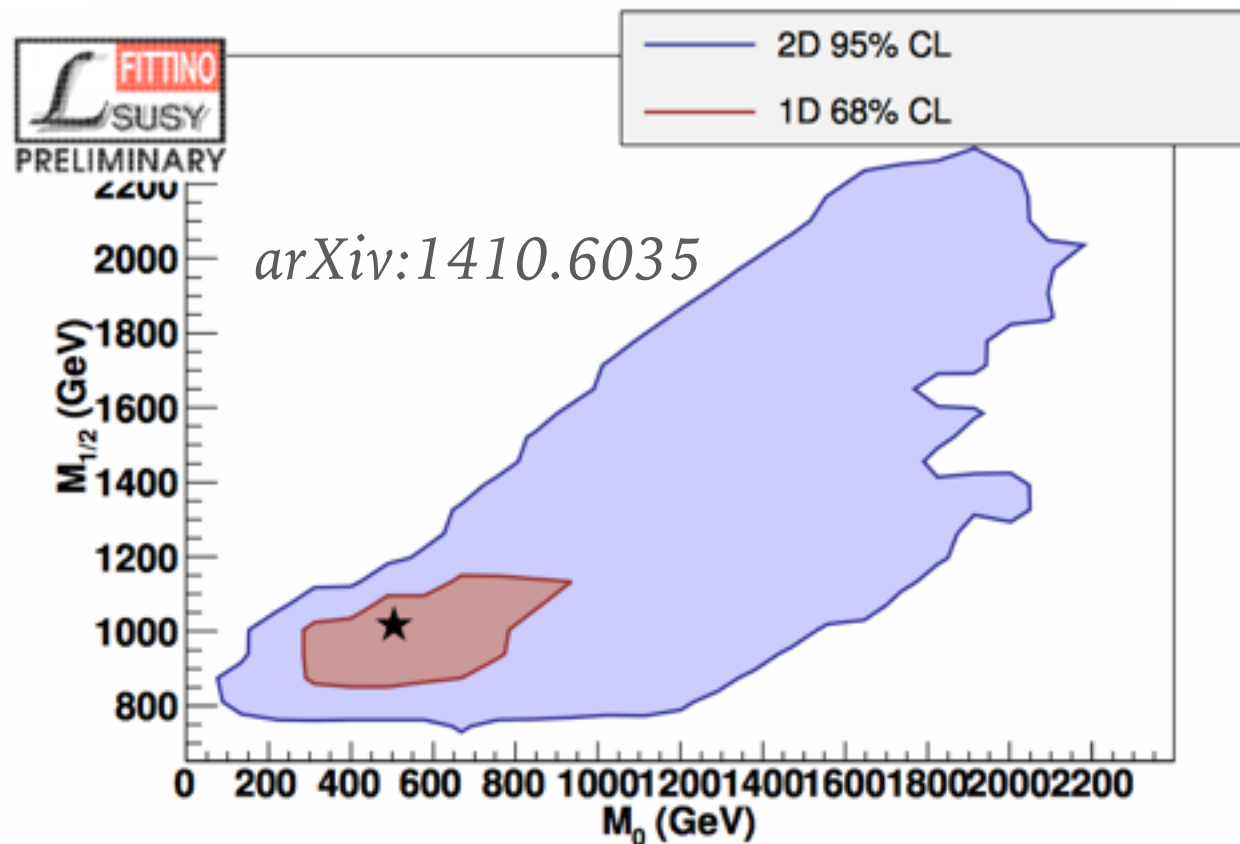
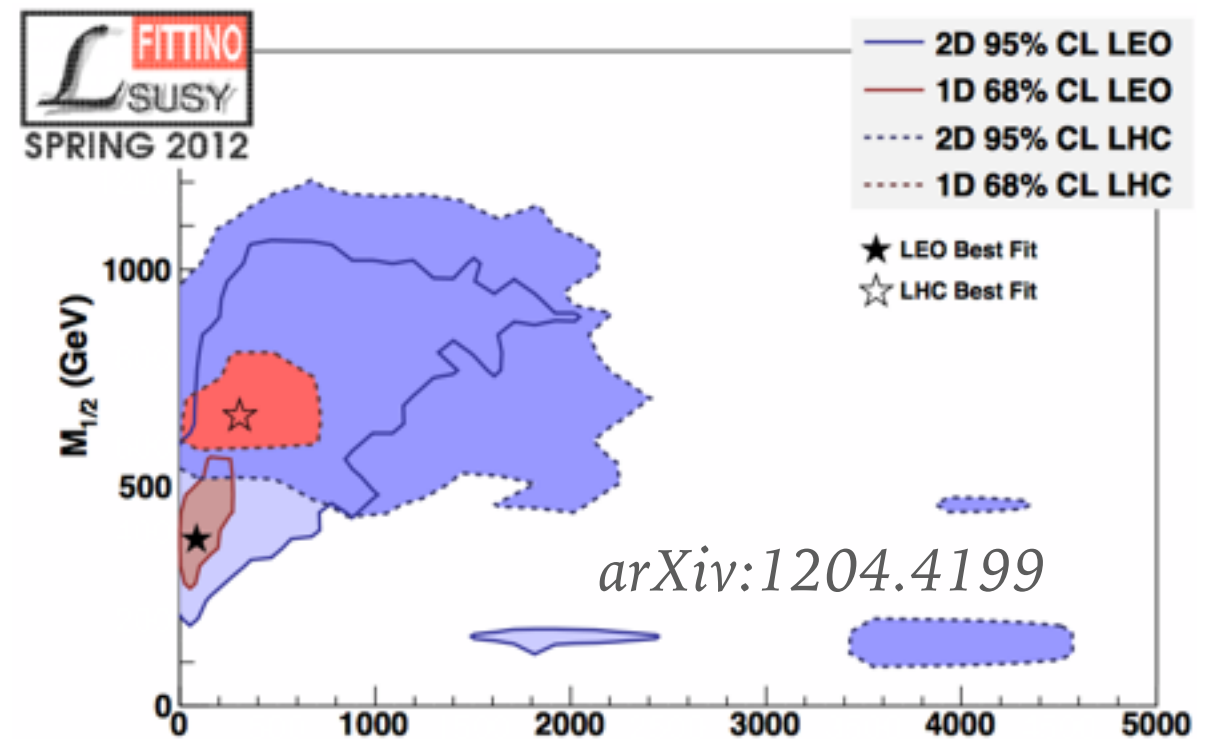
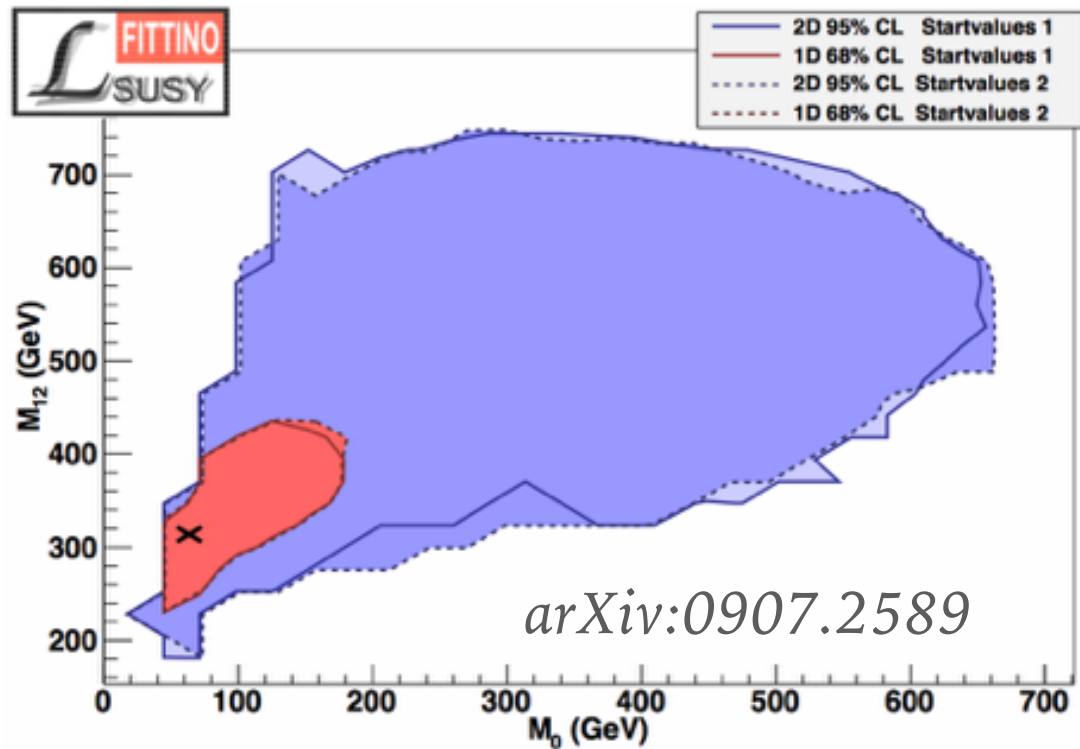
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And no significant signals have been observed in SUSY searches at LHC and DM direct searches.

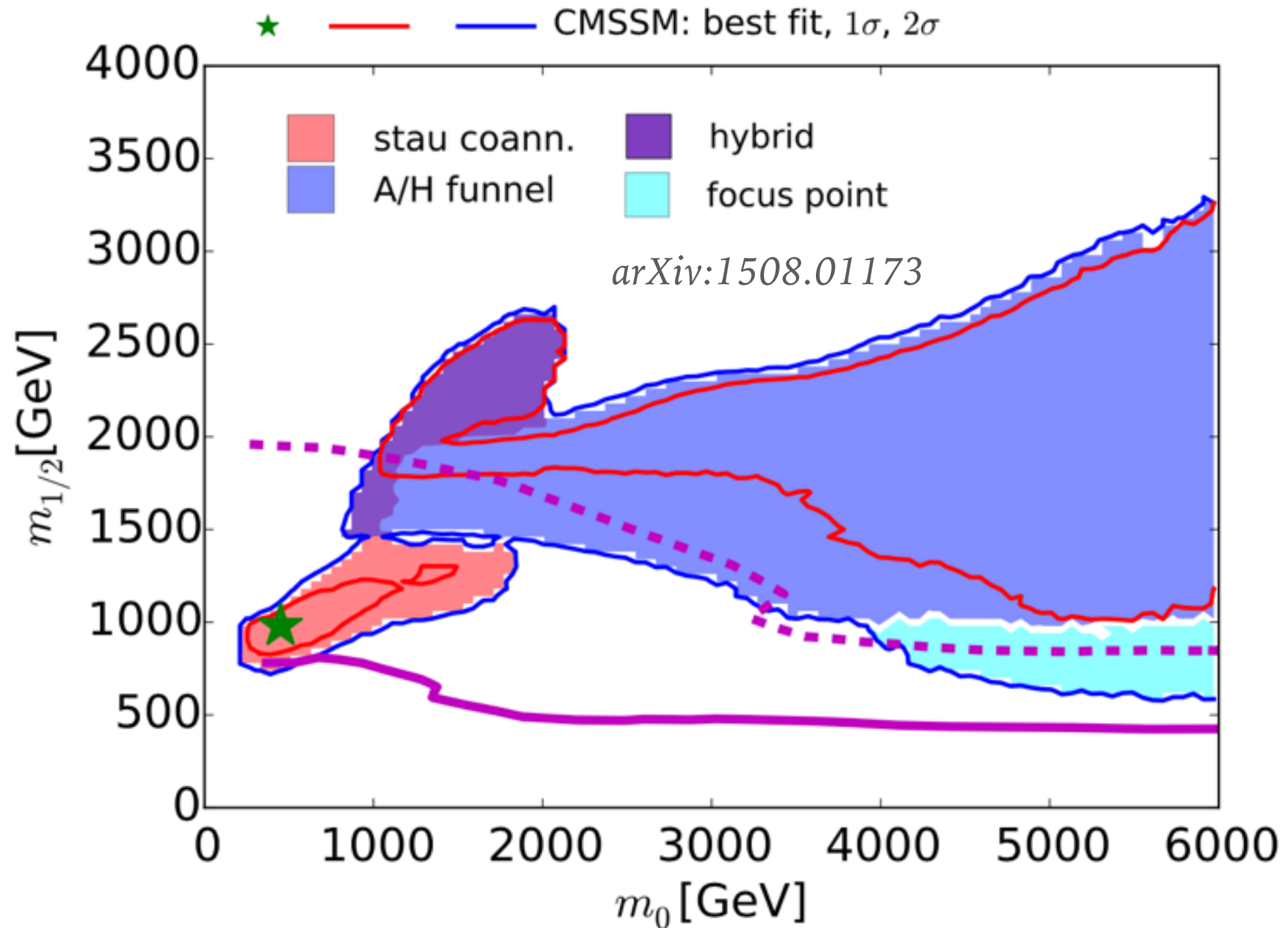


The favored parameter space changed ...

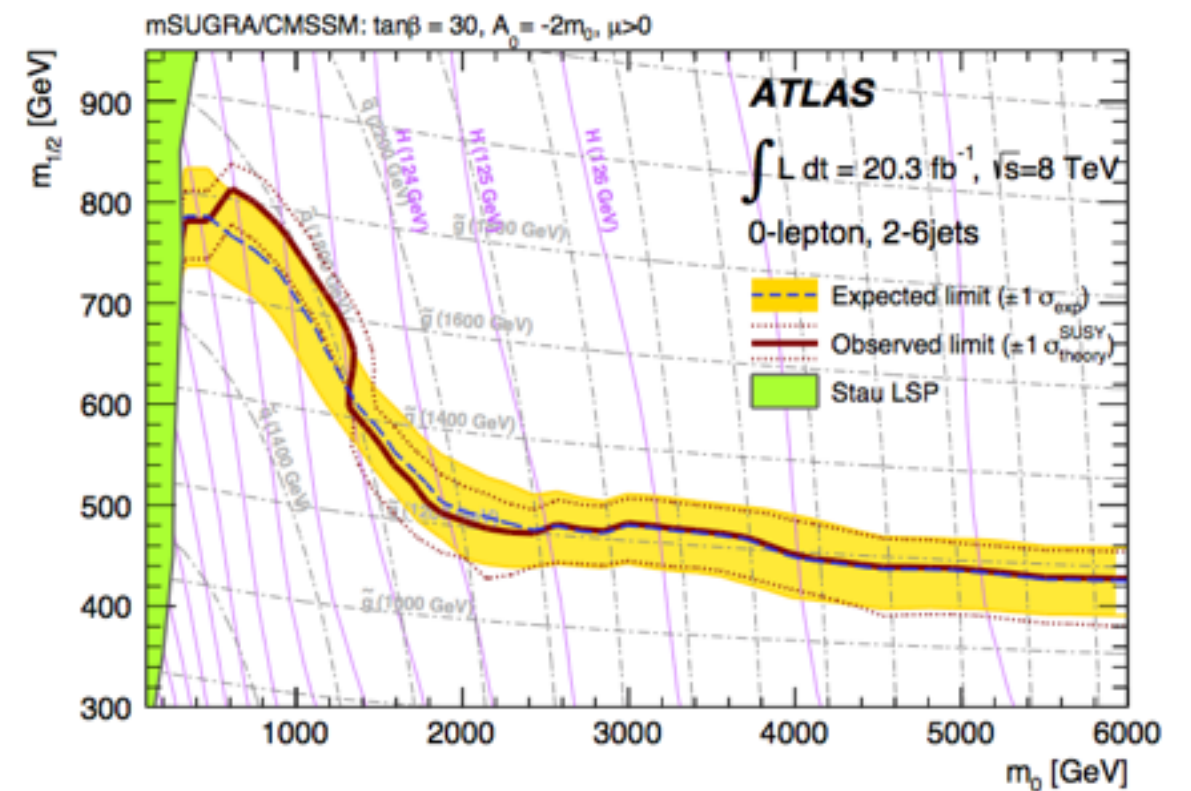
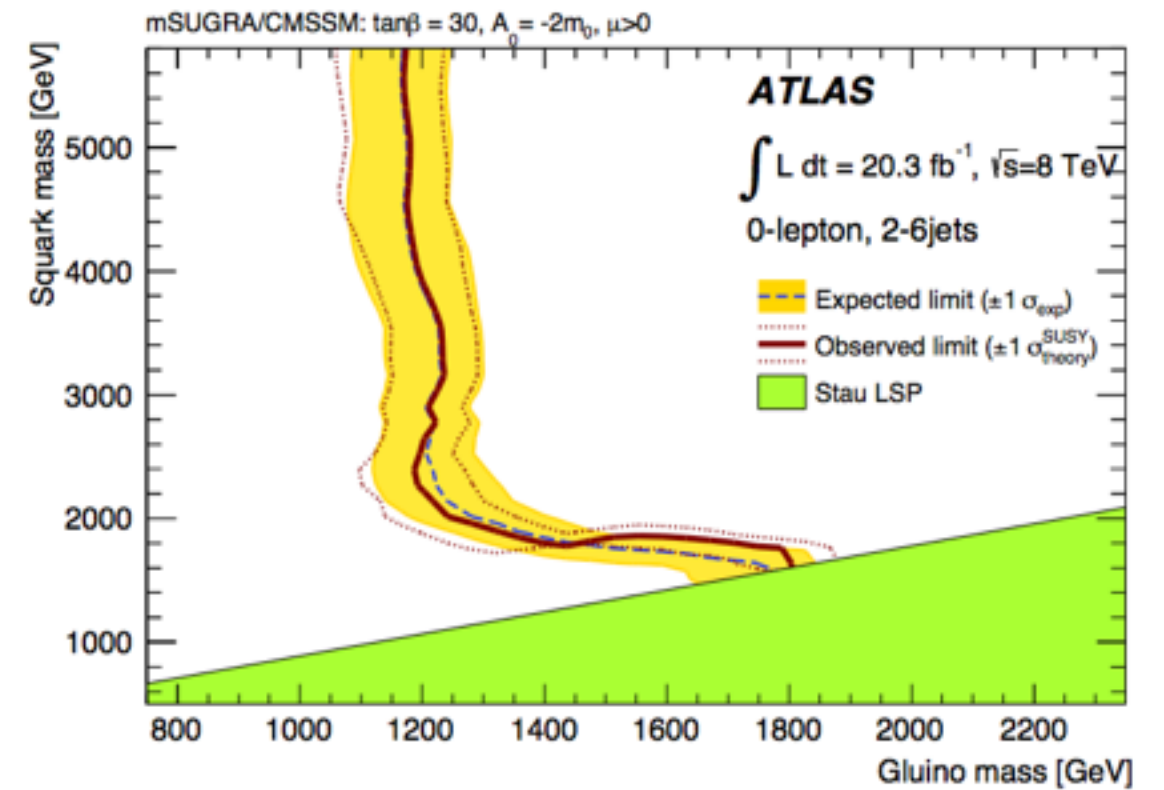
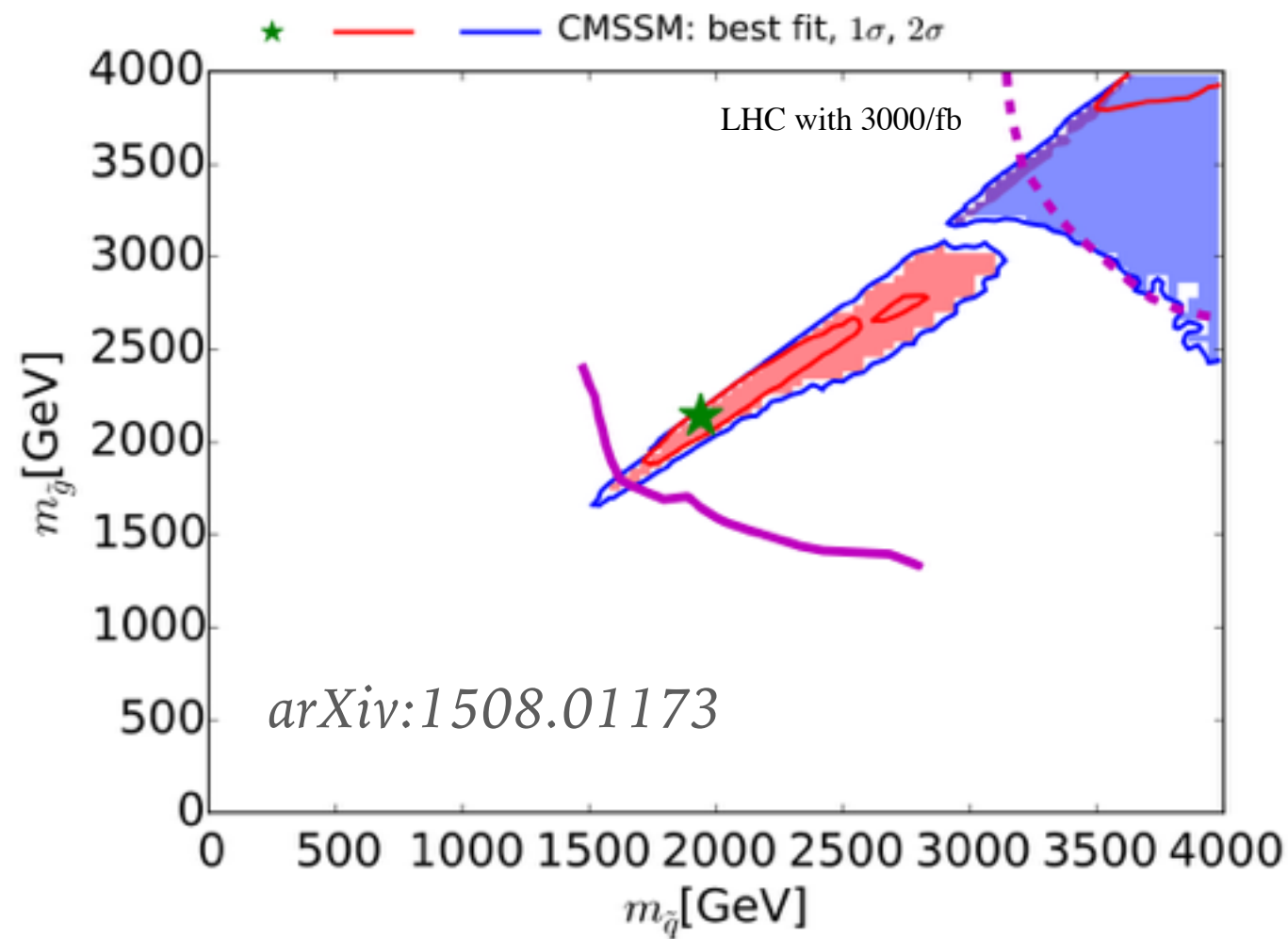


