

Multi-differential Higgs theory predictions: vetos and double-differential resummation

Luca Rottoli

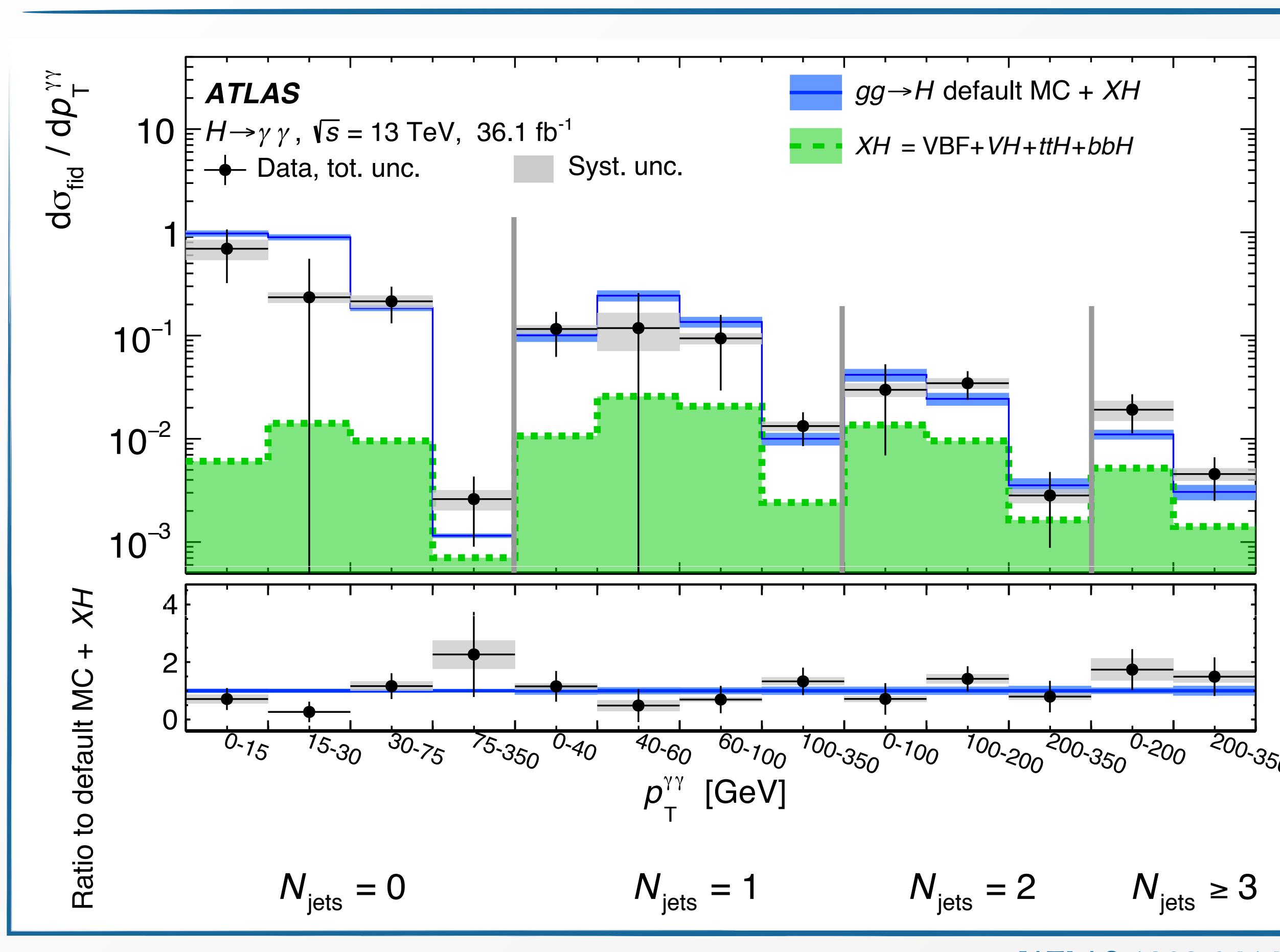
University of Milan-Bicocca & LBNL



Based on 1909.04704 with P. Monni and P. Torrielli



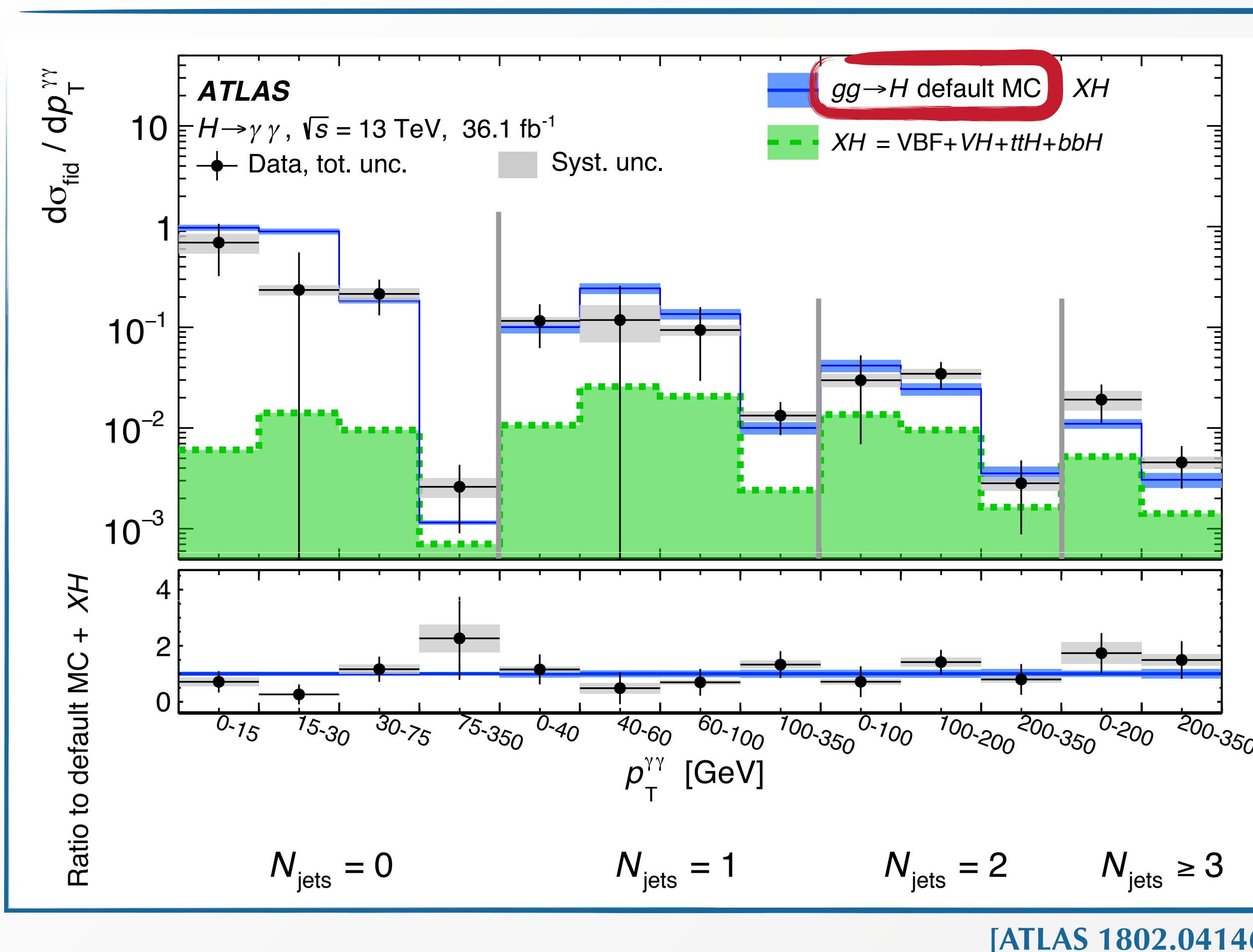
The Higgs transverse momentum



[ATLAS 1802.04146]

- Relatively easy to measure
- Sensitivity to New Physics (e.g. **light Yukawa** couplings, **trilinear** Higgs coupling) [Bishara et al. '16][Soreq et al. '16] [Bizon et al. 1610.05771]
- Experimental analyses categorize events into **jet bins** according to the jet multiplicity
- Increased sensitivity to Higgs boson kinematics, spin-CP properties, BSM effects...

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- Current description of double-differential distributions based on predictions with **NNLO+PS accuracy** [Hamilton et al. 1309.0017]
- Similar comments apply also to other analyses (e.g. VH with boosted Higgs, W^+W^- production...)

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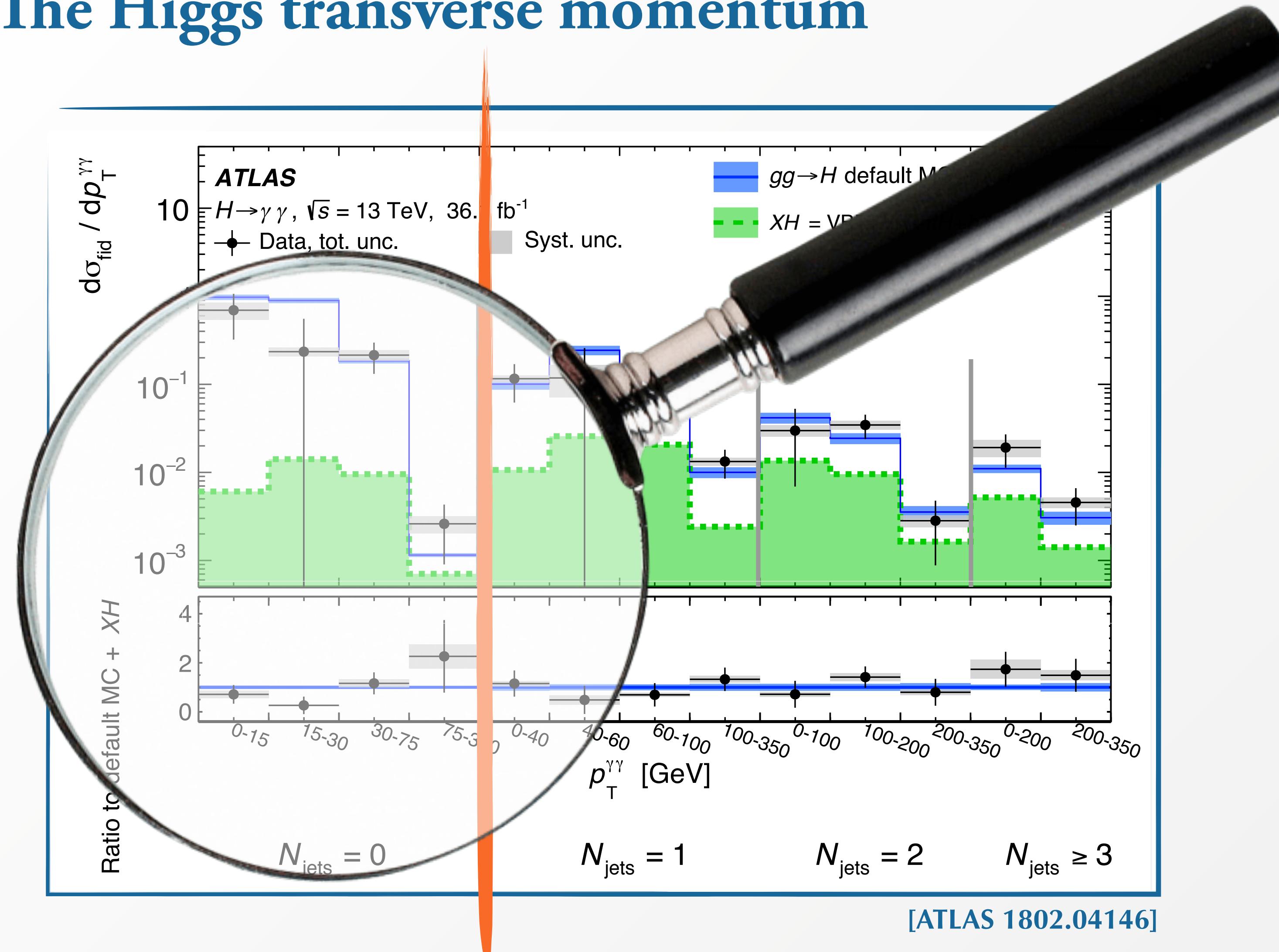
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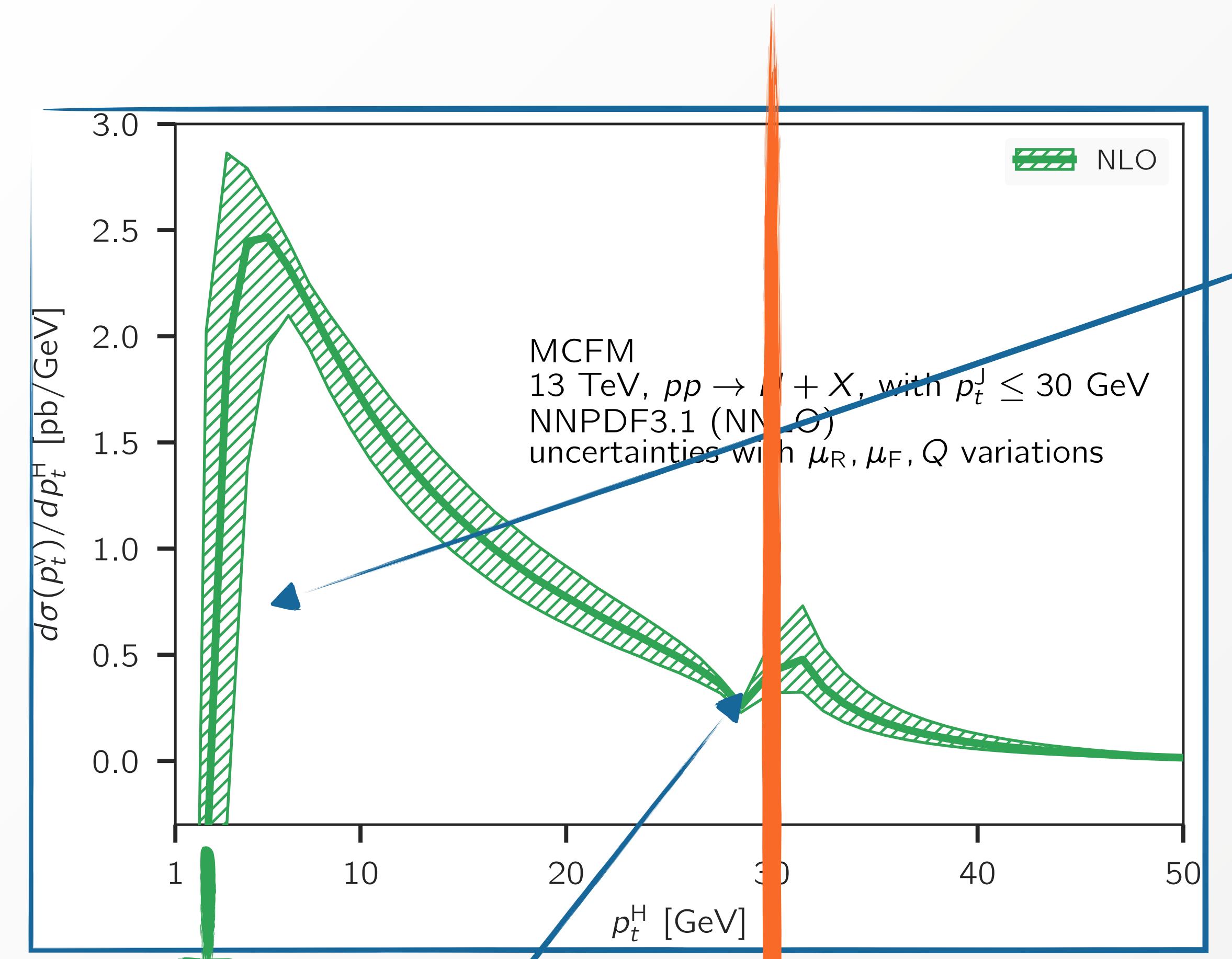


$$p_T^J \leq 30 \text{ GeV}$$

- Focus on the **zero-jet bin** $p_T^J \leq p_T^{J,v}$
- Jet veto enforced to enhance the Higgs signal with respect to its backgrounds (e.g. W^+W^- event selection) or study of different production channels (e.g. STXS)

[ATLAS 1802.04146]

The appearance of large logarithms



$$L = \ln(|\vec{p}_\perp^H + \vec{p}_\perp^J|/p_\perp^H) \quad \text{Sudakov shoulder logarithms}$$