Dark matter annihilation mechanisms

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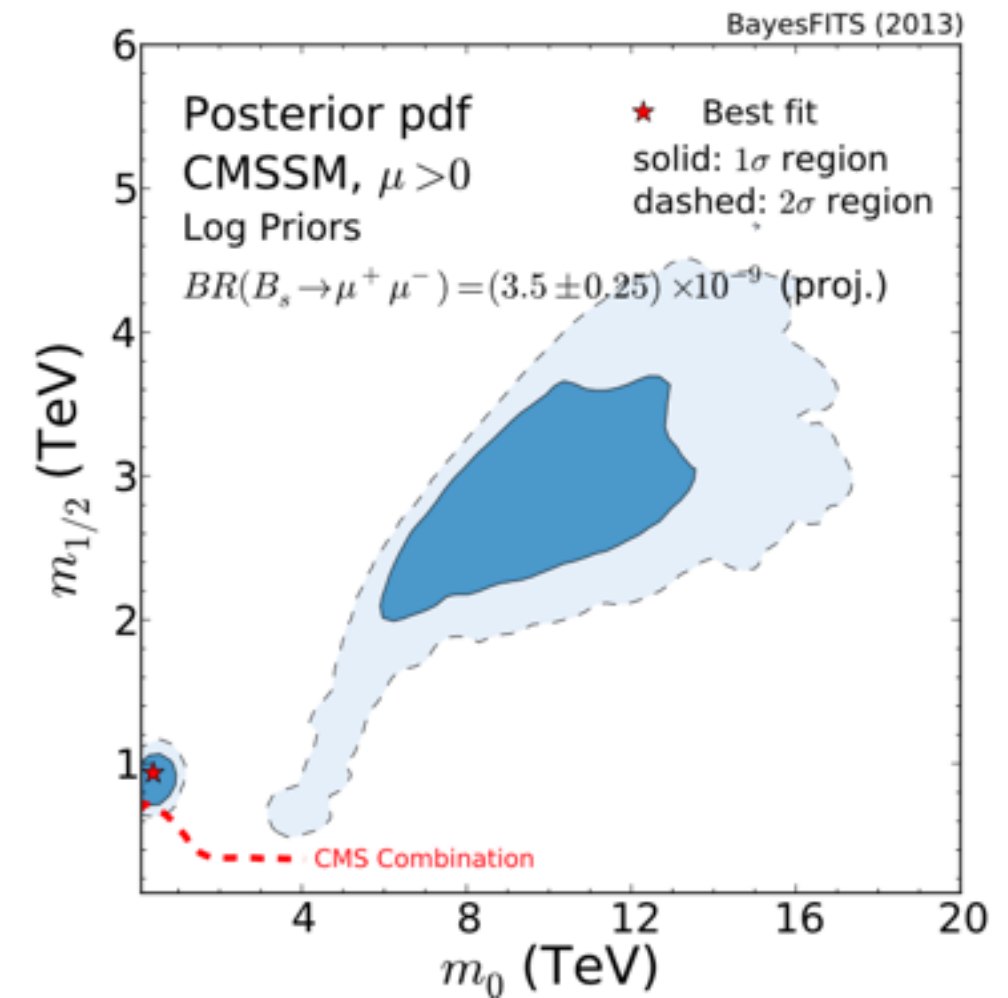
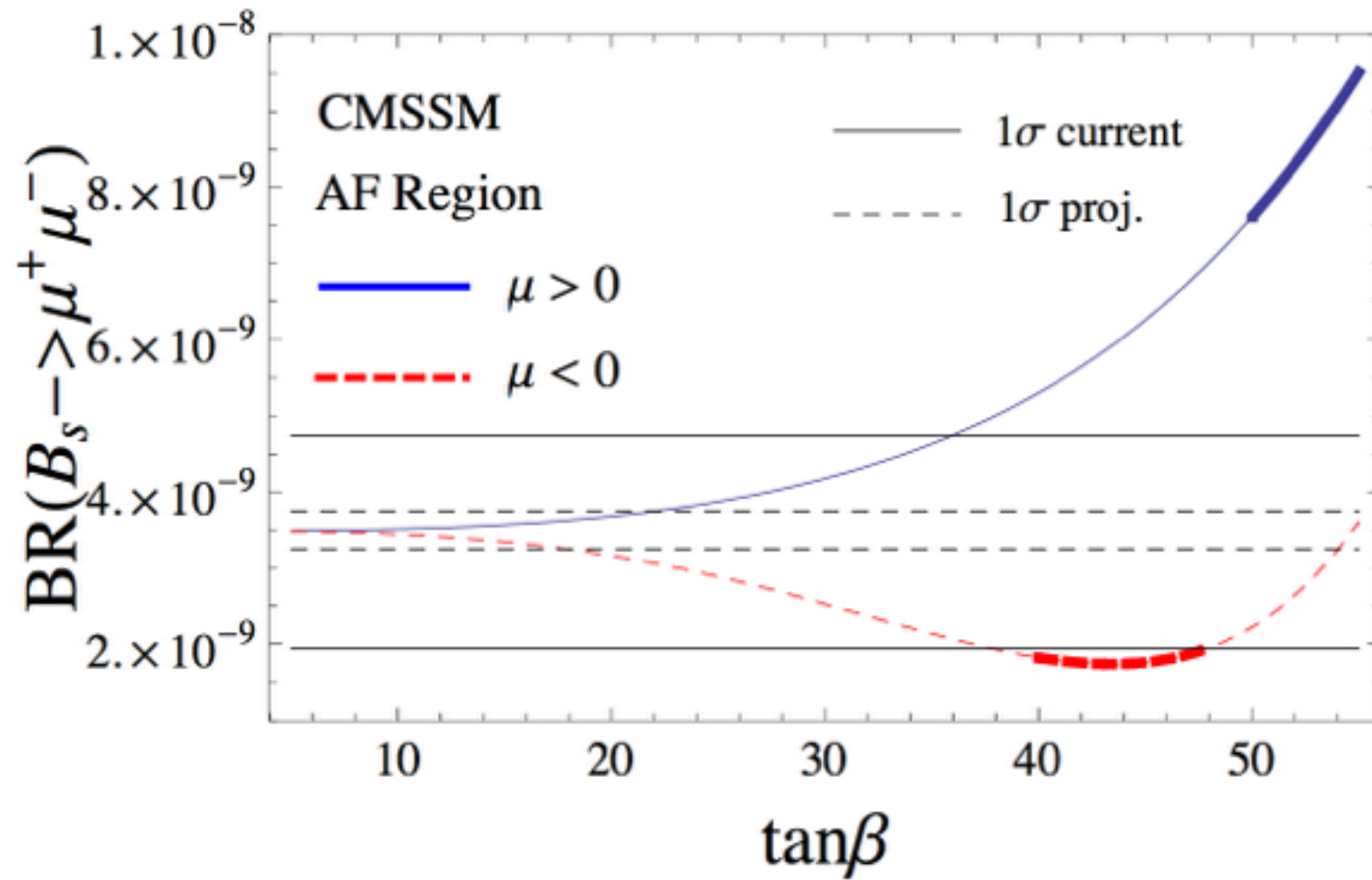


Stau coannihilation and LHC



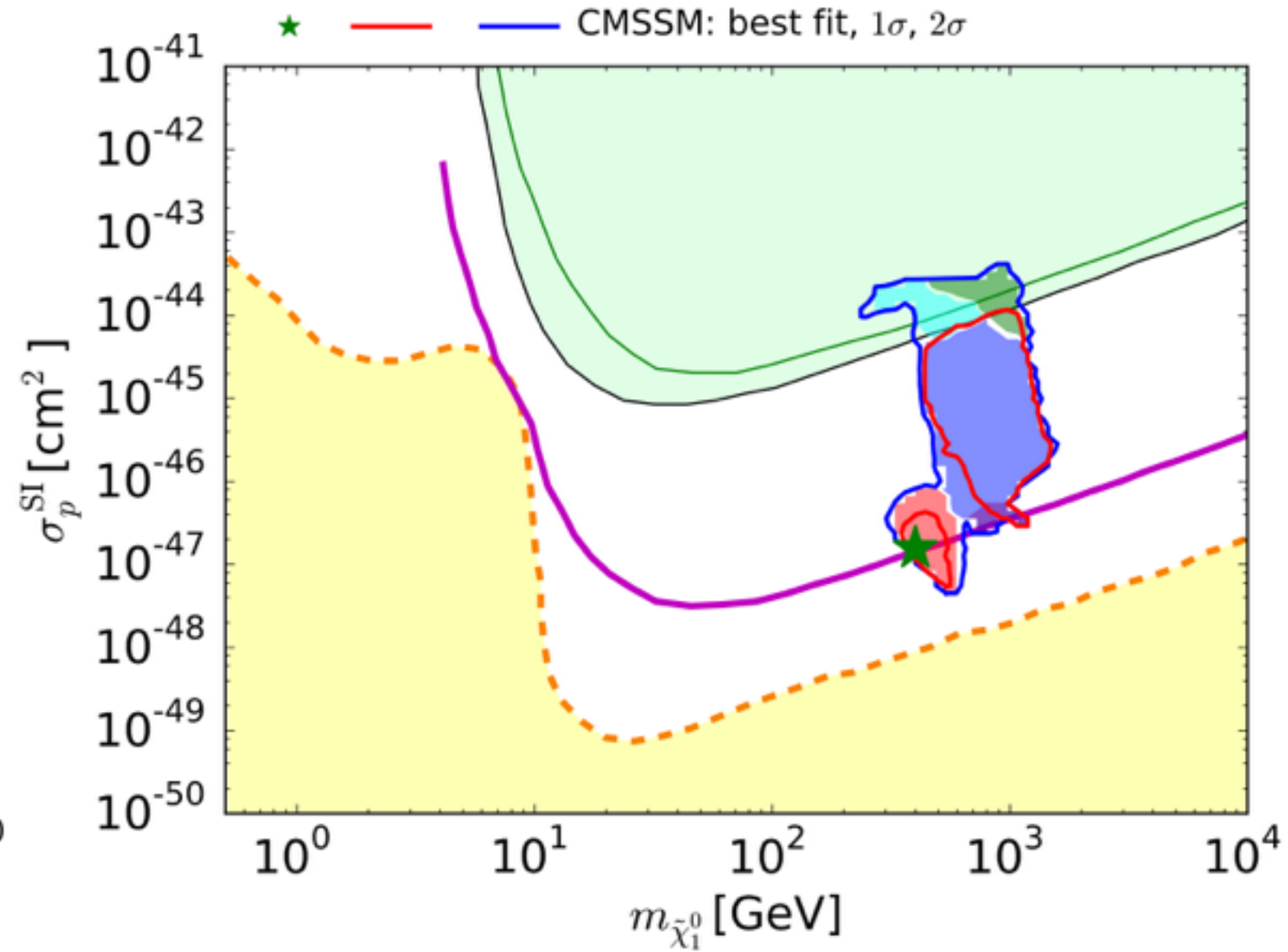
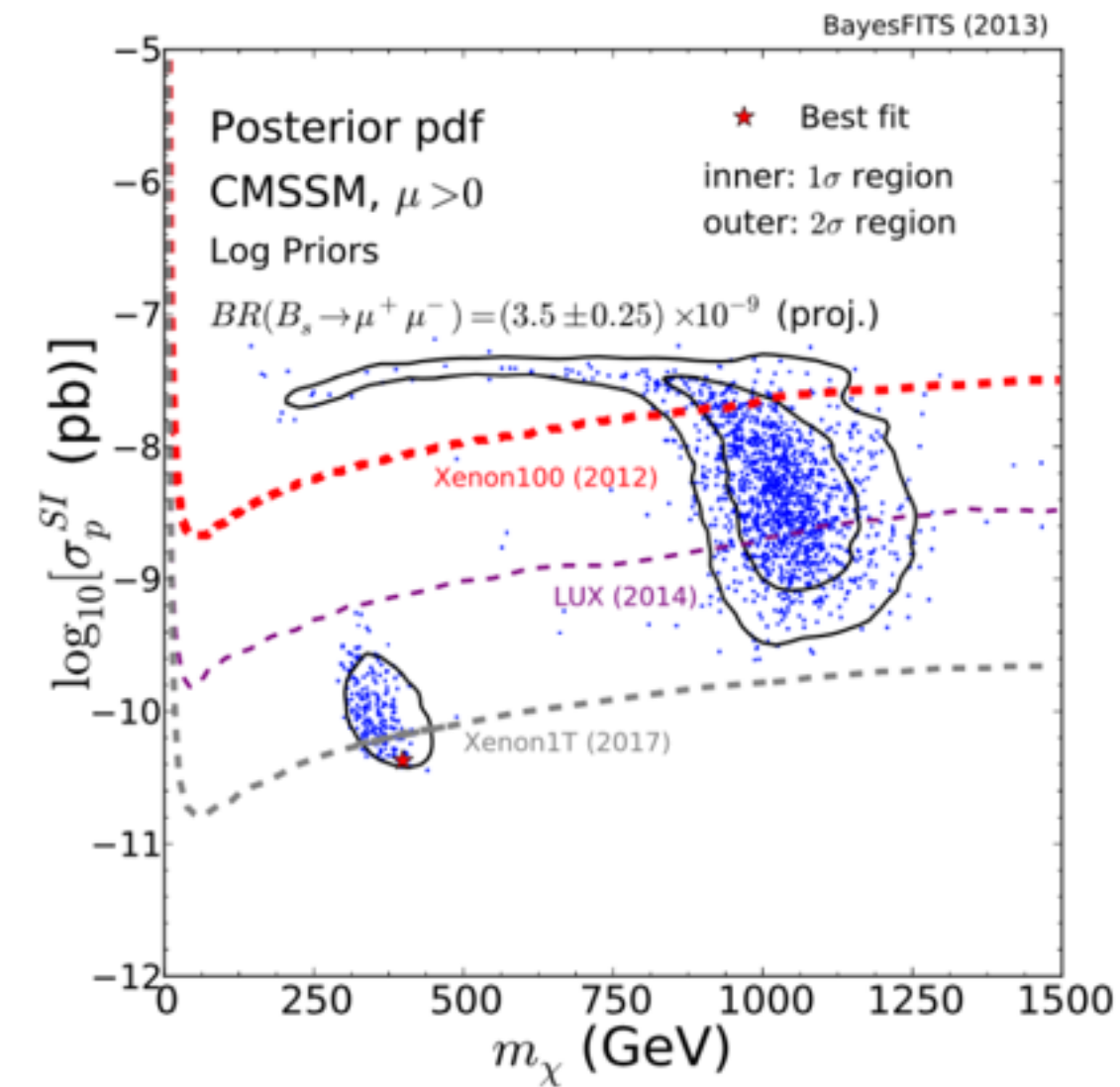
A/H Funnel and B decay