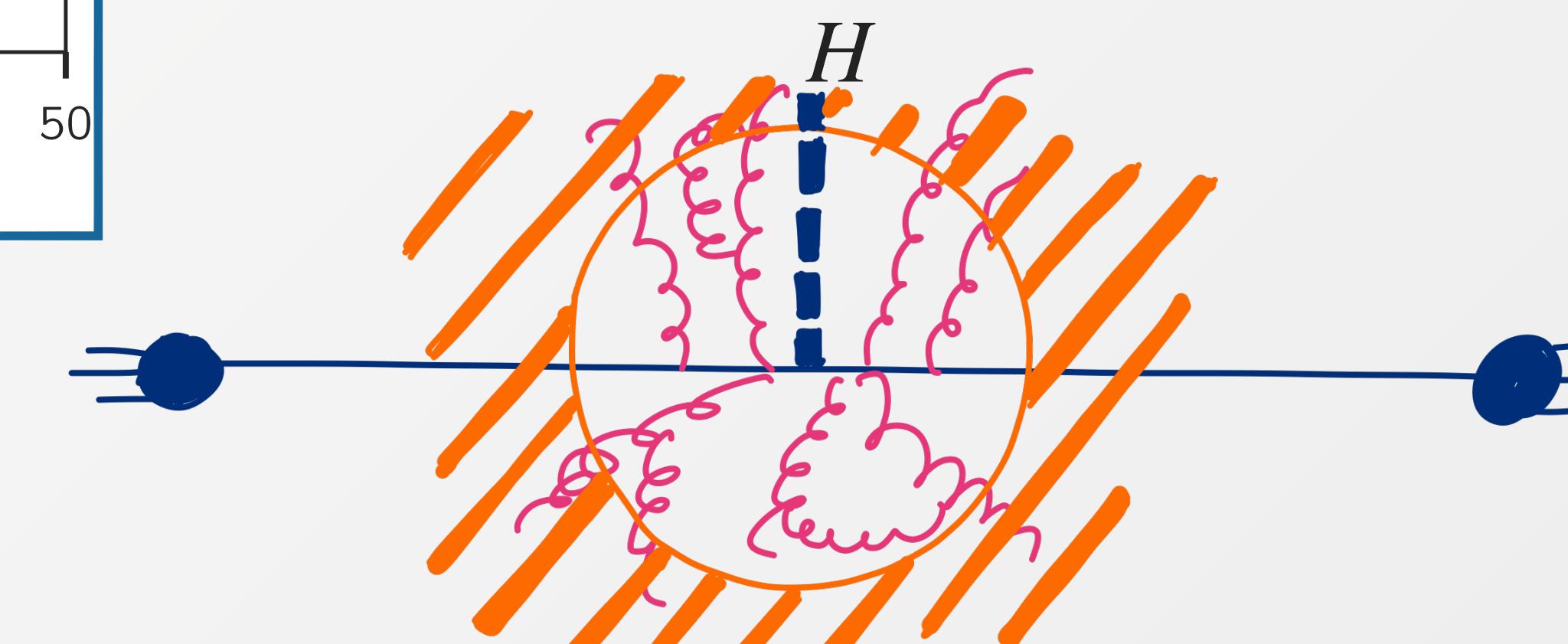
Jet veto = 30 GeV

Large **transverse momentum** logarithms

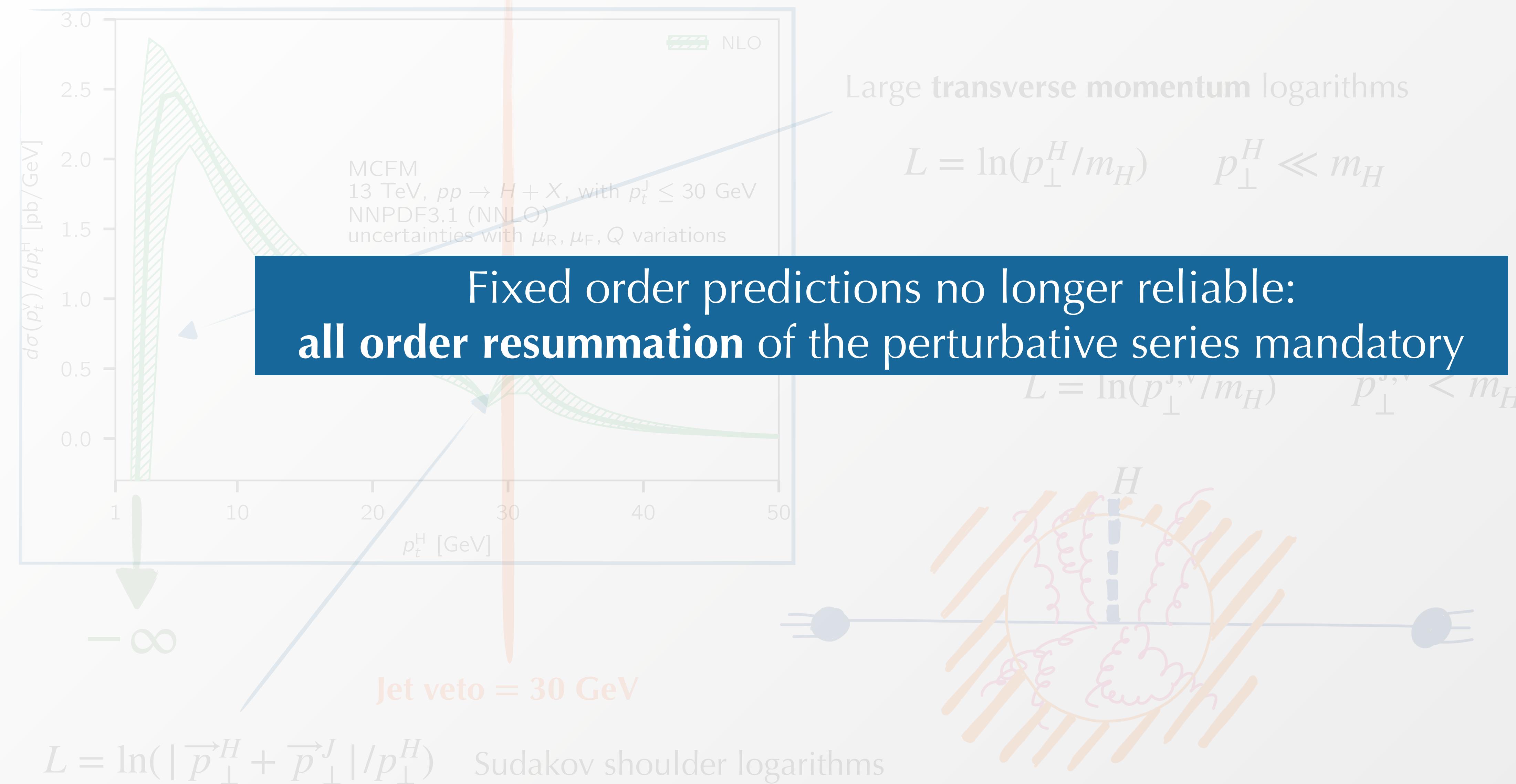
$$L = \ln(p_\perp^H/m_H) \quad p_\perp^H \ll m_H$$

Large(ish) **jet veto** logarithms

$$L = \ln(p_\perp^{J,v}/m_H) \quad p_\perp^{J,v} < m_H$$



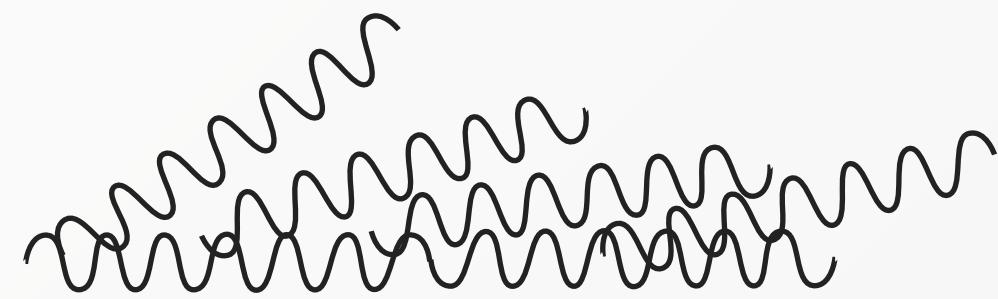
The appearance of large logarithms



An example : resummation of the transverse momentum spectrum

Resummation of transverse momentum is particularly delicate because p_\perp is a **vectorial quantity**

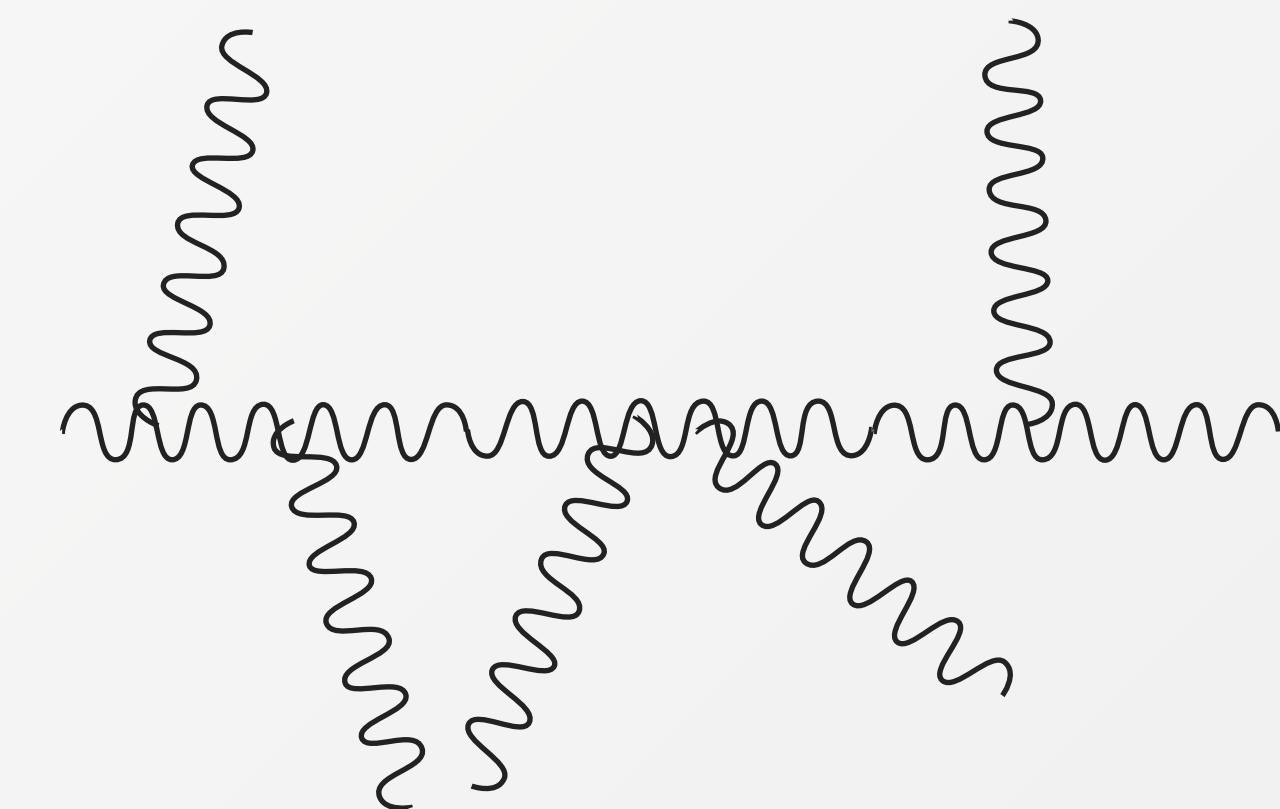
Two concurring mechanisms leading to a system with small p_\perp



$$p_\perp^2 \sim k_{t,i}^2 \ll m_H^2$$

cross section naturally suppressed as there is no phase space left for gluon emission
(Sudakov limit)

Exponential suppression



$$\sum_{i=1}^n \vec{k}_{t,i} \simeq 0$$

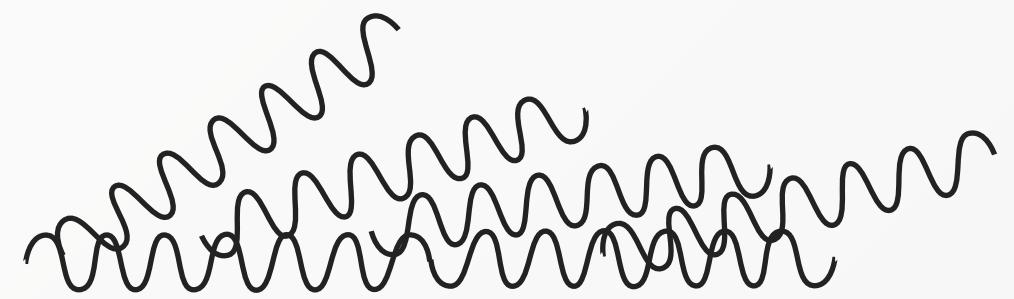
Large kinematic cancellations
 $p_\perp \sim 0$ far from the Sudakov limit

Power suppression

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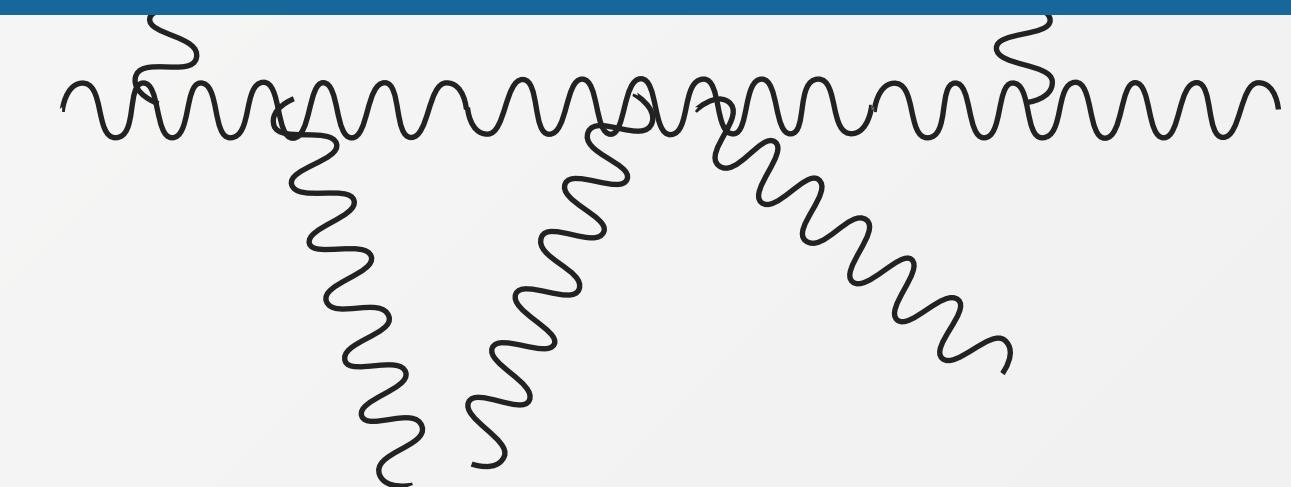


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Dominant at small p_\perp



$$\sum_{i=1}^n \vec{k}_{t,i} \simeq 0$$

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Power suppression

Resummation of the transverse momentum spectrum in direct space

Highly **non-trivial problem**. Solution to the problem recently formulated by extending the CAESAR/ARES approach to deal with observables with azimuthal cancellations: **RadISH** approach

[Monni, Re, Torrielli '16][Bizon, Monni, Re, LR, Torrielli '17]

[Banfi, Salam, Zanderighi '01, '03, '04]

[Banfi, McAslan, Monni, Zanderighi '15, '16]

[Banfi, El-Menoufi, Monni '19]

Problem recently addressed also within SCET [Ebert, Tackmann '17]

Result at NLL accuracy can be written as

$$\sigma(p_\perp) = \sigma_0 \int \frac{dv_1}{v_1} \int_0^{2\pi} \frac{d\phi_1}{2\pi} e^{-R(v_1)} \quad \text{Simple observable}$$

$$v_i = k_{t,i}/m_H, \quad \zeta_i = v_i/v_1$$

$$\times e^{R'(v_1)} R' (v_1) \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R' (\zeta_i v_1) \Theta \left(p_\perp - |\vec{k}_{t,i} + \dots + \vec{k}_{t,n+1}| \right)$$

Transfer function

Formula can be evaluated with Monte Carlo method; dependence on ϵ vanishes (as $\mathcal{O}(\epsilon)$) and result is **finite** in four dimensions

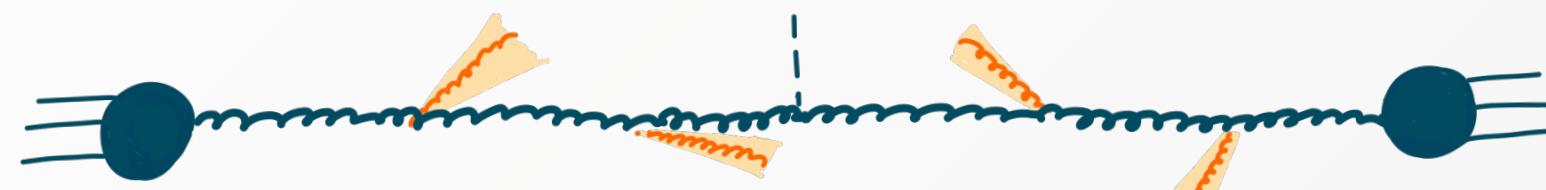
Subleading effects retained: no divergence at small p_\perp , power-like behaviour respected

Direct space formulation: generality

[Parisi, Petronzio '79; Collins, Soper, Sterman '85]

NLL result for p_{\perp}^J

$$\sigma(p_{\perp}^J) = \sigma_0 e^{Lg_1(\alpha_s \beta_0 L) + g_2(\alpha_s \beta_0 L)}$$



NLL result for p_{\perp}^H

$$\sigma(p_{\perp}^H) = \sigma_0 \int d^2 \vec{p}_{\perp}^H \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i \vec{b} \cdot \vec{p}_{\perp}^H} e^{-R_{\text{NLL}}(L)}$$



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General formula for a generic transverse observable at NLL [Bizon, Monni, Re, LR, Torrielli '17]

$$\sigma(v) = \sigma_0 \int \frac{dk_{t,1}}{k_{t,1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi} e^{-R(k_{t,1})} R' (k_{t,1}) d\mathcal{Z} \Theta (v - V(k_1, \dots, k_{n+1}))$$

$$d\mathcal{Z} = \epsilon^{R'(k_{t,1})} \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R' (\zeta_i k_{t,1})$$

$$R_{\text{NLL}}(L) = -Lg_1(\alpha_s \beta_0 L) - g_2(\alpha_s \beta_0 L)$$

$$L = \ln(k_{t,1}/M)$$

$$R'_{\text{NLL}}(k_t) = 4 \left(\frac{\alpha_s^{\text{CMW}}(k_t)}{\pi} C_A \ln \frac{m_H}{k_t} - \alpha_s(k_t) \beta_0 \right)$$

CMW scheme

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differential control in momentum space provides guidance to
double-differential resummation

General

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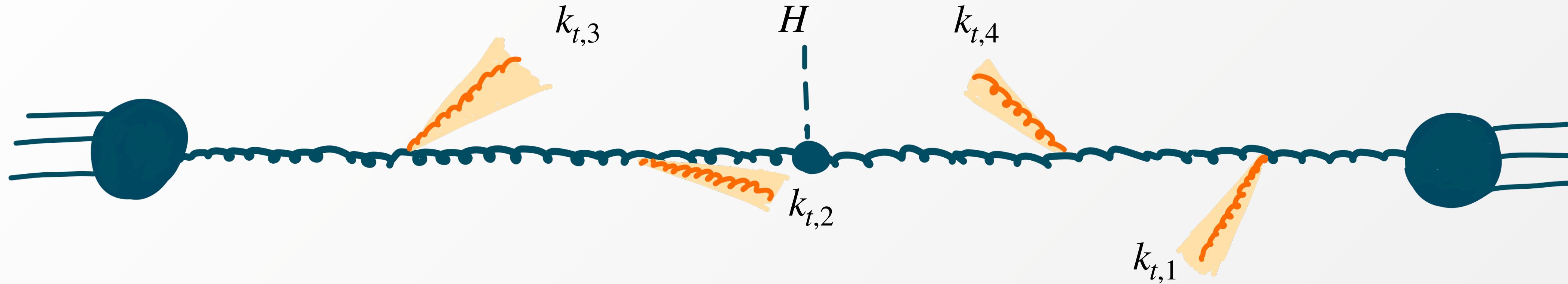
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CMW scheme