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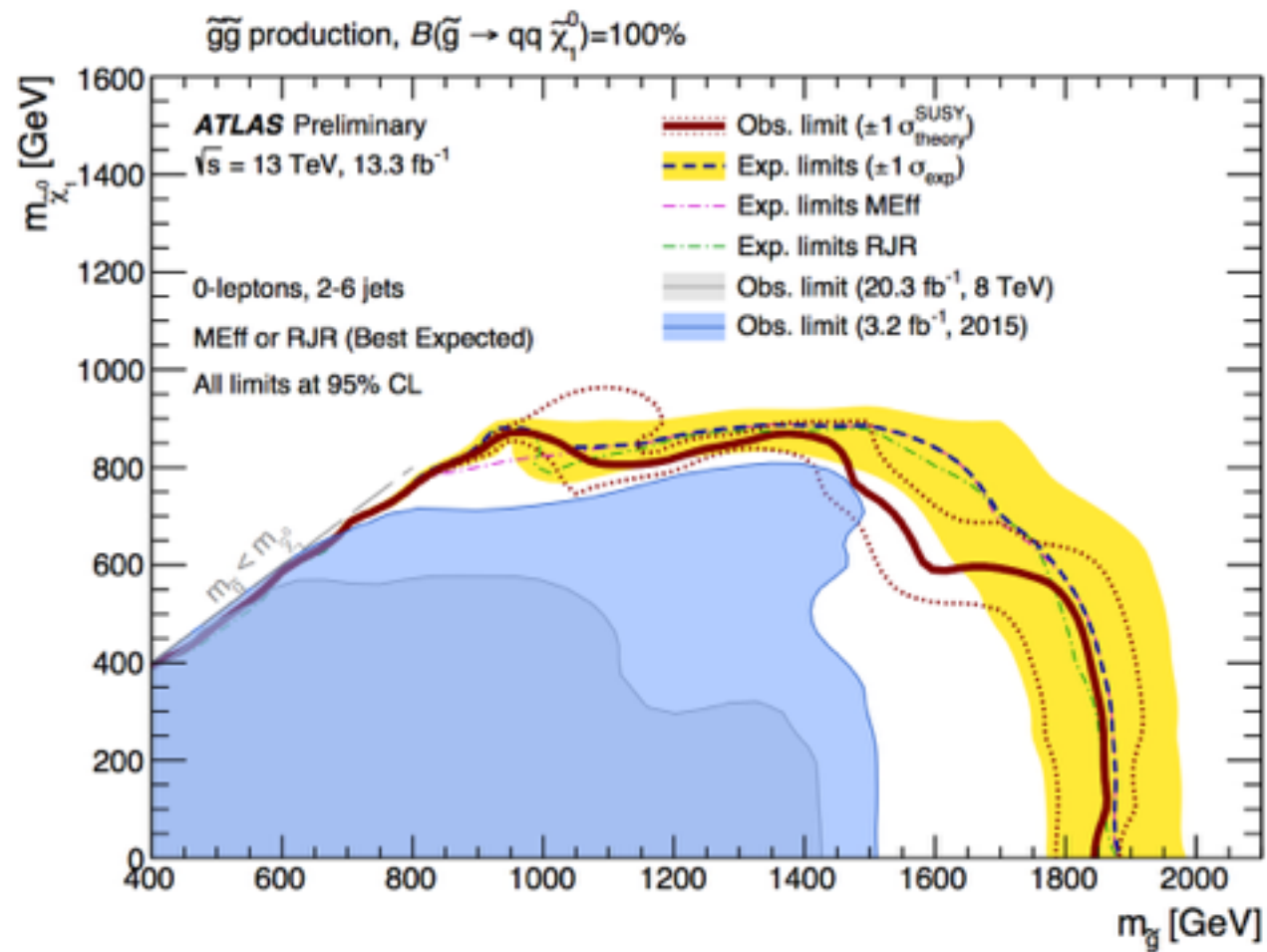


Focus point region and DM direct detection

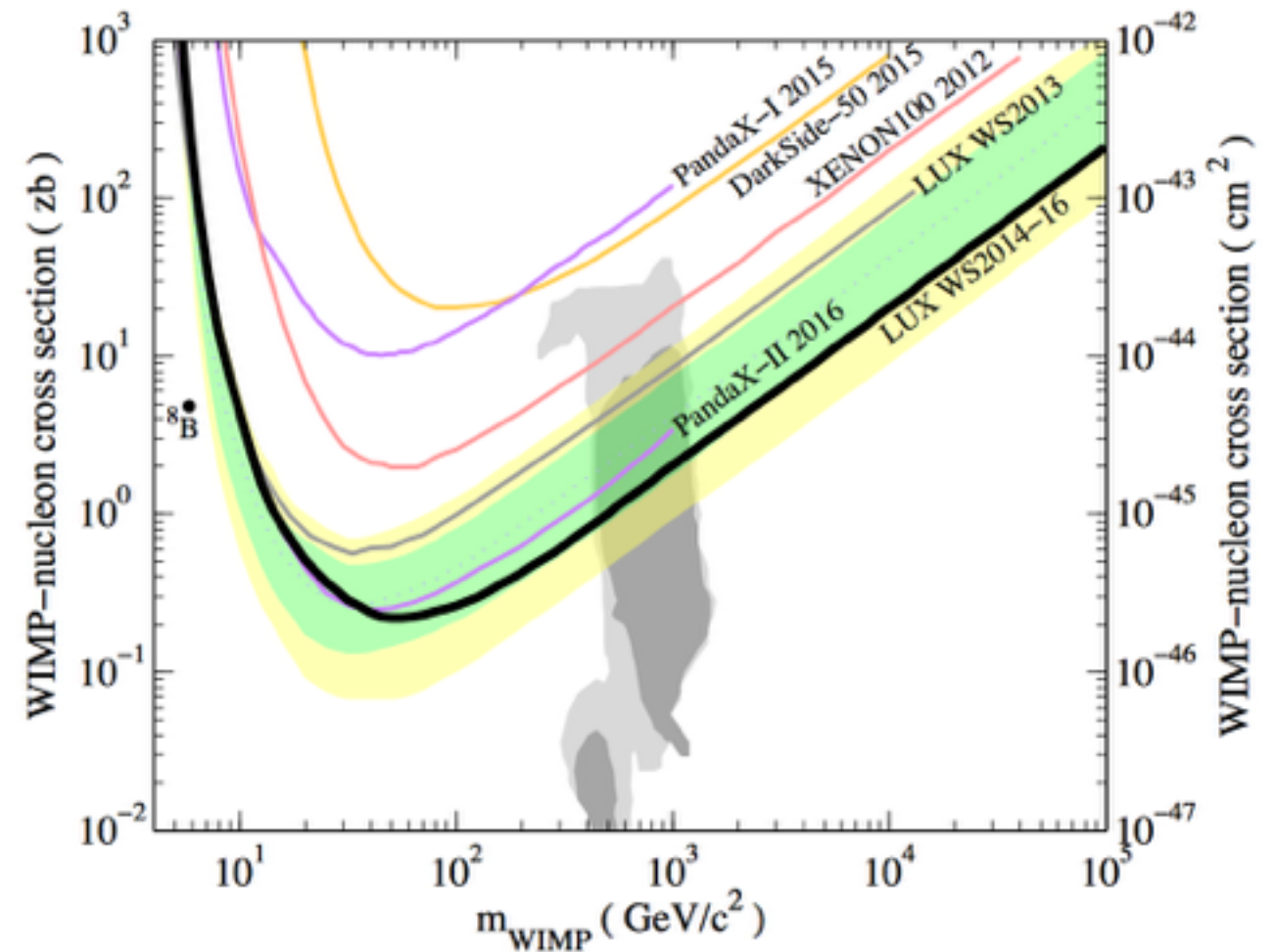
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13 TeV LHC and LUX 2016

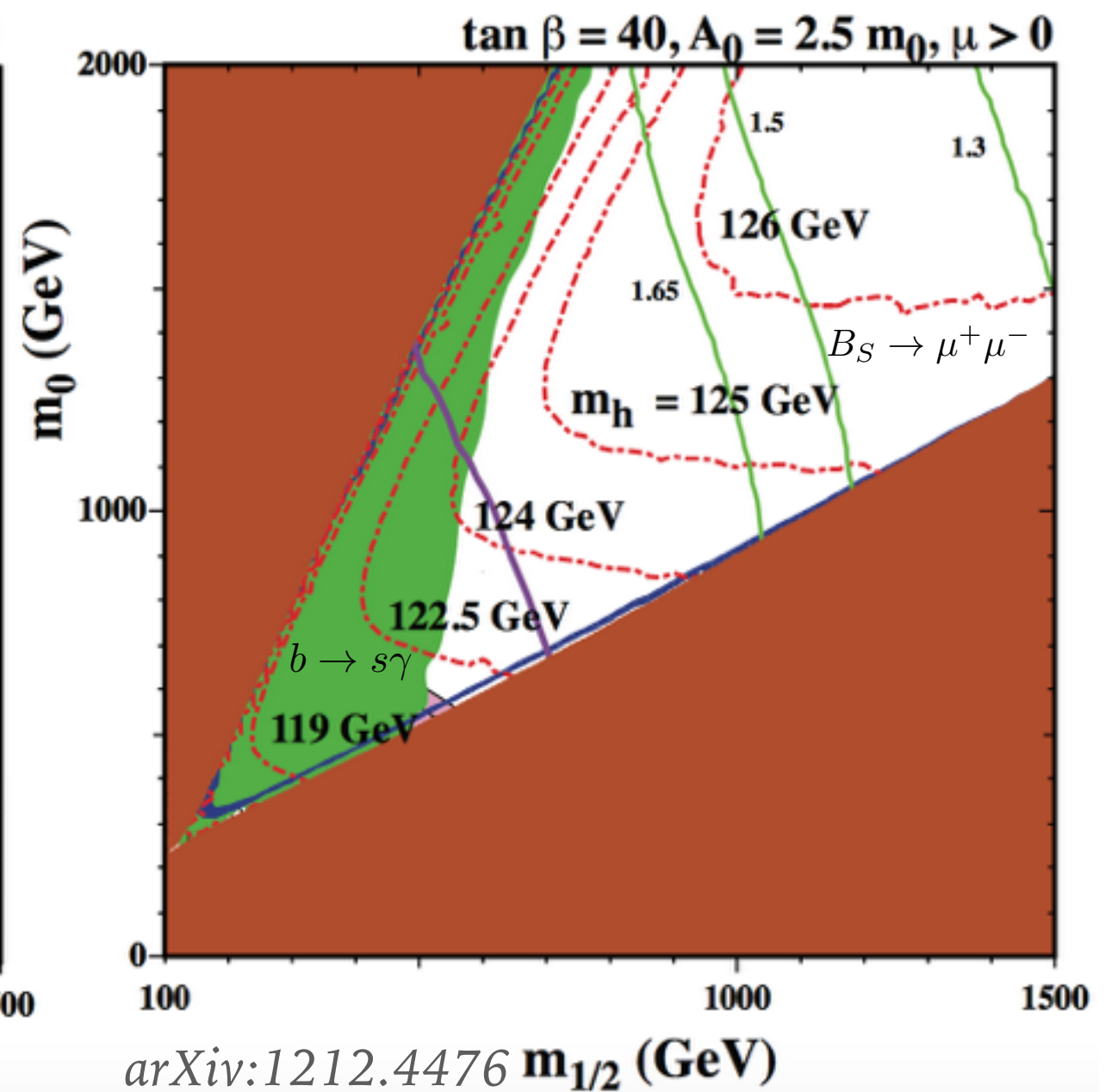
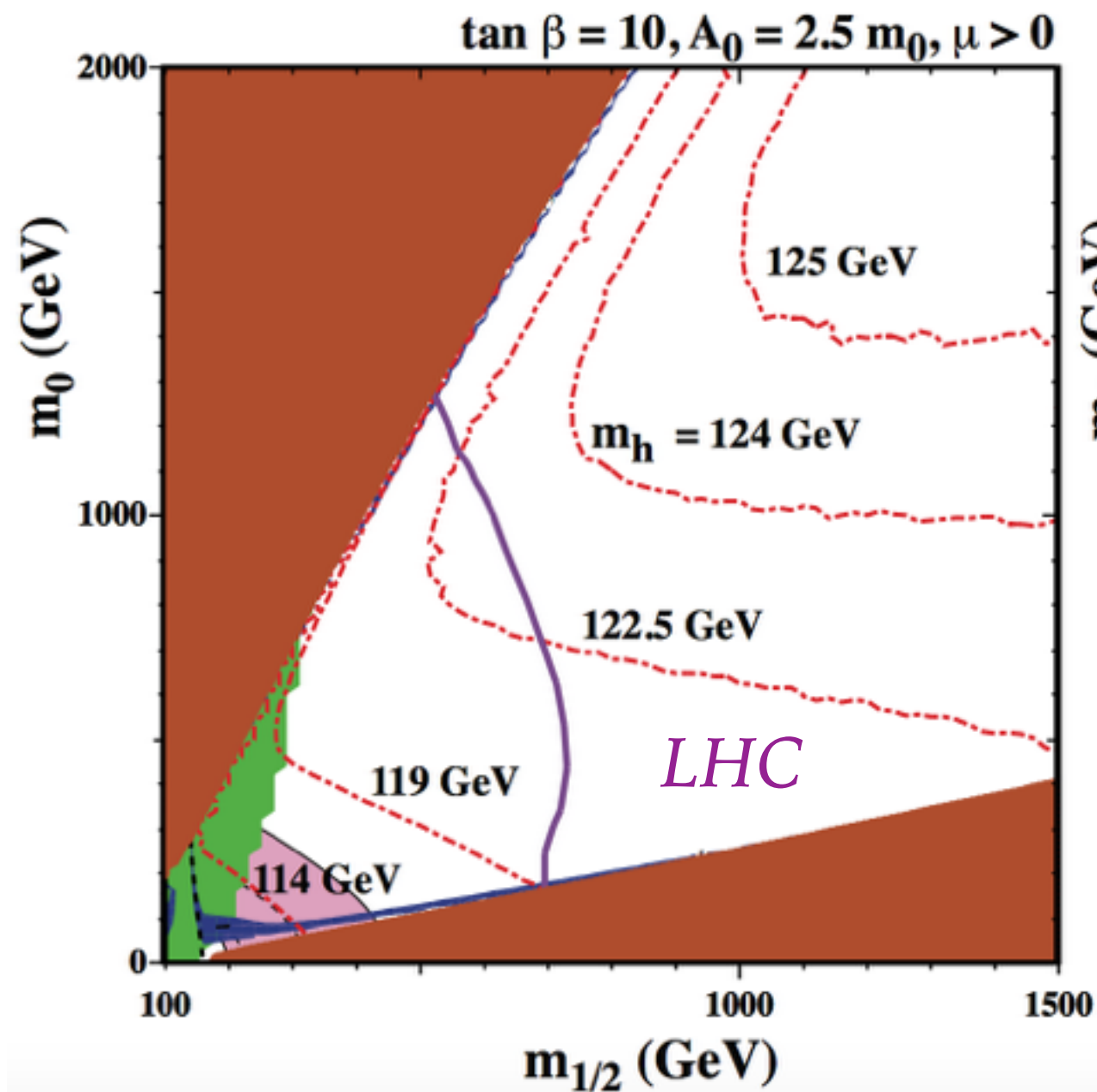