Double-differential resummation in direct space

At NLL, emissions are **strongly ordered** in angle. k_t -clustering algorithms associate **each emission** to a **different jet**



Double-differential resummation obtained by **combining measurement functions!**

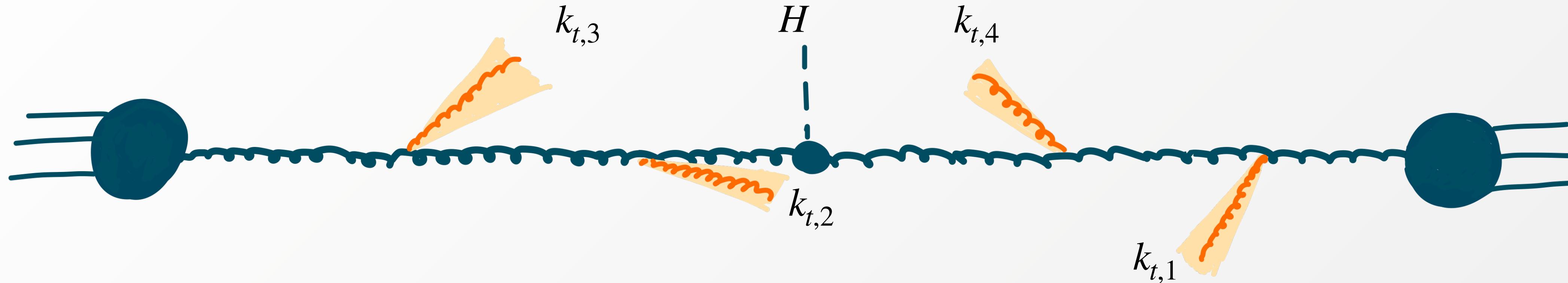
At NLL, now including parton luminosities

$$\sigma^{\text{incl}}(p_\perp^H) = \int \frac{dk_{t,1}}{k_{t,1}} \frac{d\phi_1}{2\pi} d\mathcal{Z} \partial_L \left[-e^{-R_{\text{NLL}}(k_{t,1})} \mathcal{L}_{\text{NLL}}(k_{t,1}) \right] \Theta \left(p_\perp^H - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1}| \right)$$

$$\sigma^{\text{incl}}(p_\perp^{\text{J,v}}) = \int \frac{dk_{t,1}}{k_{t,1}} \frac{d\phi_1}{2\pi} d\mathcal{Z} \partial_L \left[-e^{-R_{\text{NLL}}(k_{t,1})} \mathcal{L}_{\text{NLL}}(k_{t,1}) \right] \Theta \left(p_\perp^{\text{J,v}} - \max \{k_{t,1}, \dots, k_{t,n+1}\} \right)$$

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$$\mathcal{L}_{\text{NLL}}(\mu) = M_{gg \rightarrow H}^2 f_g(x_1, \mu) f_g(x_2, \mu)$$

At higher orders, it contains also coefficient functions and hard-virtual coefficients

Double-differential resummation in direct space at NNLL

Additional corrections must be included at NNLL [Banfi et al. '12][Becher et al. '13][Stewart et al. '14]

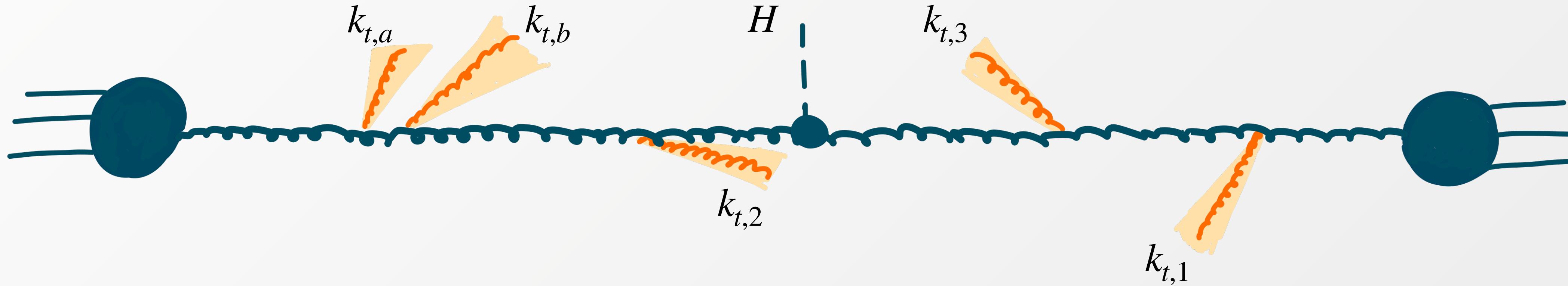
$$\sigma^{\text{NNLL}}(p_{\perp}^{\text{J,v}}, p_{\perp}^H) = \sigma_{\text{incl}}^{\text{NNLL}}(p_{\perp}^{\text{J,v}}, p_{\perp}^H) + \sigma_{\text{clust}}^{\text{NNLL}}(p_{\perp}^{\text{J,v}}, p_{\perp}^H) + \sigma_{\text{corr}}^{\text{NNLL}}(p_{\perp}^{\text{J,v}}, p_{\perp}^H)$$

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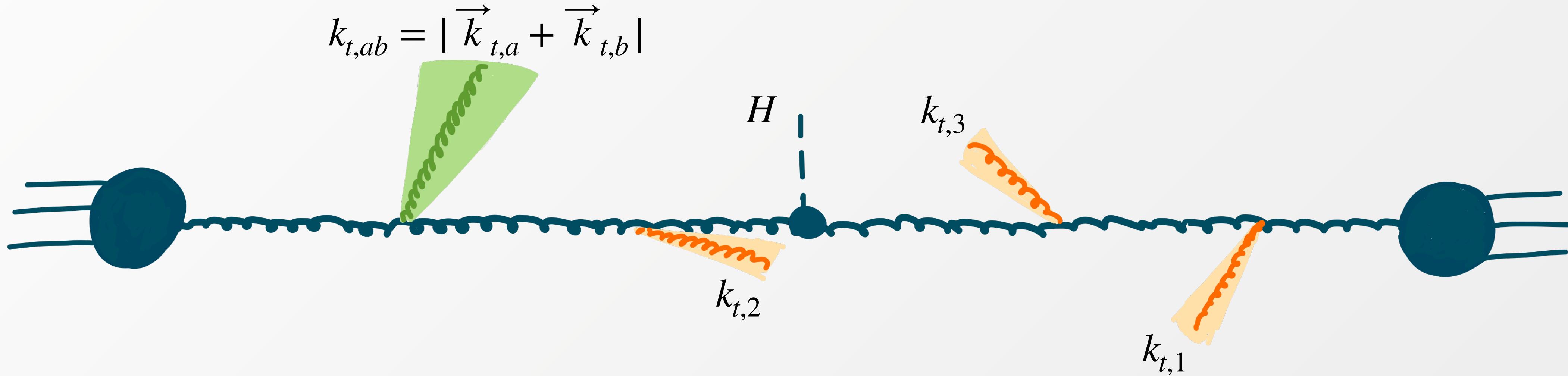


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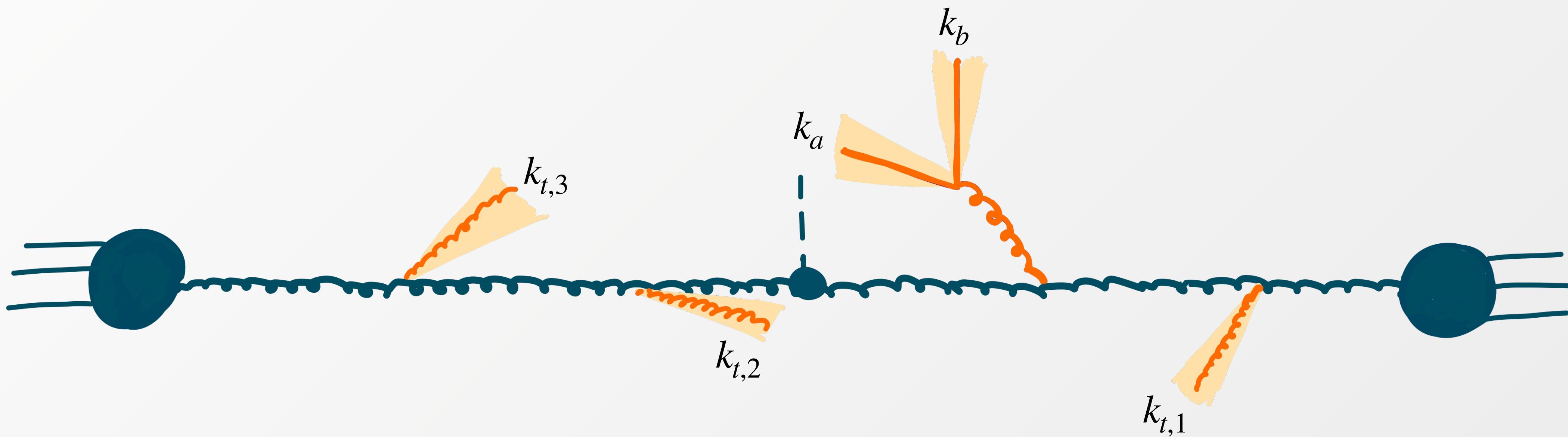


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- **correlated correction:** amends the inclusive treatment of the correlated squared amplitude for two emission accounting for configurations where the two **correlated** emissions (non abelian) are not clustered in the same jet

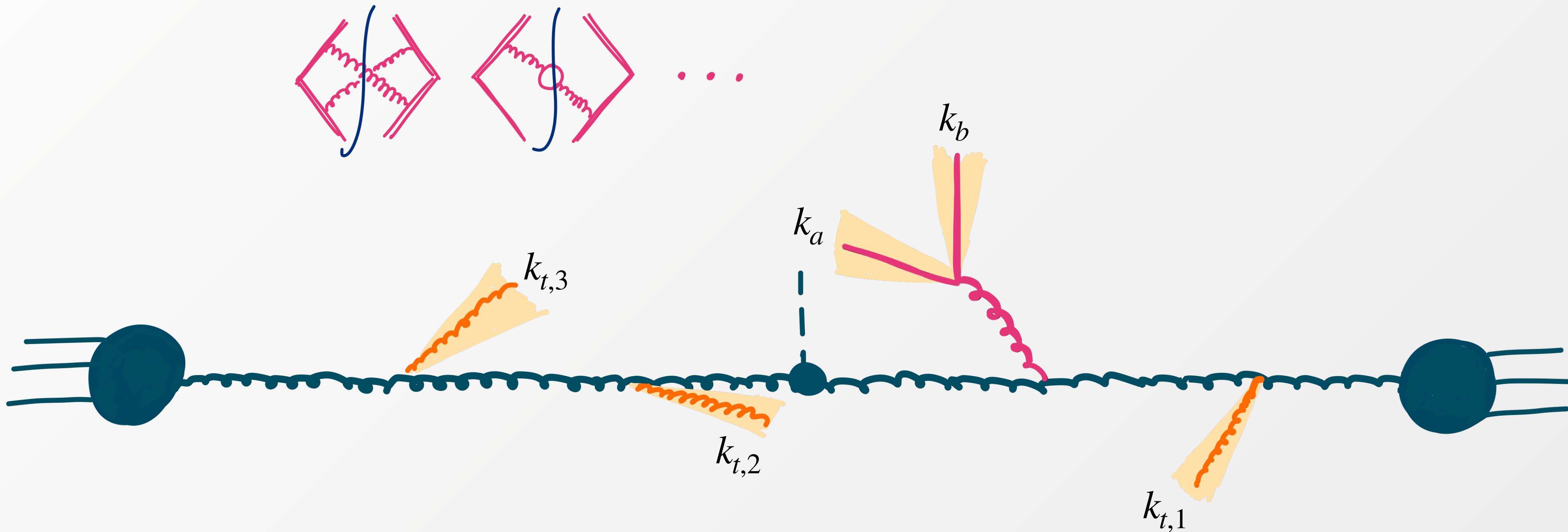


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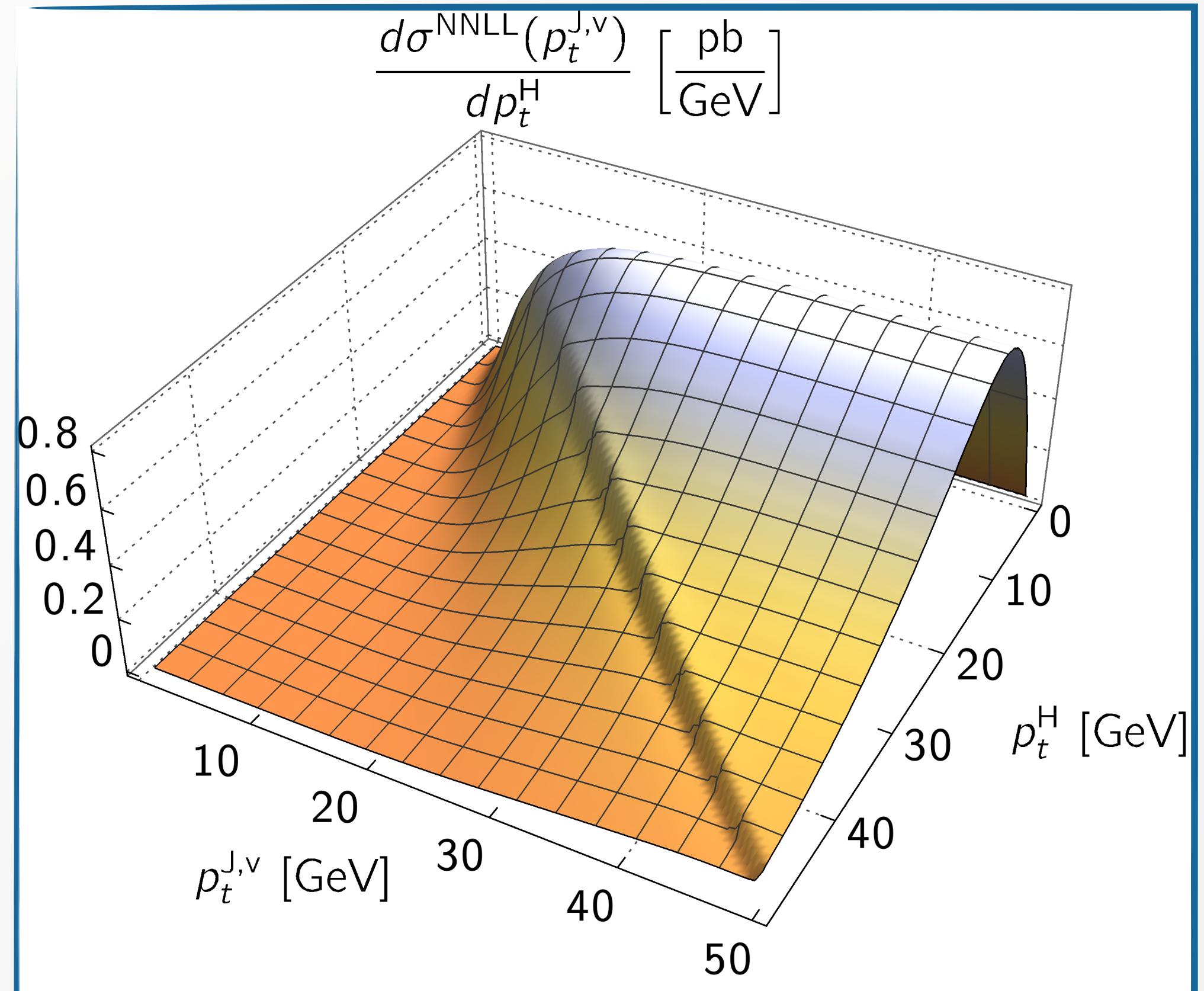
The same philosophy used at NLL can be applied at NNLL

where e.g

$$\begin{aligned}\sigma_{\text{clust}}^{\text{NNLL}}(p_{\perp}^{\text{J,v}}, p_{\perp}^H) \simeq & \int_0^\infty \frac{dk_{t,1}}{k_{t,1}} \frac{d\phi_1}{2\pi} \int d\mathcal{Z} e^{-R(k_{t,1})} 8 C_A^2 \frac{\alpha_s^2(k_{t,1})}{\pi^2} \Theta\left(p_{\perp}^{\text{J,v}} - \max_{i>1}\{k_{t,i}\}\right) \\ & \times \int_0^{k_{t,1}} \frac{dk_{t,s_1}}{k_{t,s_1}} \frac{d\phi_{s_1}}{2\pi} \int_{-\infty}^\infty d\Delta\eta_{1s_1} J_{1s_1}(R) \left[\Theta\left(p_{\perp}^{\text{J,v}} - |\vec{k}_{t,1} + \vec{k}_{t,s_1}| \right) - \Theta\left(p_{\perp}^{\text{J,v}} - k_{t,1} \right) \right] \\ & \times \Theta\left(p_{\perp}^H - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1} + \vec{k}_{t,s_1}| \right)\end{aligned}$$