Stau coannihilation



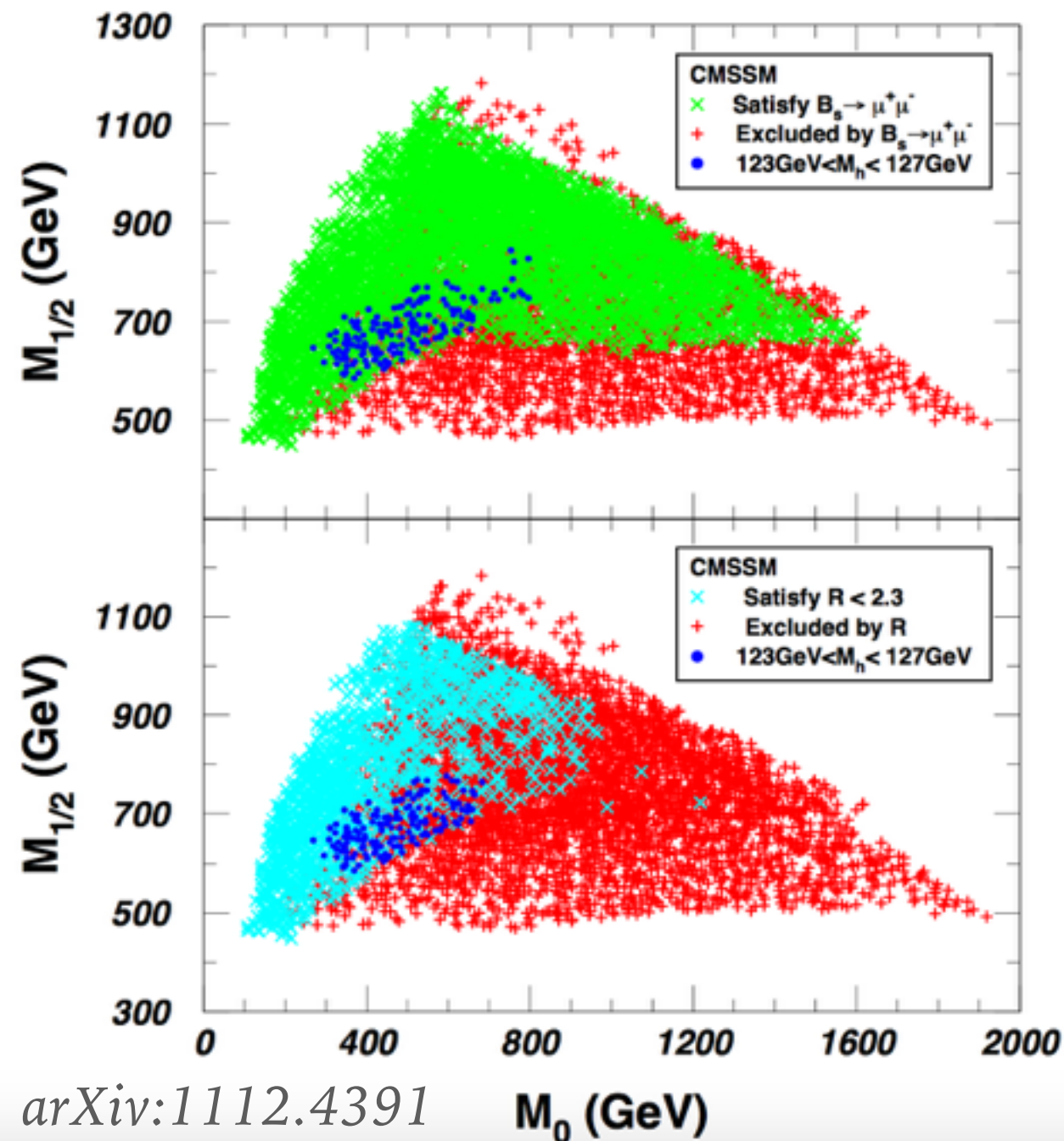
Focus point region

Statistical framework



Statistical framework

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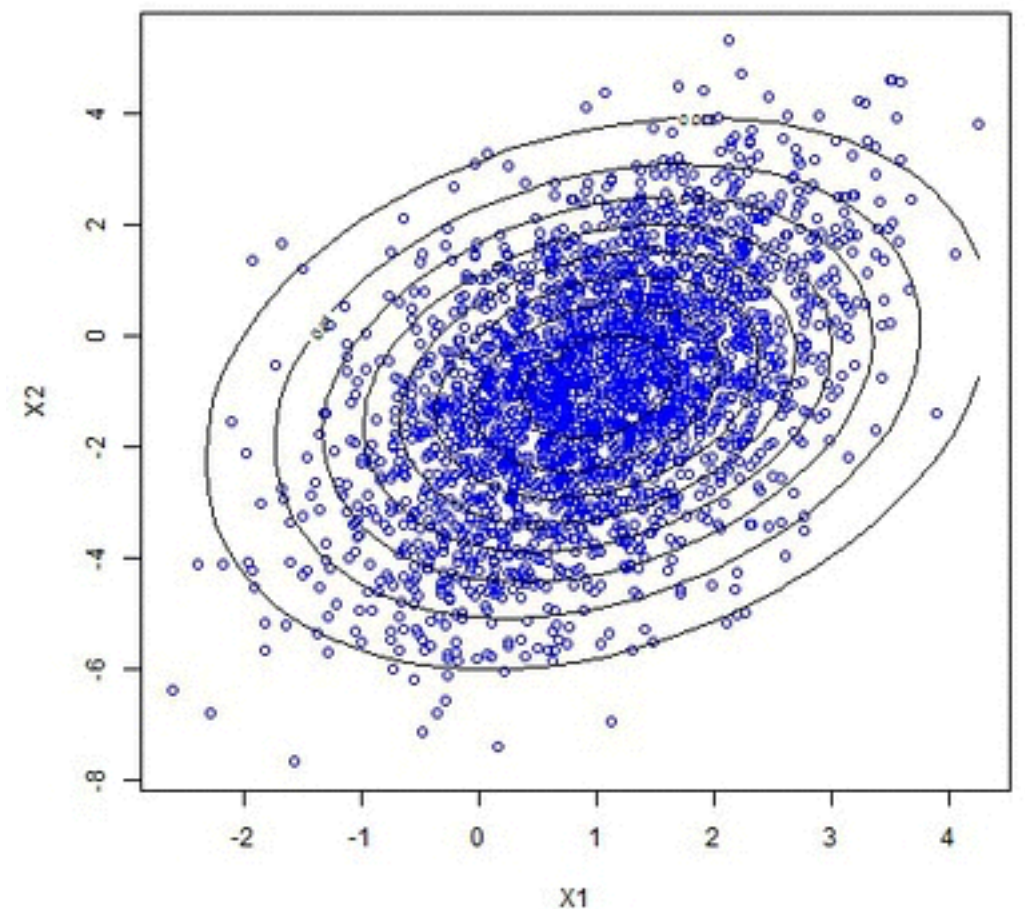
arXiv:1112.4391

M_0 (GeV)

Junjie Cao, Zhaoxia Heng, Dongwei Li, Jin Min Yang

Markov chain Monte
Carlo (MCMC)

Multinest



Likelihood function

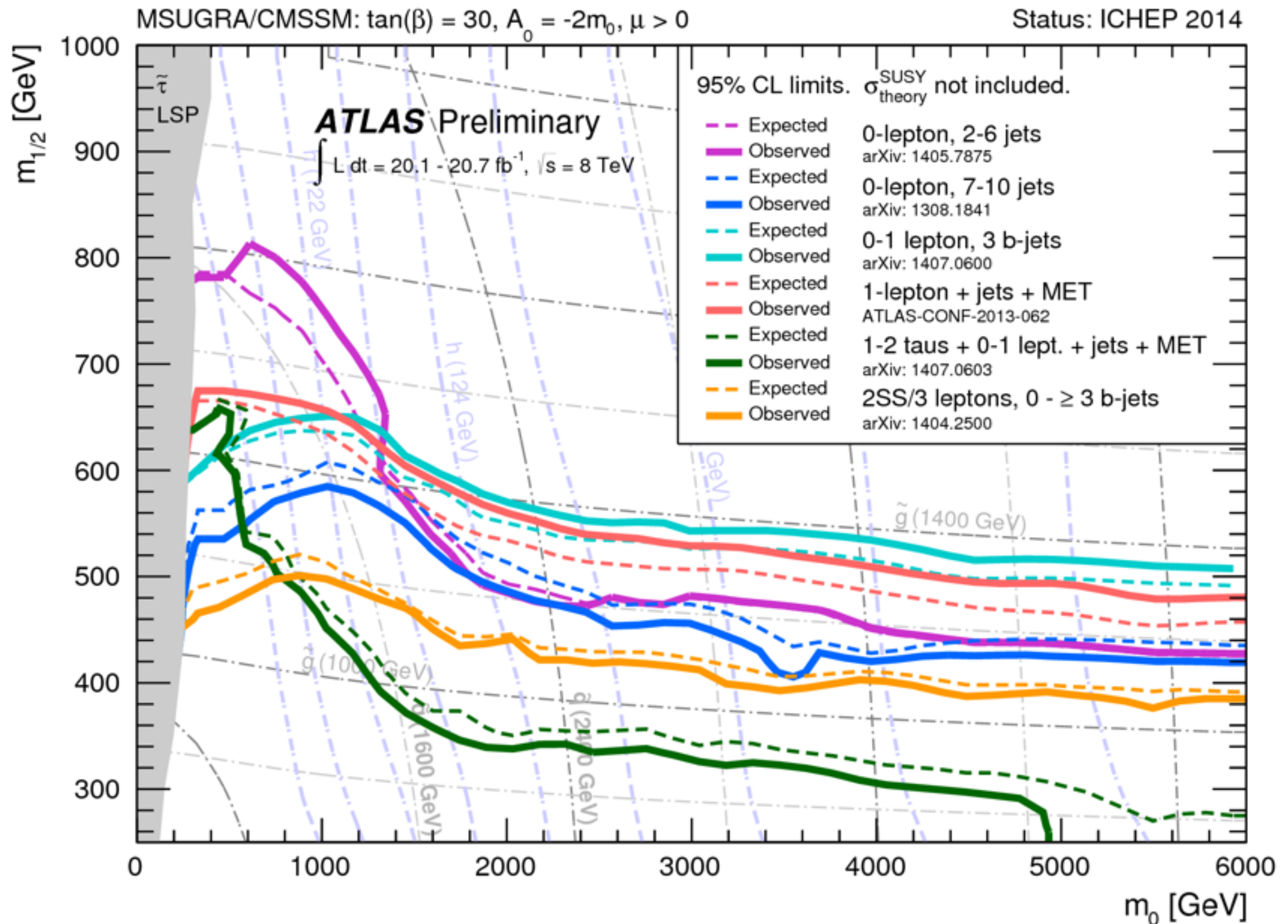
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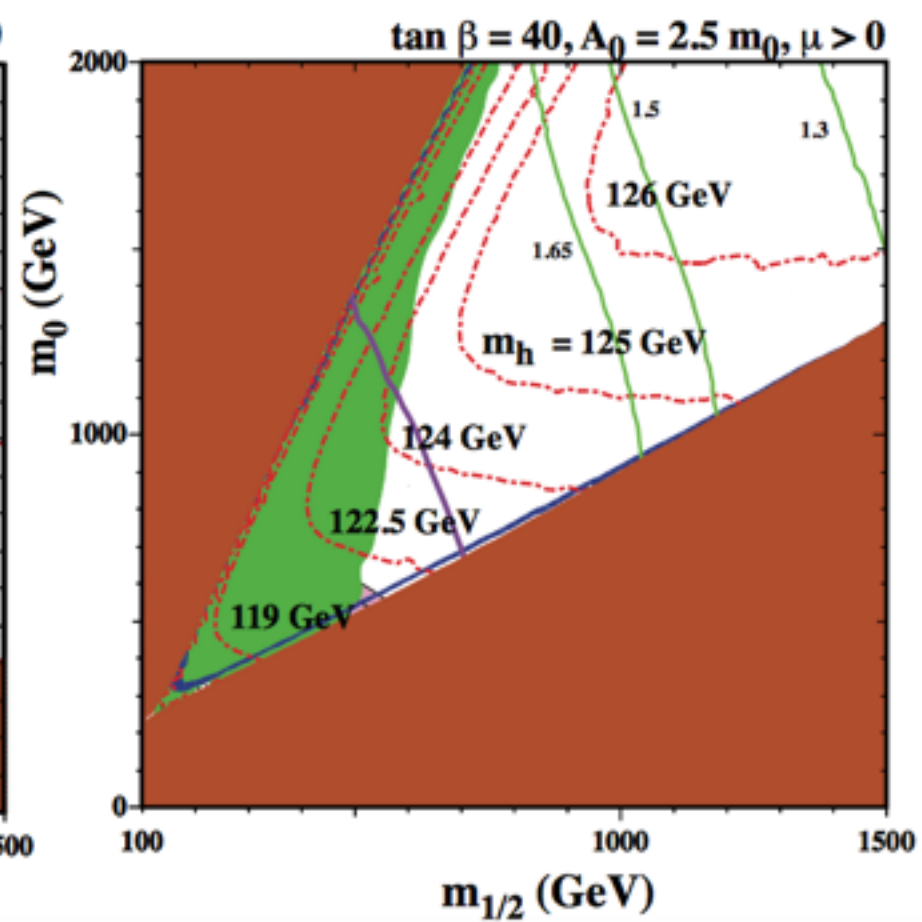
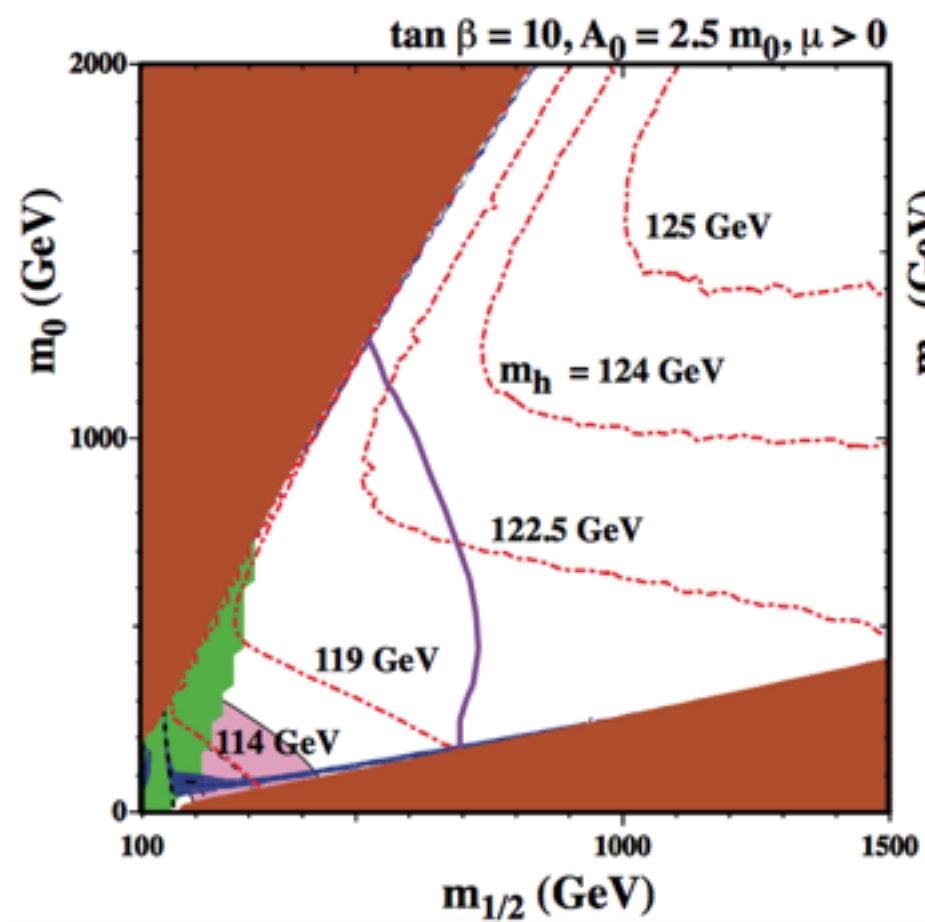
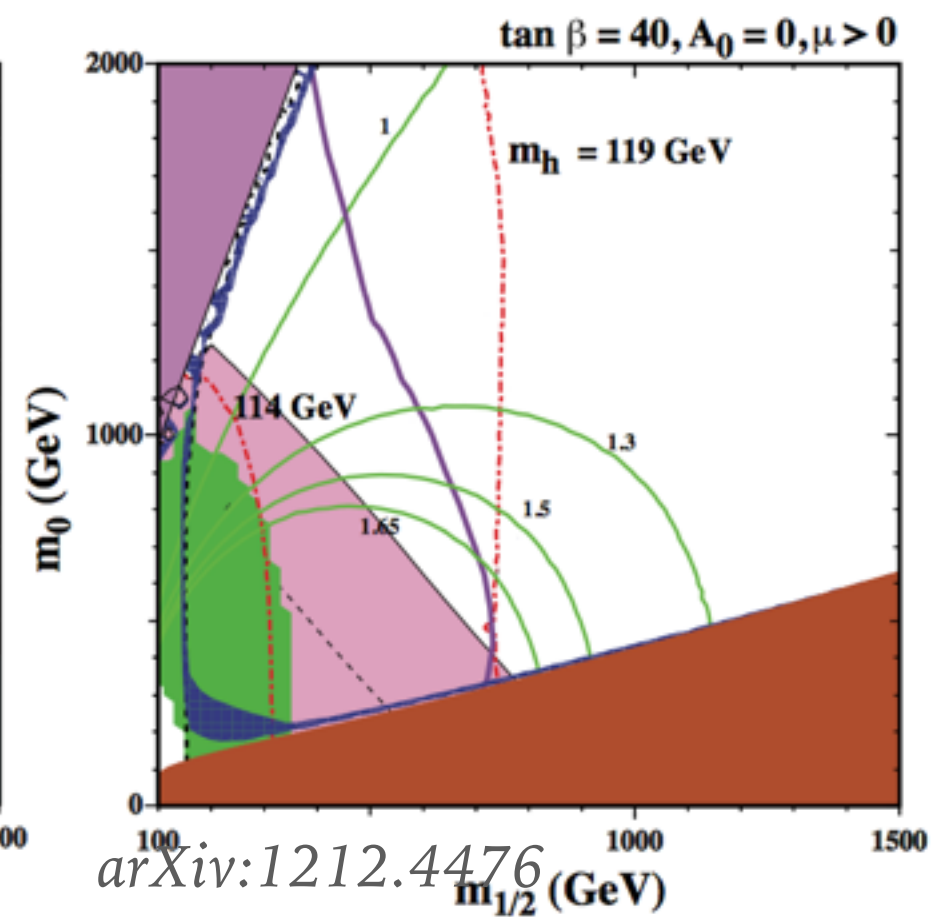
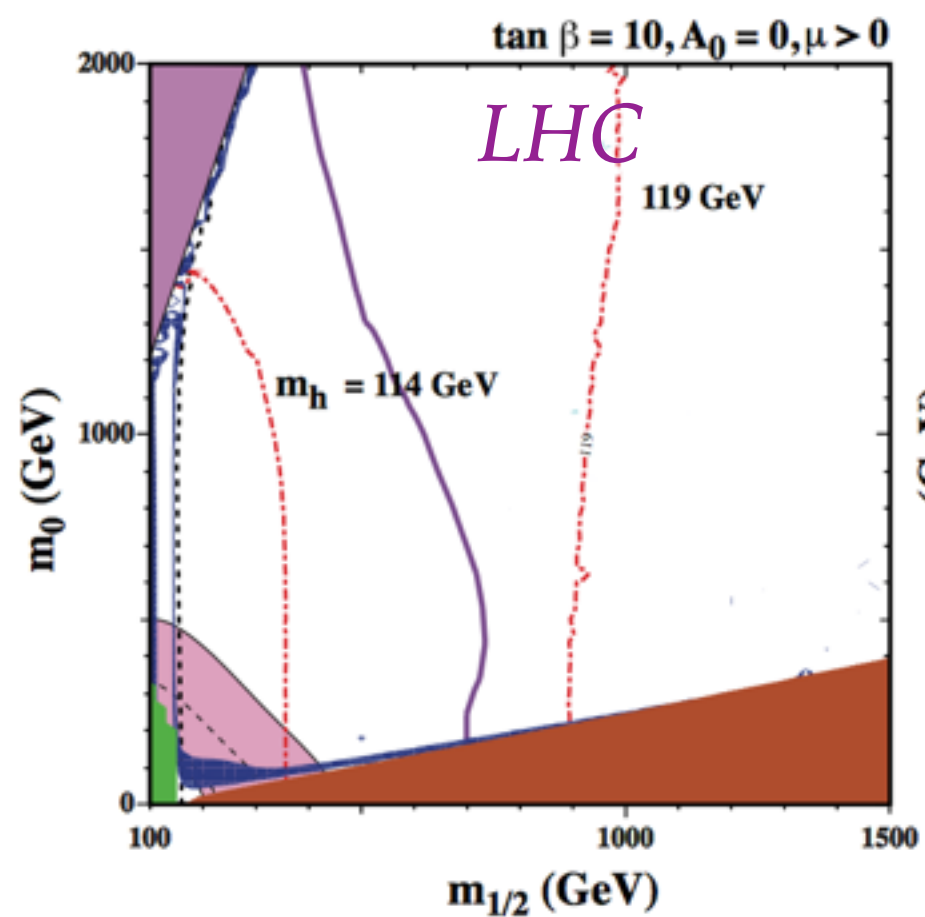
$$p(\Theta|\mathbf{D}) = \frac{p(\mathbf{D}|\Theta)p(\Theta)}{p(\mathbf{D})} \quad p(\mathbf{D}|\Theta) \equiv \mathcal{L}(\Theta)$$

$$\ln\mathcal{L} = \ln\mathcal{L}_{WMAP} + \ln\mathcal{L}_{EW} + \ln\mathcal{L}_B + \ln\mathcal{L}_{g-2} + \ln\mathcal{L}_{m_h} \\ + \ln\mathcal{L}_{LHC} + \ln\mathcal{L}_{LUX} + \ln\mathcal{L}_{LEP} + \ln\mathcal{L}_{HiggsCoupling}$$