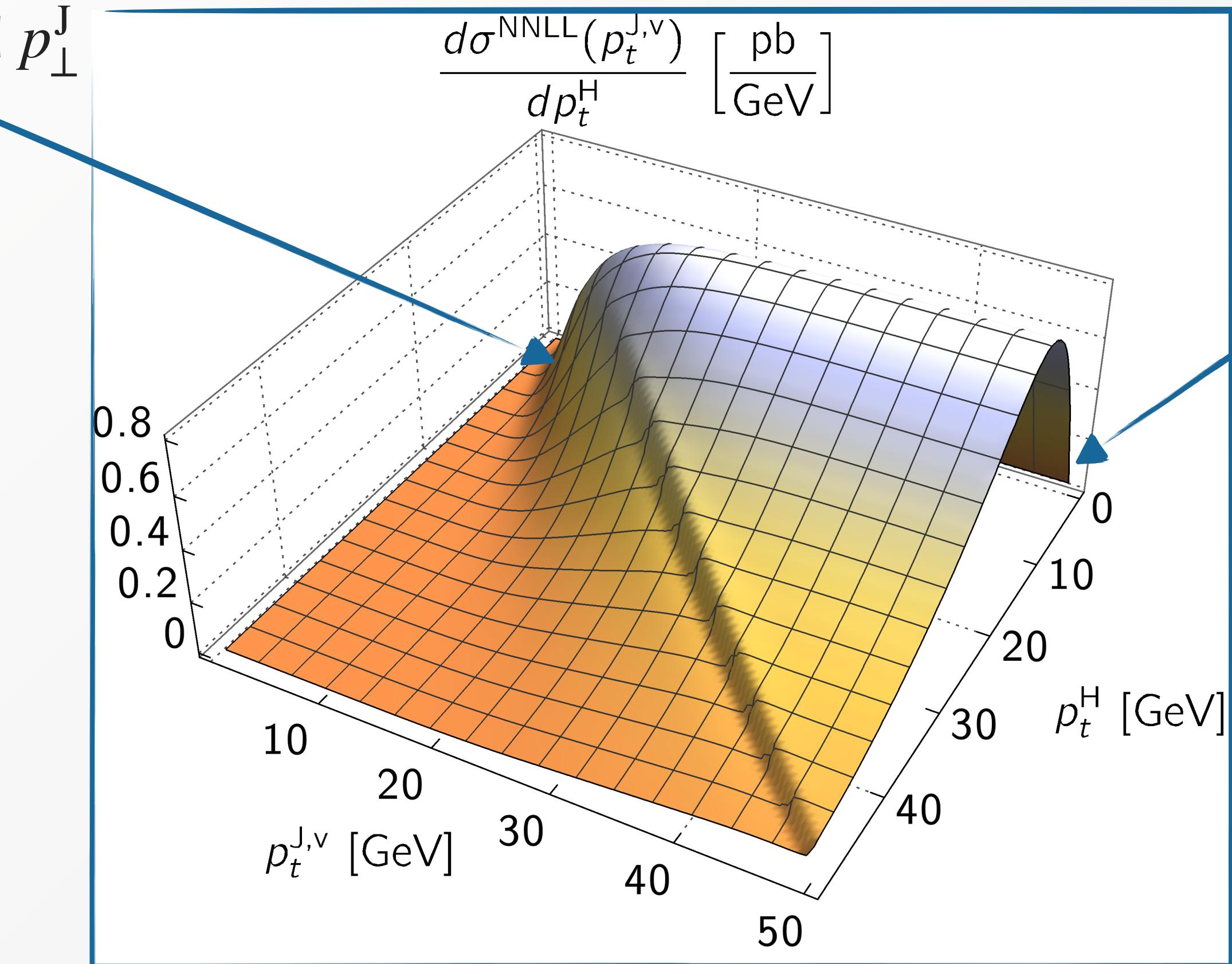
And analogously for other contributions

NNLL cross section differential in p_t^H , cumulative in $p_t^J \leq p_{\perp}^{J,v}$



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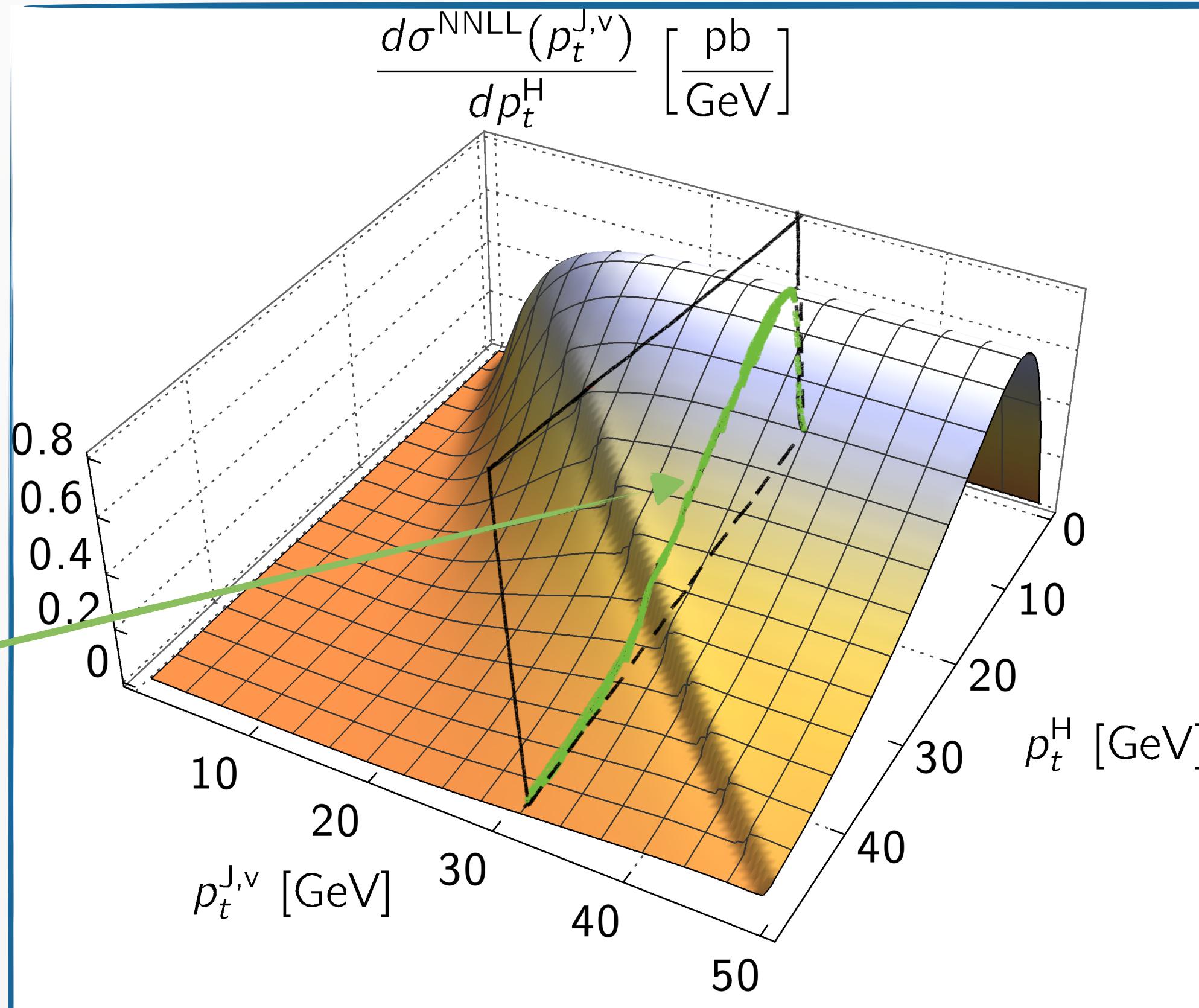
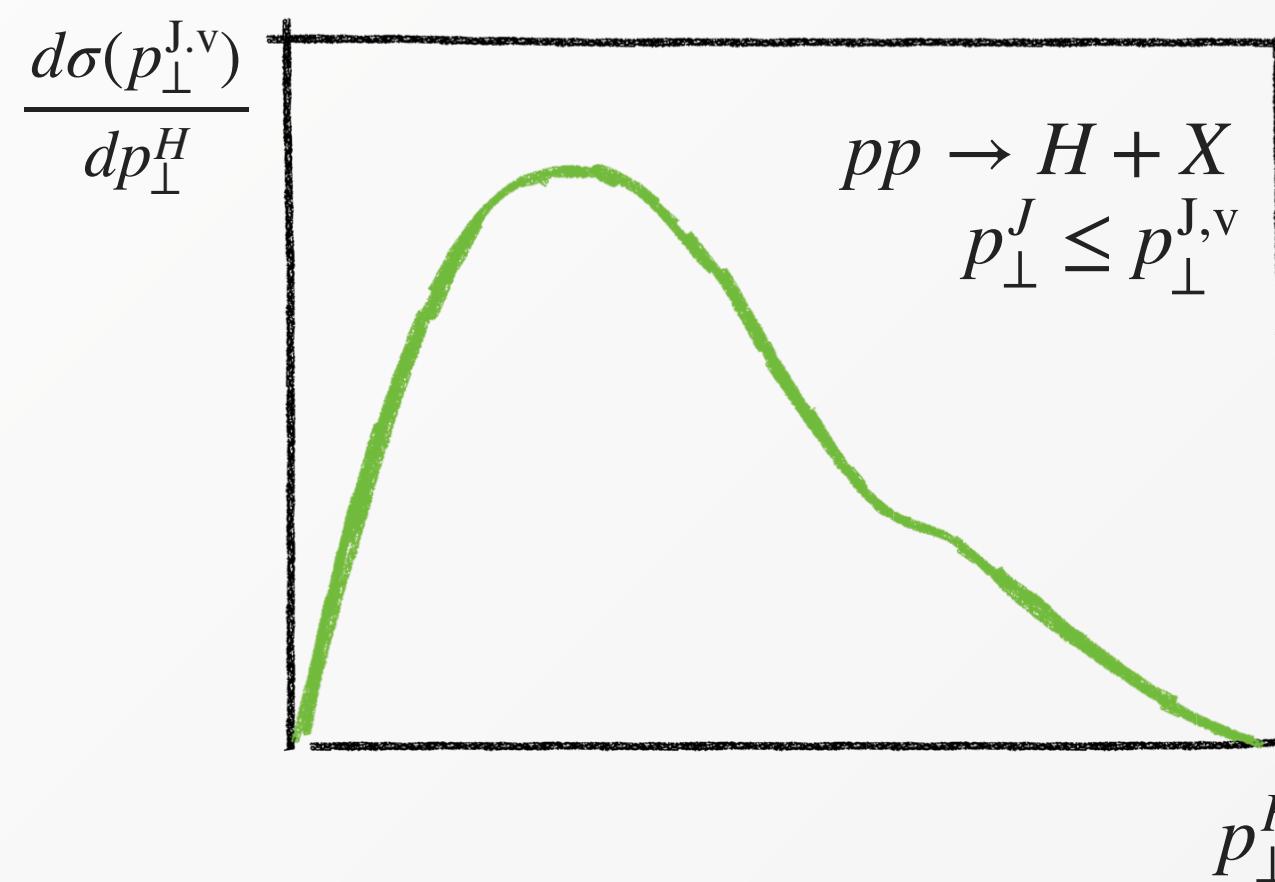
Sudakov suppression at small p_\perp^J