Observable	μ	$\sigma(\text{exp})$	$\tau(\text{th})$	References
$\Omega_\chi h^2$	0.1186	0.0031	0.012	arXiv:1303.5076
$\Delta a_\mu \times 10^{10}$	28.7	8	10.0	Eur.Phys.J. C71, 1515 (2011)
$BR(B \rightarrow X_s \gamma) \times 10^4$	3.43	0.22	0.24	arXiv:1412.7515
$BR(B_s^0 \rightarrow \mu^+ \mu^-) \times 10^9$	2.9	0.7	0.29	arXiv:1411.4413
$R(B^- \rightarrow \tau^- \hat{\nu}_\tau)$	1.04	0.24	0.24	arXiv:1412.7515
m_h	125.36	0.41	2.0	arXiv:1406.3827
M_W	80.385	0.015	0.01	Chin. Phys. C 38 (2014) 090001
$\sin^2 \hat{\theta}(M_Z)(\overline{MS})$	0.23153	0.00016	0.00010	arXiv:hep-ex/0509008

Limits from LHC SUSY search





arXiv:1212.4476

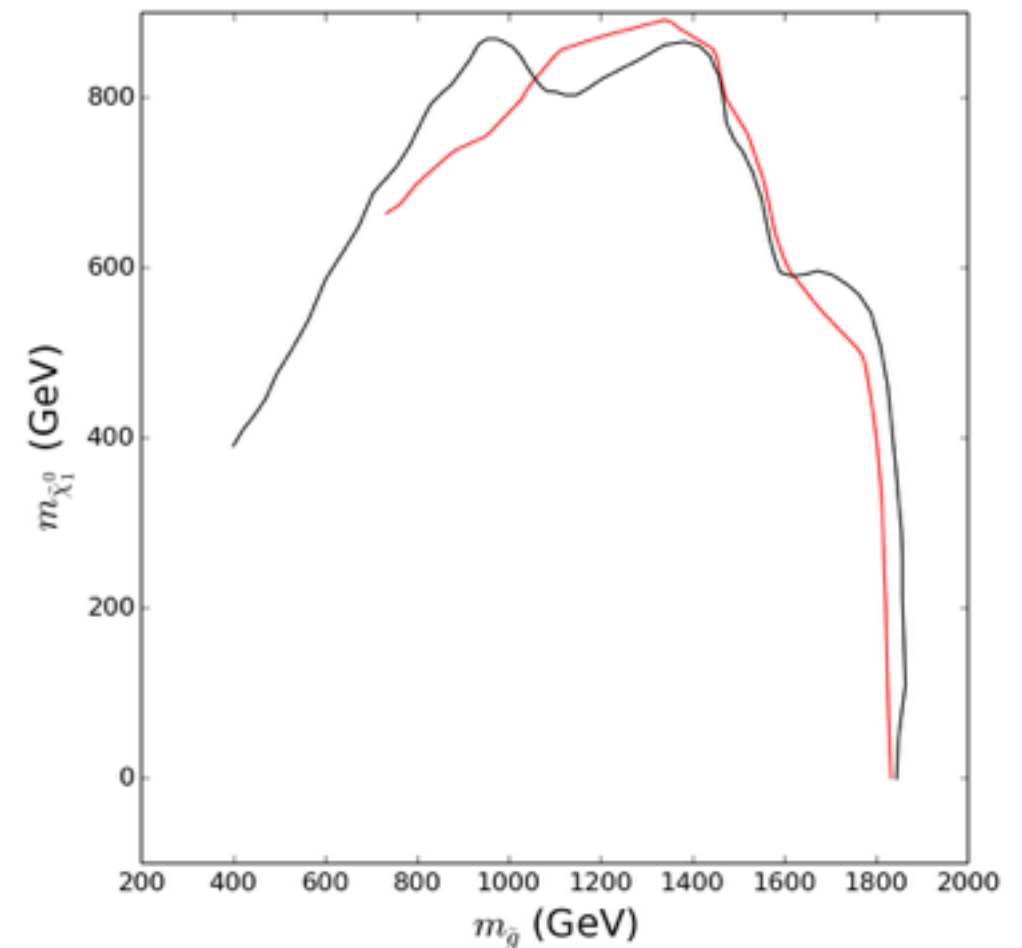
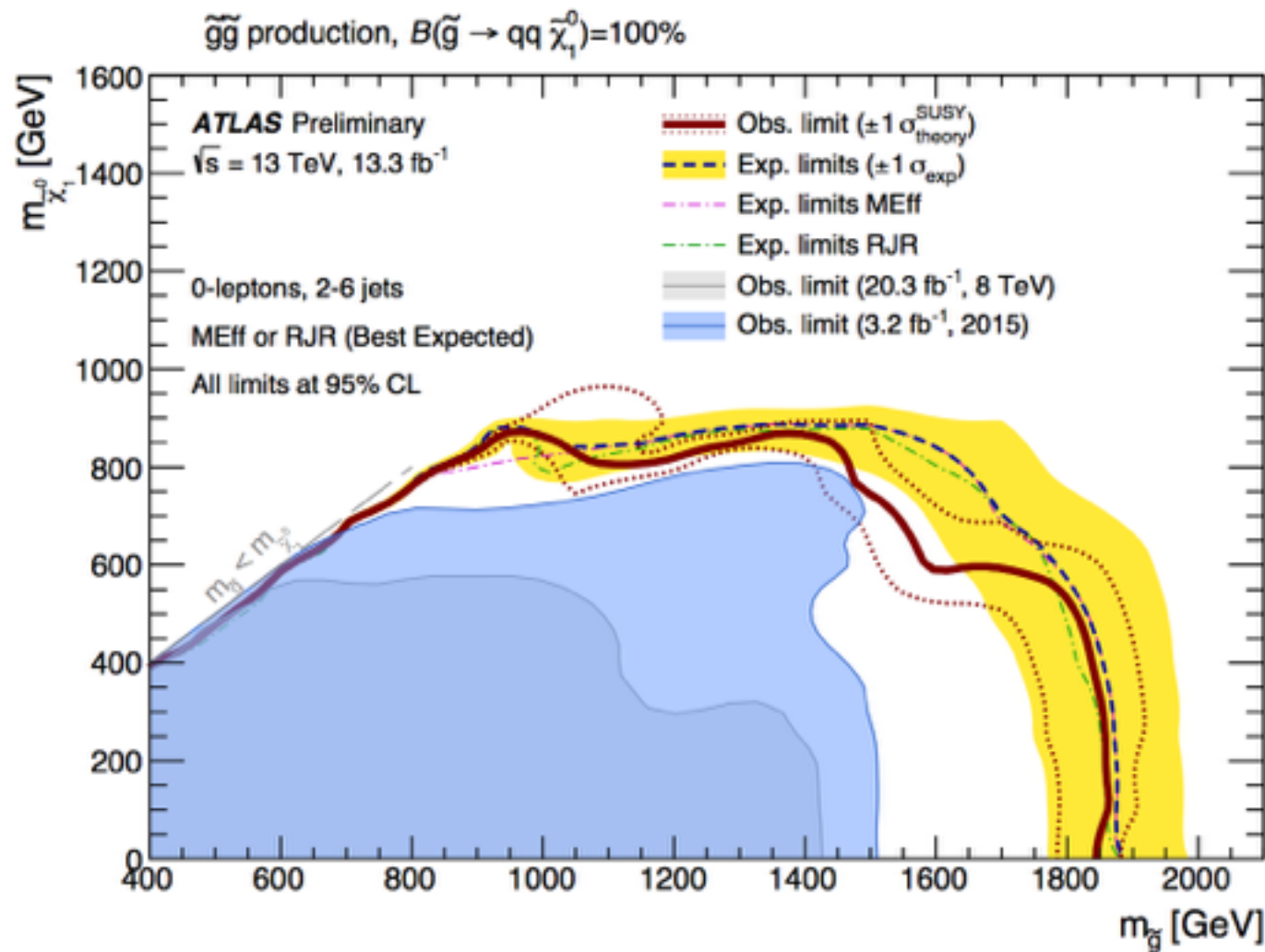
ATLAS 13TeV 13.3 /fb 2~6 jets + MET

The jets+ E^{miss} Meff-based search

Targeted signal	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$				
Requirement	Signal Region				
	Meff-2j-800	Meff-2j-1200	Meff-2j-1600	Meff-2j-2000	Meff-3j-1200
$E_{\text{T}}^{\text{miss}}$ [GeV] >	250				
$p_{\text{T}}(j_1)$ [GeV] >	200	250			600
$p_{\text{T}}(j_2)$ [GeV] >	200	250			50
$p_{\text{T}}(j_3)$ [GeV] >	–				50
$ \eta(j_{1,2}) <$	0.8	1.2			–
$\Delta\phi(\text{jet}_{1,2,3}, E_{\text{T}}^{\text{miss}})_{\min} >$	0.8				0.4
$\Delta\phi(\text{jet}_{1,3}, E_{\text{T}}^{\text{miss}})_{\min} >$	0.4				0.2
$E_{\text{T}}^{\text{miss}} / \sqrt{H_{\text{T}}} \text{ [GeV}^{1/2}] >$	14	16	18	20	16
$m_{\text{eff}}(\text{incl.}) \text{ [GeV]} >$	800	1200	1600	2000	1200

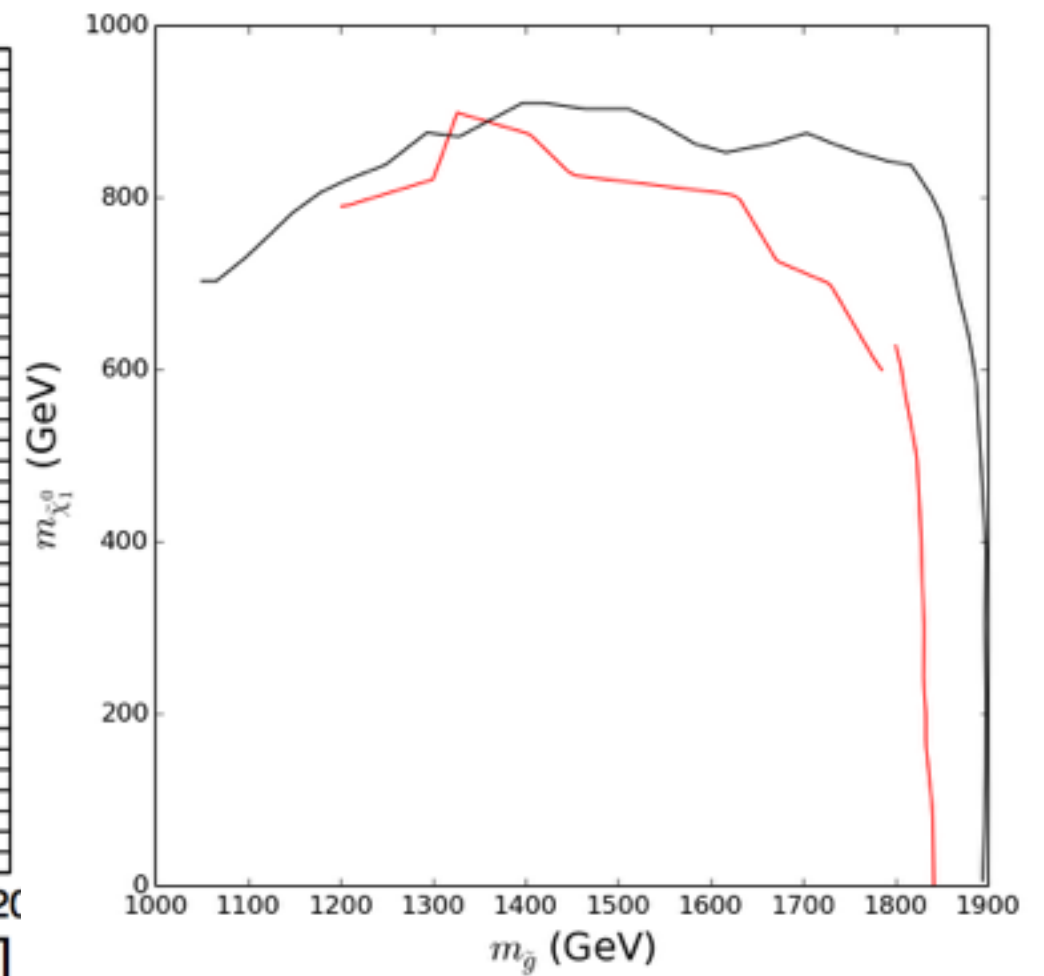
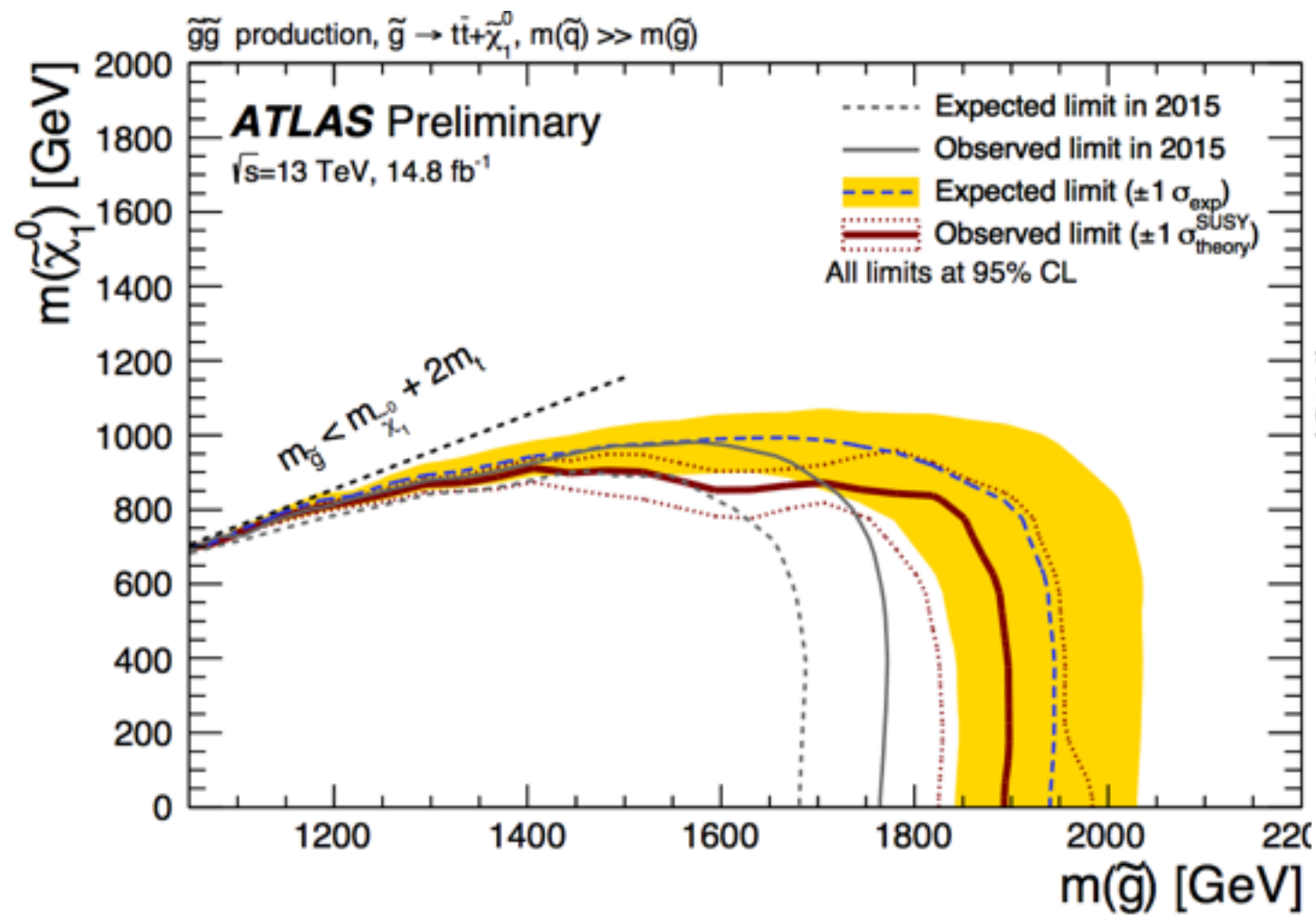
The Recursive Jigsaw Reconstruction technique

Targeted signal	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$					
Requirement	Signal Region					
	RJR-S1		RJR-S2		RJR-S3	
$H_{1,1}^{\text{PP}}/H_{2,1}^{\text{PP}} \geq$	0.6		0.55		0.5	
$H_{1,1}^{\text{PP}}/H_{2,1}^{\text{PP}} \leq$	0.95		0.96		0.98	
$p_{\text{PP},z}^{\text{lab}}/(p_{\text{PP},z}^{\text{lab}} + H_{T,2,1}^{\text{PP}}) \leq$	0.5		0.55		0.6	
$p_{j2,T}^{\text{PP}}/H_{T,2,1}^{\text{PP}} \geq$	0.16		0.15		0.13	
$\Delta_{\text{QCD}} >$	0.001					
	RJR-S1a	RJR-S1b	RJR-S2a	RJR-S2b	RJR-S3a	RJR-S3b
$H_{T,2,1}^{\text{PP}}$ [GeV] >	1000	1200	1400	1600	1800	2000
$H_{1,1}^{\text{PP}}$ [GeV] >	1000		1400		1600	



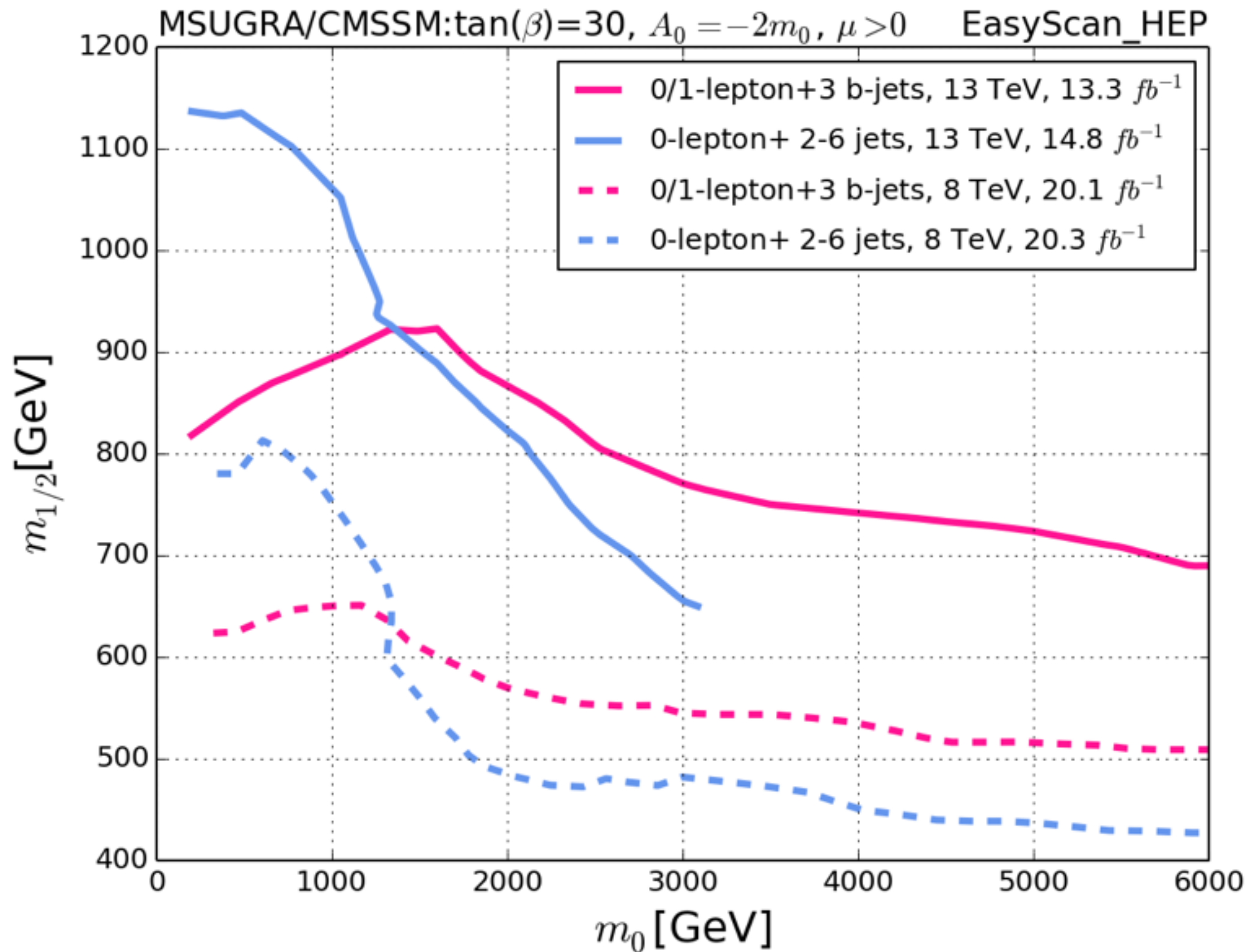
ATLAS 13TeV 14.8 /fb 0/1 lepton + 3 bjets + MET

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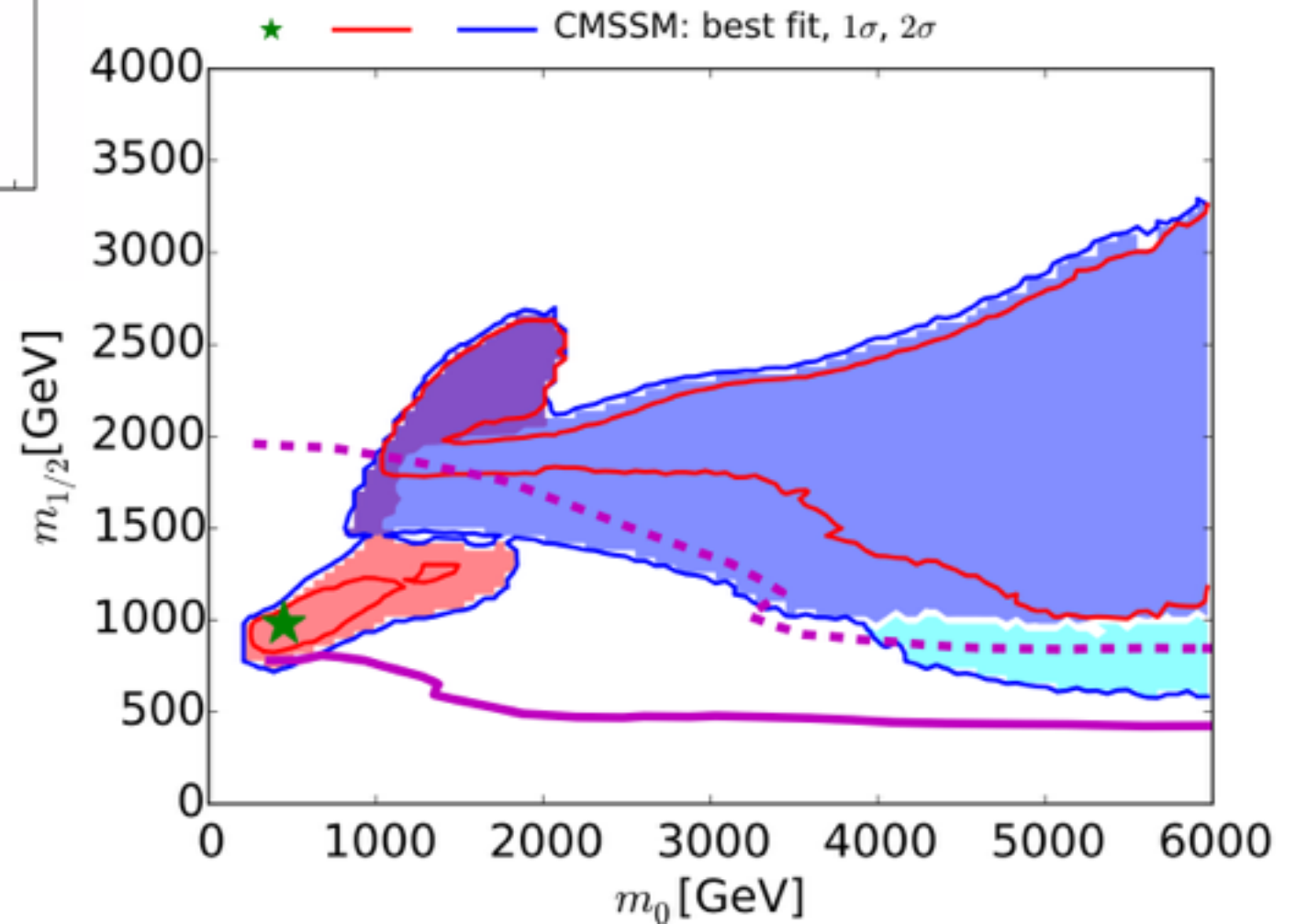
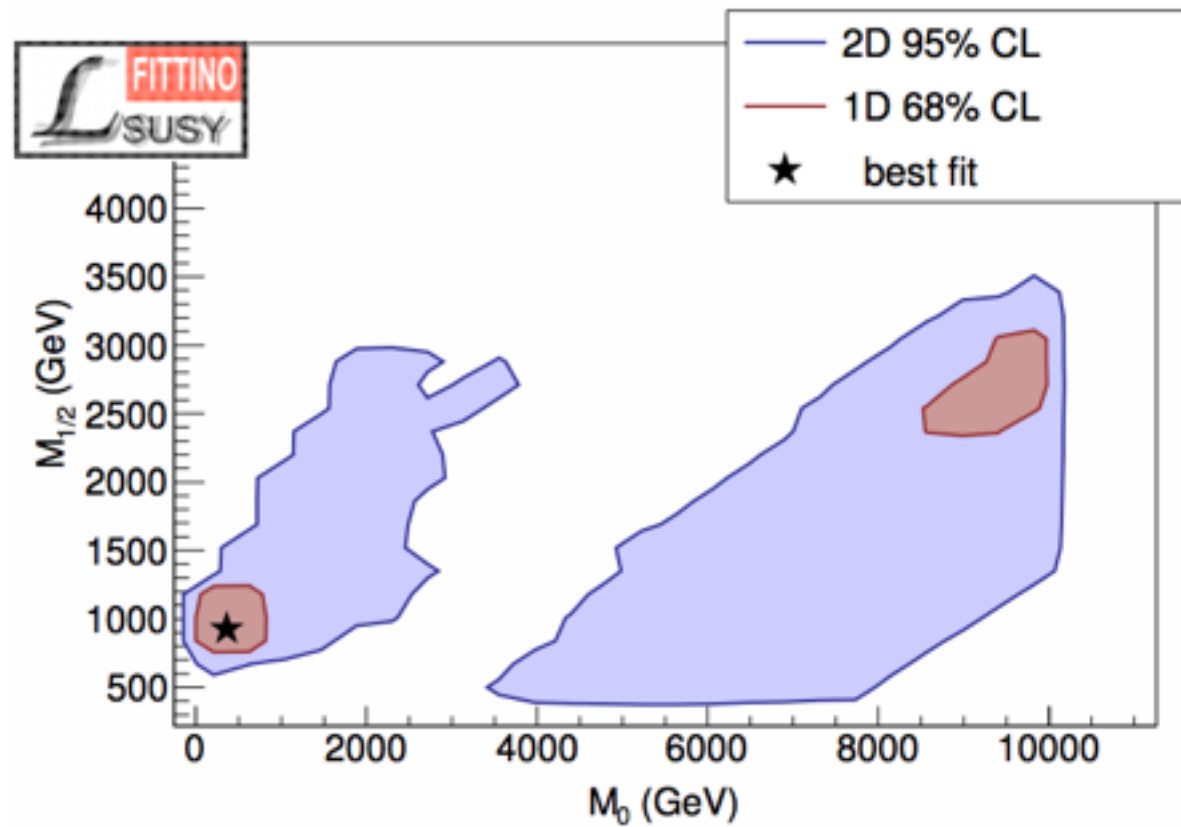
Limits from LHC SUSY search

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Estimate the impact of 13 TeV LHC

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Impact of 13 TeV LHC and LUX WS2014-16

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Input parameter	Prior fuction	Range
M_0 (GeV)	Flat, Log	100, 10000
$M_{1/2}$ (GeV)	Flat, Log	100, 4000
A_0 (GeV)	Flat	-10000, 10000
$\tan\beta$	FLAS	2.0, 65.0
$\text{sign}(\mu)$	Fixed	+1
m_t	Gaussian	172.9 ± 0.91

CMSSM	ATLAS 7 TeV	32.6/23	8.8%	340	910	2670	12
$\mu > 0$	ATLAS _{20/fb} (low)	35.8/23	4.3%	670	1040	3440	21
	ATLAS _{20/fb} (high)	35.1/23	5.1%	5650	2100	-780	51
CMSSM	ATLAS _{20/fb} (low)	38.9/23	2.0%	330	970	3070	10
$\mu < 0$	ATLAS _{20/fb} (high)	36.6/23	3.6%	6650	2550	-3150	39

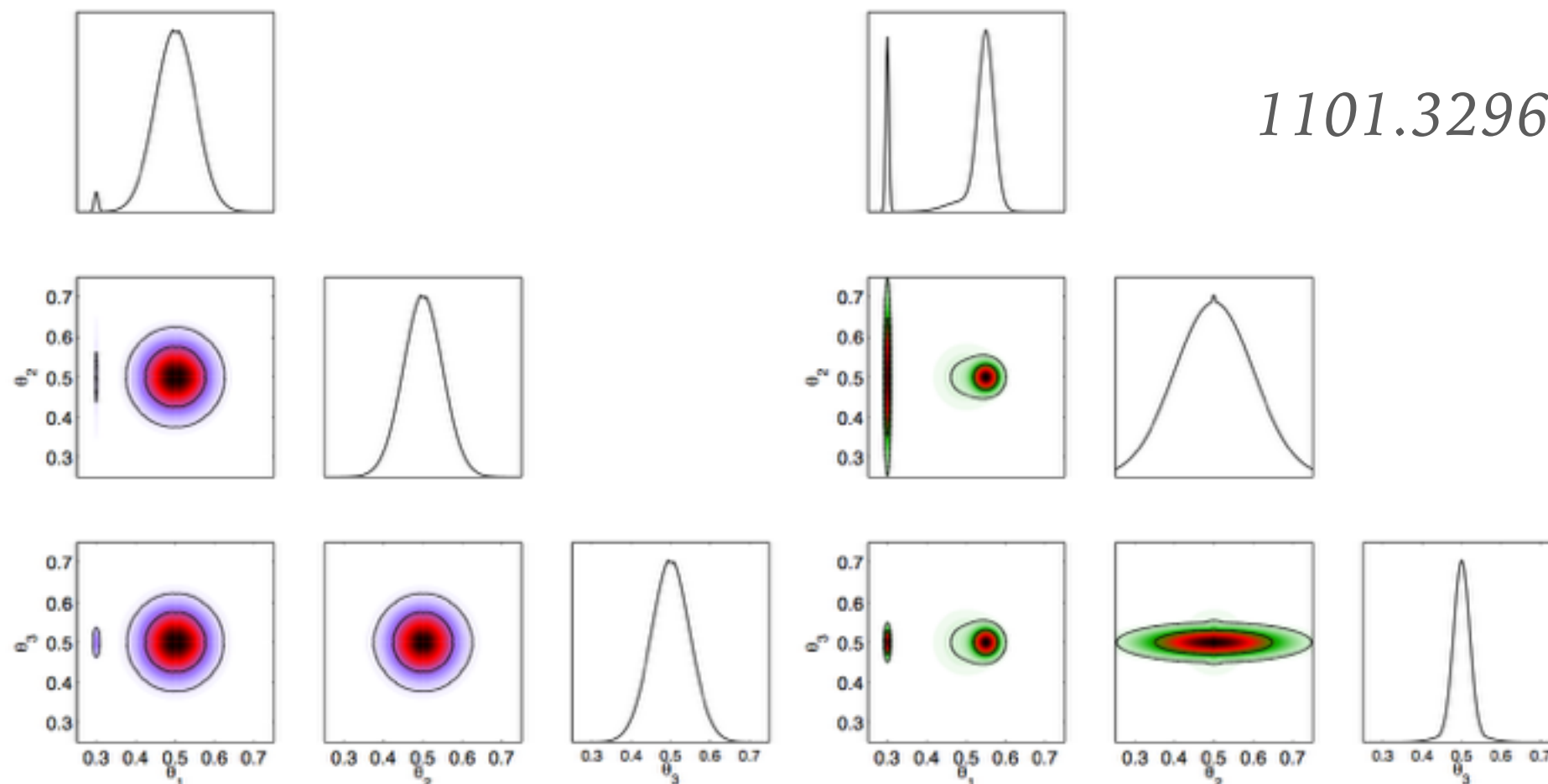
Constraints on a single parameter of interest

The Bayesian marginal pdf:

$$p(\theta_i|\mathbf{D}) = \int p(\boldsymbol{\Theta}|\mathbf{D})d\theta_1\dots d\theta_{i-1}d\theta_{i+1}\dots d\theta_n.$$

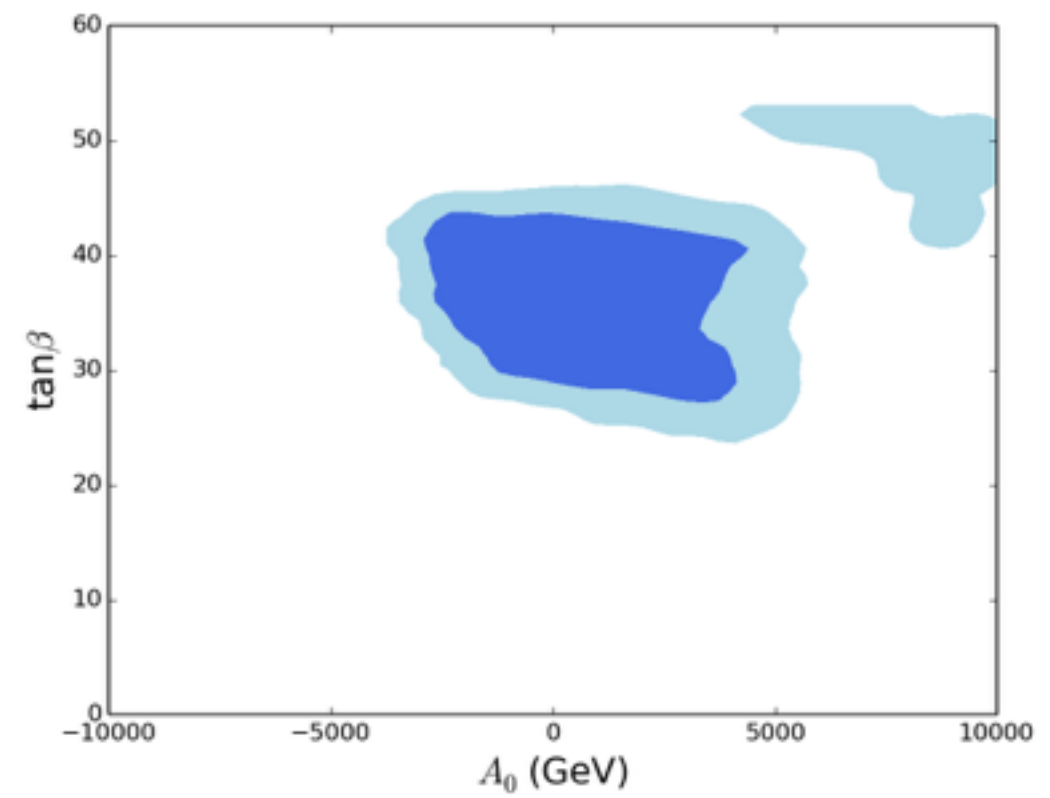
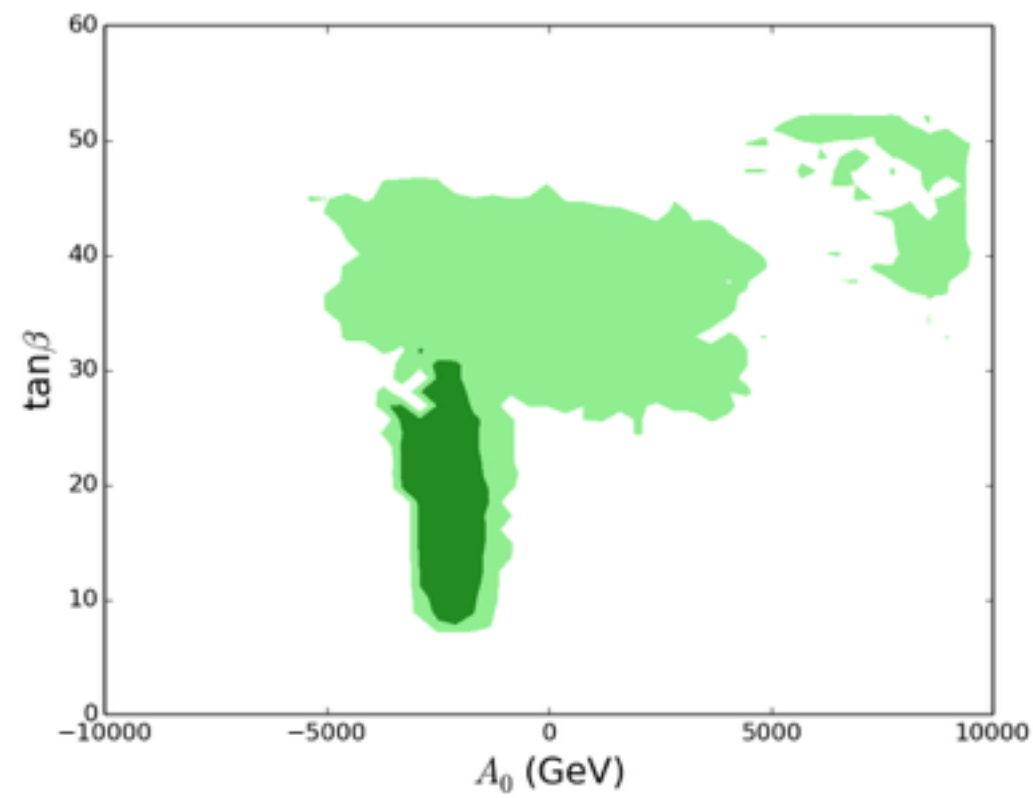
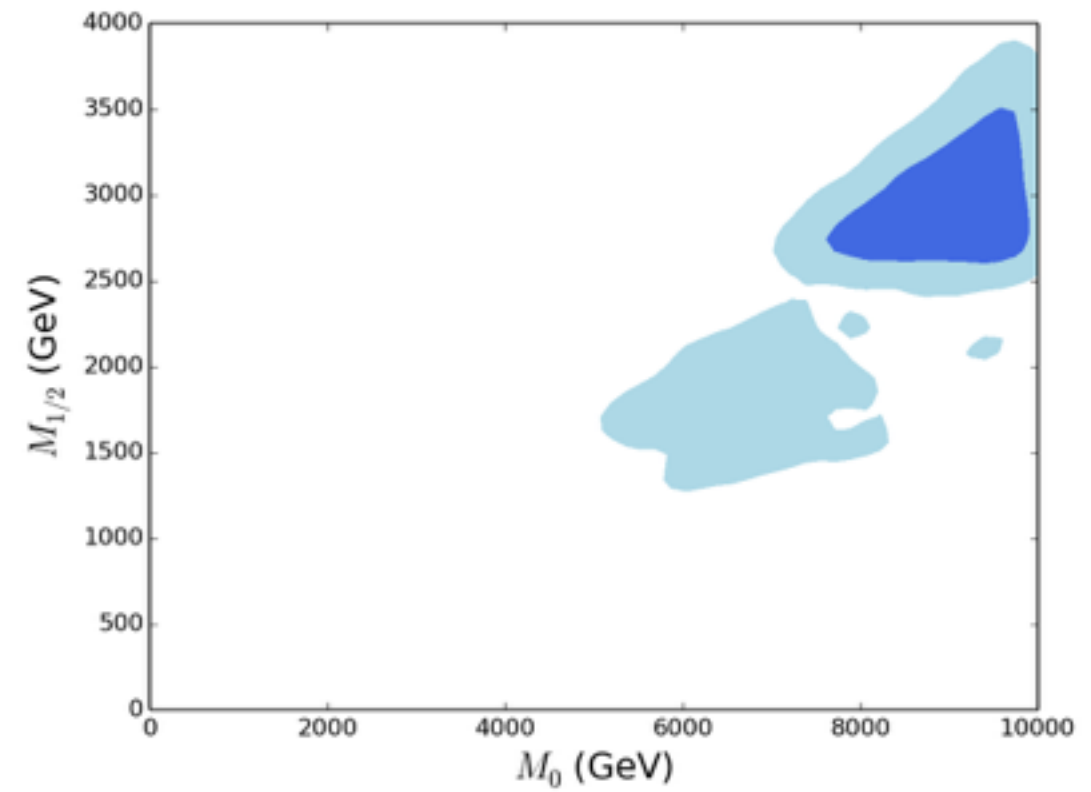
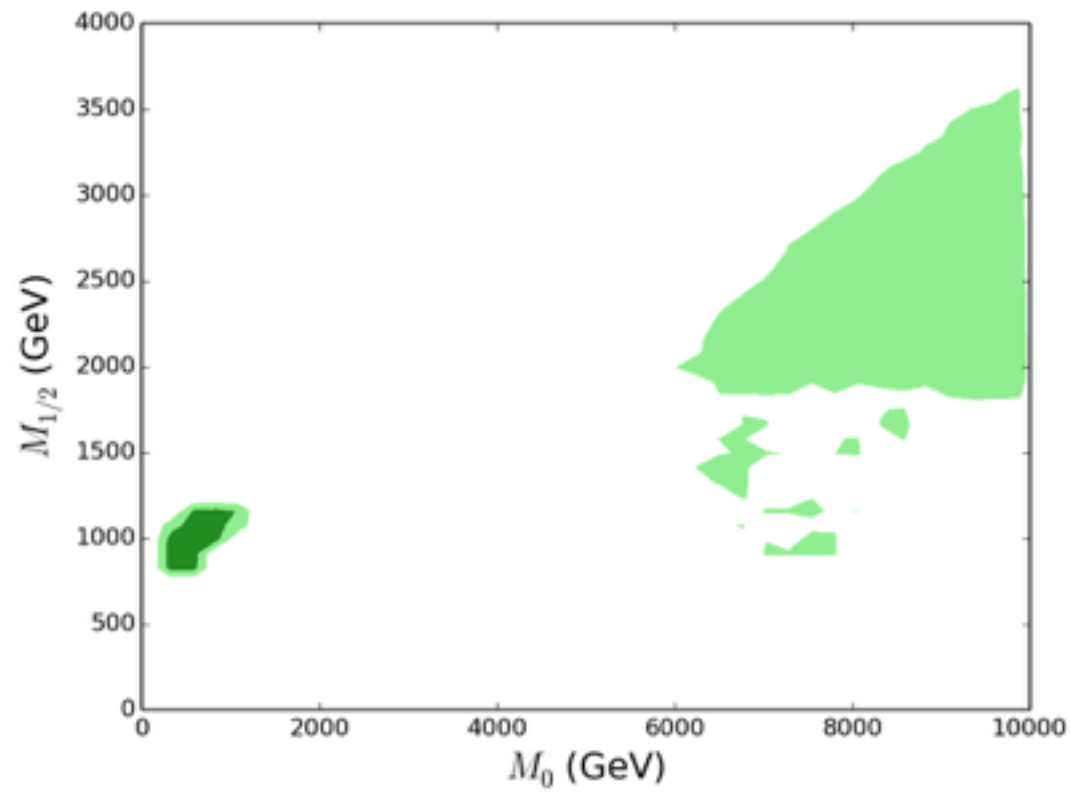
The frequentist profile likelihood function:

$$\mathcal{L}(\theta_i) = \max_{\theta_1,\dots,\theta_{i-1},\theta_{i+1},\dots,\theta_n} \mathcal{L}(\boldsymbol{\Theta}).$$



Impact of 13 TeV LHC and LUX WS2014-16

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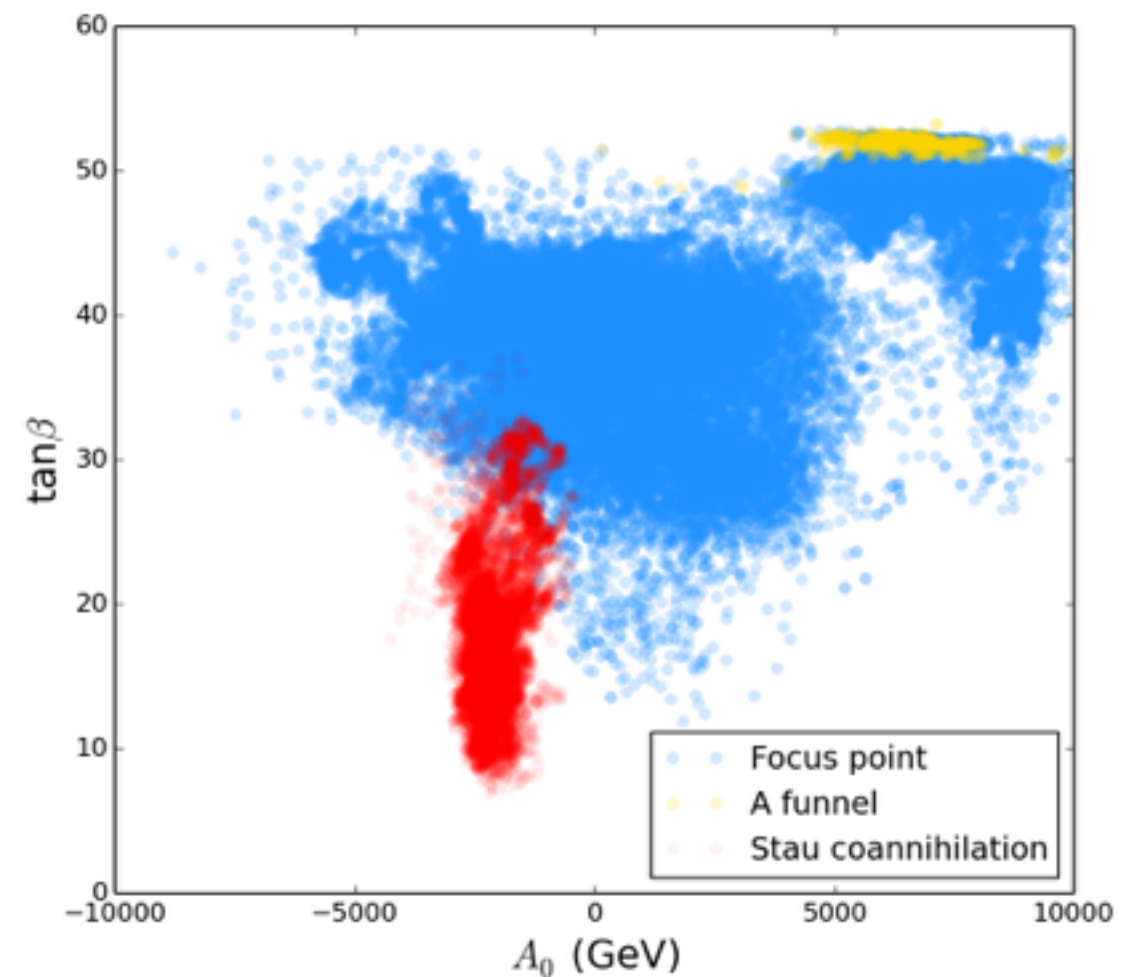
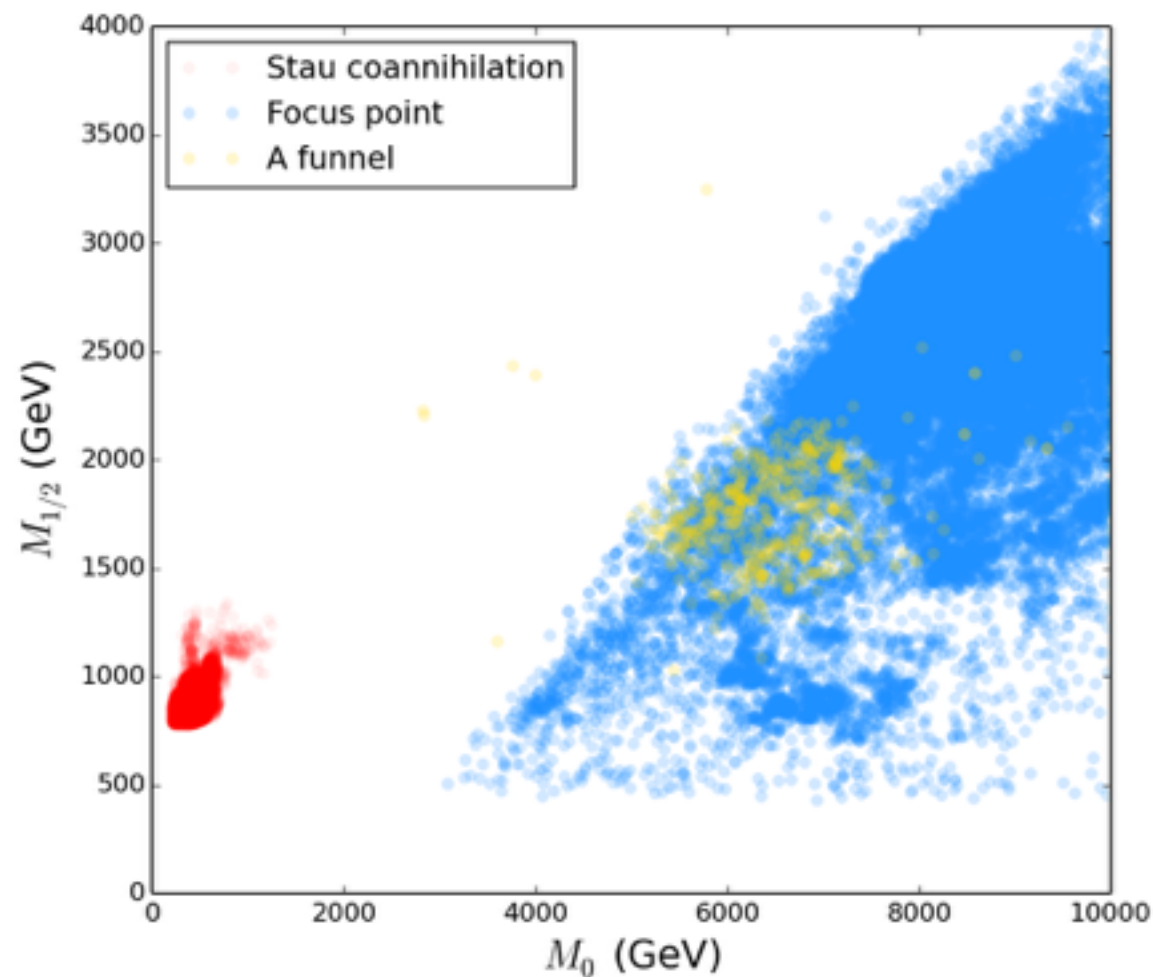
Dark matter annihilation mechanisms

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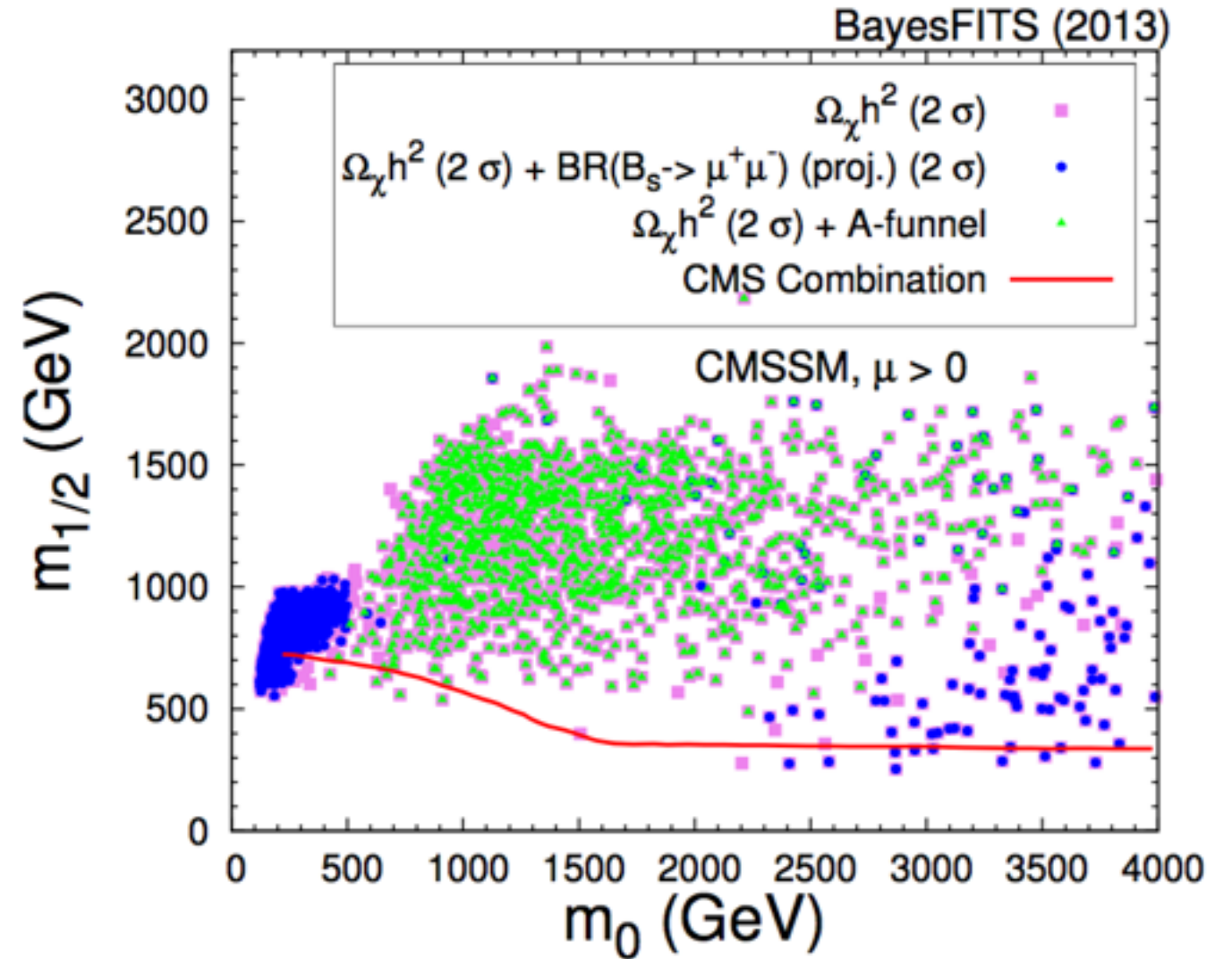
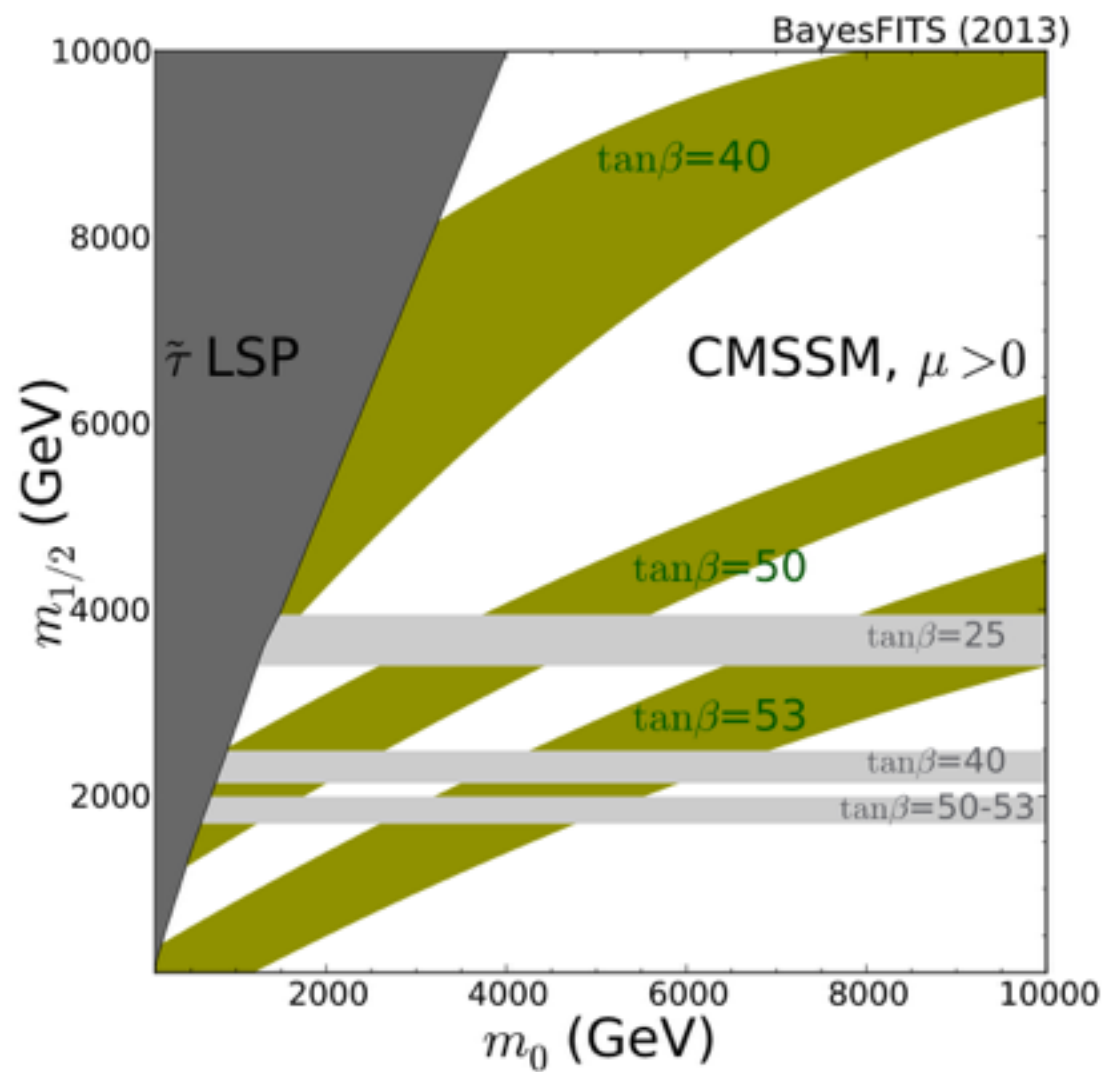
$\tilde{\tau}_1$ coannihilation: $m_{\tilde{\tau}_1}/m_{\tilde{\chi}_1^0} - 1 < 0.15$

A/H funnel: $|m_A/2m_{\tilde{\chi}_1^0} - 1| < 0.2$

Focus Point: $|\mu/m_{\tilde{\chi}_1^0} - 1| < 0.4$



A Funnel region



$BR(B_s \rightarrow \mu^+ \mu^-)_{\text{current}} \times 10^9$	3.2	$+1.5 - 1.2, 10\%$ (0.32)	Gaussian	[5]
$BR(B_s \rightarrow \mu^+ \mu^-)_{\text{proj}} \times 10^9$	3.5 (3.2*)	0.18 (0.16*), 5% [0.18 (0.16*)]	Gaussian	[5]

$$3.5 + \sqrt{0.18^2 + 0.18^2} = 4$$

$BR(B_s^0 \rightarrow \mu^+ \mu^-) \times 10^9$	2.9	0.7	0.29	arXiv:1411.4413
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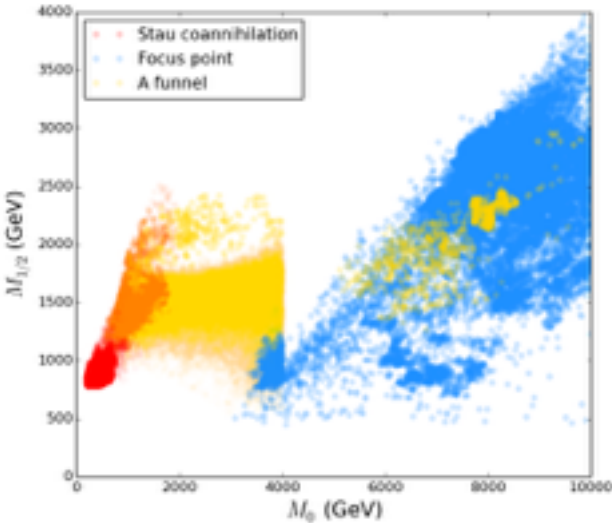
$$2.9 + \sqrt{0.7^2 + 0.29^2} = 4.4$$

Before the conclusion

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Observable	Mean value	Standard deviation		Ref.	1508.05951		
		Experimental	Theoretical				
m_W [GeV]	80.385	0.015	0.01	[57]	$a_\mu - a_\mu^{\text{SM}}$	$(28.7 \pm 8.0) \times 10^{-10}$	[82, 83]
$\sin^2 \theta_{\text{eff}}^{\text{lept}}$	0.231 53	0.00016	0.00010	[58]	$\sin^2 \theta_{\text{eff}}$	0.23113 ± 0.00021	[84]
Γ_Z [GeV]	2.495 2	0.0023	0.001	[59]	m_t	$(173.34 \pm 0.27 \pm 0.71) \text{ GeV}$	[85]
Γ_Z^{inv} [GeV]	0.499	0.0015	0.001	[57]	m_W	$(80.385 \pm 0.015) \text{ GeV}$	[86]
σ_{had}^0 [nb]	41.540	0.037	-	[59]	Δm_s	$(17.719 \pm 0.036 \pm 0.023) \text{ ps}^{-1}$	[87]
R_ℓ^0	20.767	0.025	-	[59]	$\mathcal{B}(B_s \rightarrow \mu\mu)$	$(2.90 \pm 0.70) \times 10^{-9}$	[88]
R_b^0	0.216 29	0.00066	-	[59]	$\mathcal{B}(b \rightarrow s\gamma)$	$(3.43 \pm 0.21 \pm 0.07) \times 10^{-4}$	[89]
R_c^0	0.172 1	0.003	-	[59]	$\mathcal{B}(B \rightarrow \tau\nu)$	$(1.05 \pm 0.25) \times 10^{-4}$	[87]
$BR(B \rightarrow X_s \gamma) \times 10^4$	3.55	0.26	0.30	[57]	1606.03577		
$\frac{BR(B_\mu \rightarrow \tau\nu)}{BR(B_\mu \rightarrow \tau\nu)_{\text{SM}}}$	1.62	0.57	-	[60]			
$BR(B_s^0 \rightarrow \mu^+ \mu^-) \times 10^9$	2.9	1.1	0.38	[61]	1608.00872		
$\Omega_\chi h^2$	0.118 6	0.0031	0.012	[62]			
m_h [GeV]	125.36	0.41	2.0	[63]			

1608.00872



i	Observable $\mu_i(\theta)$	Constraint $D_i^{\text{non-DCS}}$	Likelihood function $L[D_i^{\text{non-DCS}} \mu_i(\theta)]$	Comment
1	$B(b \rightarrow s\gamma)$ [45]	$(3.43 \pm 0.21^{\text{stat}} \pm 0.24^{\text{th}} \pm 0.07^{\text{sys}}) \times 10^{-4}$	Gaussian	reweight
2	$\mathcal{B}(B_s \rightarrow \mu\mu)$ [46]	$(2.9 \pm 0.7 \pm 0.29^{\text{th}}) \times 10^{-9}$	Gaussian	reweight
3	$R(B \rightarrow \tau\nu)$ [45]	1.04 ± 0.34	Gaussian	reweight
4	Δa_μ [47]	$(26.1 \pm 6.3^{\text{exp}} \pm 4.9^{\text{SM}} \pm 10.0^{\text{SUSY}}) \times 10^{-10}$	Gaussian	
5	$\alpha_s(m_Z)$ [48]	0.1184 ± 0.0007	Gaussian	
6	m_t [49]	$173.20 \pm 0.87^{\text{stat}} \pm 1.3^{\text{sys}} \text{ GeV}$	Gaussian	reweight
7	$m_b(m_b)$ [48]	$4.19^{+0.18}_{-0.06} \text{ GeV}$	Two-sided Gaussian	
8	m_h	LHC: $m_h^{\text{low}} = 120 \text{ GeV}$, $m_h^{\text{high}} = 130 \text{ GeV}$	1 if $m_h^{\text{low}} \leq m_h \leq m_h^{\text{high}}$ 0 if $m_h < m_h^{\text{low}}$ or $m_h > m_h^{\text{high}}$	reweight
9	μ_h	CMS and ATLAS in LHC Run 1, Tevatron	LILITH 1.01 [50, 51]	post-MCMC
10	sparticle masses	LEP [52] (via MICROMEGAS [53–55])	1 if allowed 0 if excluded	

13 TeV LHC and LUX WS2014-16

CMSSM

Most of the Stau coannihilation are excluded by 13 TeV LHC.

Some of the focus point region are excluded by the new LUX data.

There are some parameter space in A funnel region which will be very hard to be detected.

13 TeV LHC and LUX WS2014-16

CMSSM

TANK YOU!

张阳

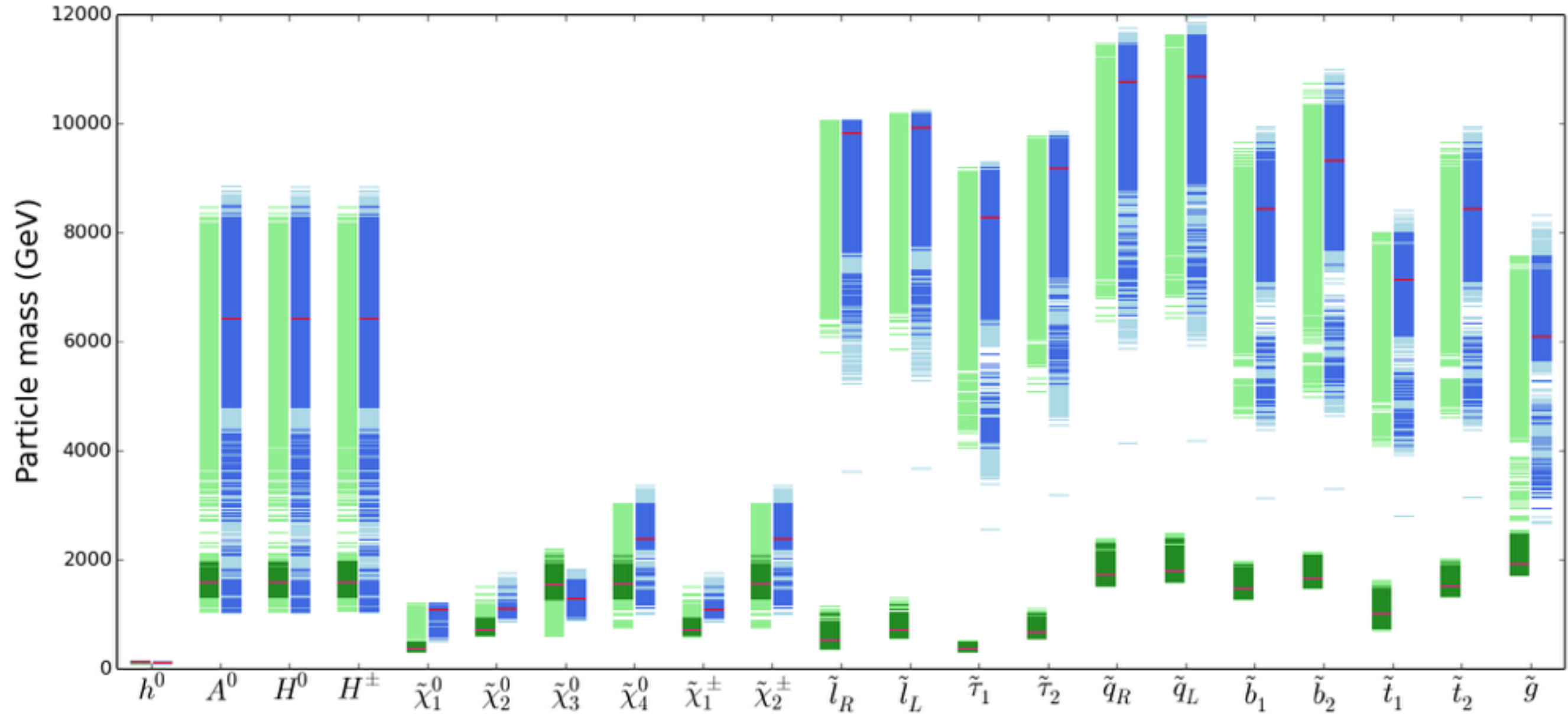
中科院理论物理所

2016.09.23 郑州大学

BACKUP

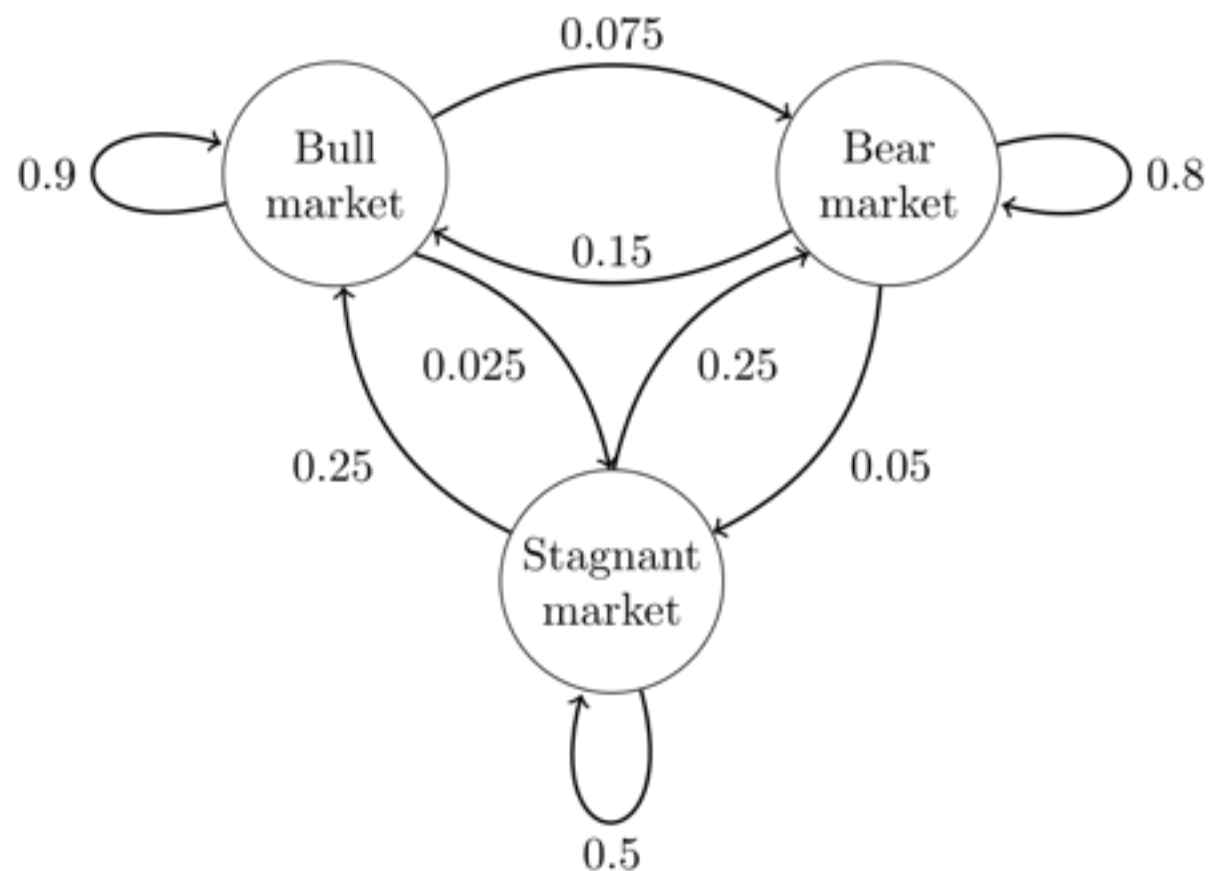
The Higgs and supersymmetric particle mass spectrum

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MARKOV CHAIN

A discrete-time Markov chain is a sequence of random variables $X_1, X_2, X_3, \dots$ with the Markov property, namely that the probability of moving to next state depends only on the present state and not on the previous states.



$$\begin{aligned} \mathbf{x}^{(n+3)} &= [0 \quad 1 \quad 0] \begin{bmatrix} 0.9 & 0.075 & 0.025 \\ 0.15 & 0.8 & 0.05 \\ 0.25 & 0.25 & 0.5 \end{bmatrix}^3 \\ &= [0 \quad 1 \quad 0] \begin{bmatrix} 0.7745 & 0.17875 & 0.04675 \\ 0.3575 & 0.56825 & 0.07425 \\ 0.4675 & 0.37125 & 0.16125 \end{bmatrix} \\ &= [0.3575 \quad 0.56825 \quad 0.07425]. \end{aligned}$$

$$\lim_{N \rightarrow \infty} P^N = \begin{bmatrix} 0.625 & 0.3125 & 0.0625 \\ 0.625 & 0.3125 & 0.0625 \\ 0.625 & 0.3125 & 0.0625 \end{bmatrix}$$

MARKOV CHAIN MONTE CARLO

构造一个转移矩阵为 P 的马氏链, 使得该马氏链的平稳分布恰好是 $p(x)$, 那么我们从任何一个初始状态 x_0 出发沿着马氏链转移, 得到一个转移序列 $x_0, x_1, x_2, \dots, x_n, x_{n+1}, \dots$, 如果马氏链在第 n 步已经收敛了, 于是我们就得到了 $p(x)$ 的样本 $x_n, x_{n+1}, \dots$ 。

METROPOLISHASTINGSSAMPLER (π_0, q, T, x_0)

1 $\mathcal{S} \leftarrow \emptyset.$

2 **for** $t \leftarrow 1$ **to** T ,

3 Sample $x_* \sim q(\cdot | x_{t-1})$ and $u \sim \mathcal{U}_{(0,1)}.$

4 Form the acceptance ratio

$$\rho = \min \left(1, \frac{\pi_0(x_*)}{q(x_* | x_{t-1})} \frac{q(x_{t-1} | x_*)}{\pi_0(x_{t-1})} \right).$$

5 **if** $u < \rho$, **then** $x_t \leftarrow x_*$ **else** $x_t \leftarrow x_{t-1}.$

6 $\mathcal{S} \leftarrow \mathcal{S} \cup \{x_t\}.$