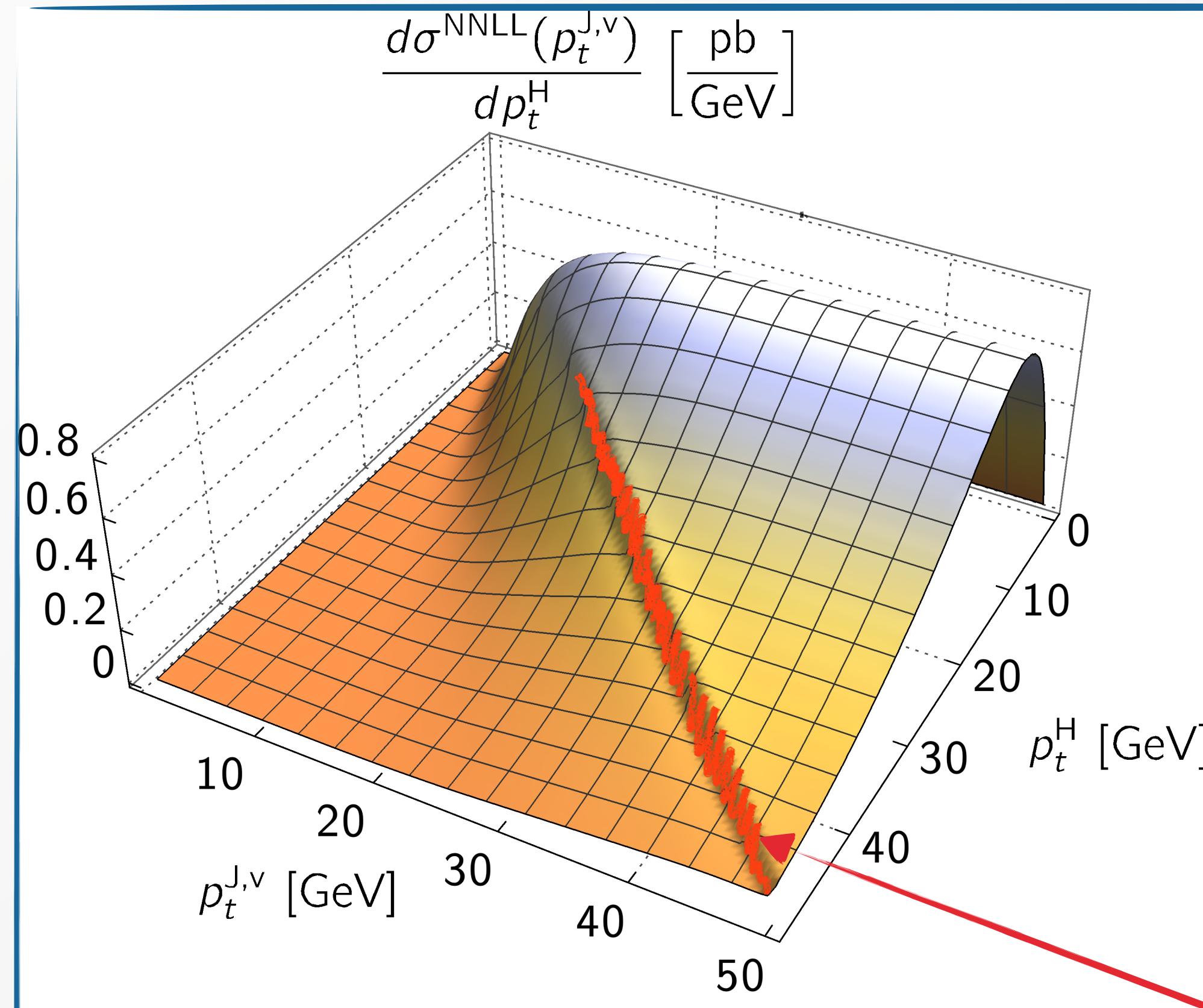
Peaked structure (Sudakov) + power-like suppression at very small p_\perp^H

NNLL cross section differential in p_{\perp}^H , cumulative in $p_{\perp}^J \leq p_{\perp}^{J,v}$

At a given value of $p_{\perp}^{J,v}$ it corresponds to the p_{\perp}^H cross section in the 0-jet bin



NNLL cross section differential in p_t^H , cumulative in $p_\perp^J \leq p_\perp^{J,v}$



Logarithms associated to the Shoulder are resummed in the limit $p_\perp^H \sim p_\perp^{J,v} \ll m_H$

[Catani, Webber '97]

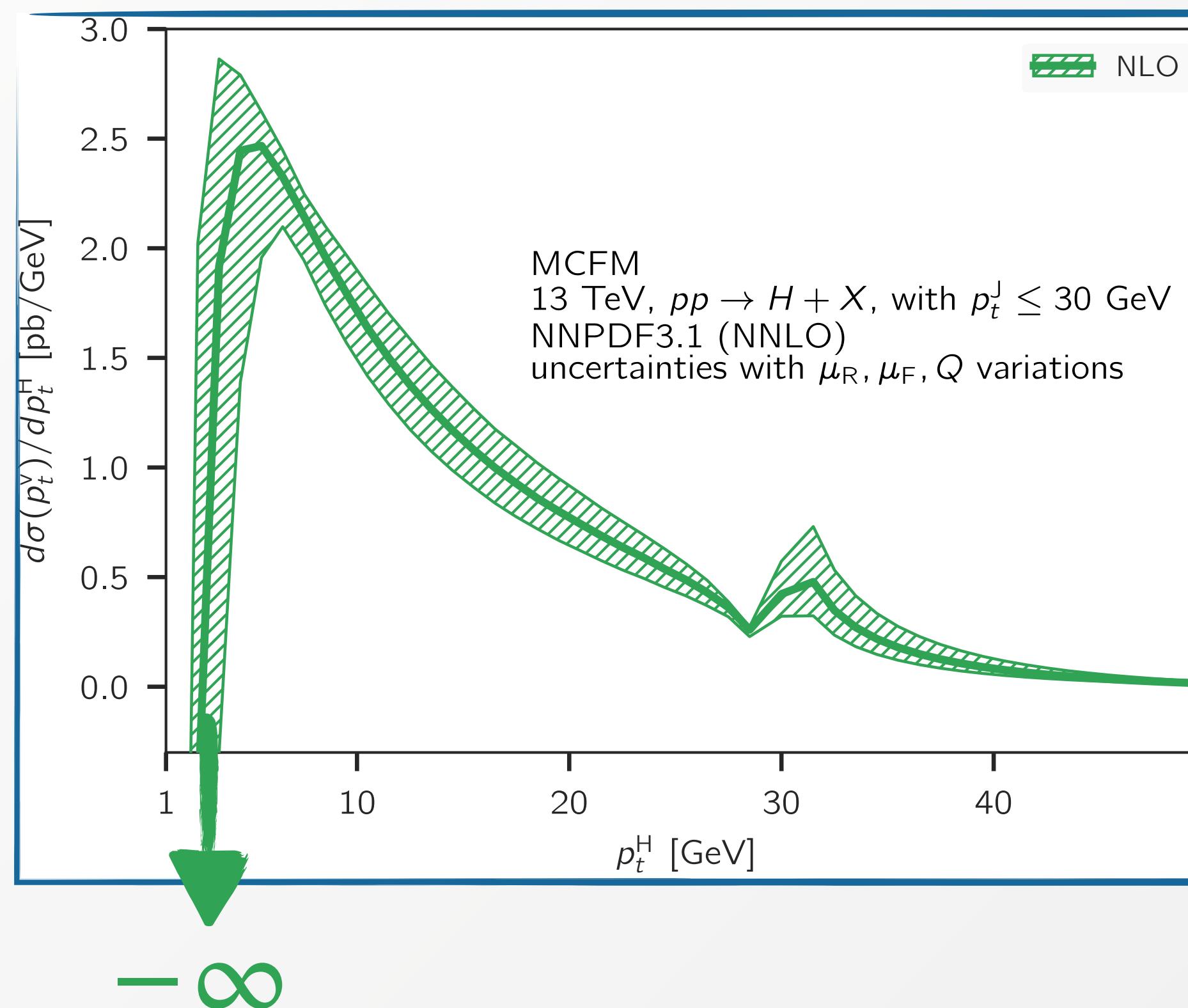
Sudakov shoulder: integrable singularity beyond LO at $p_\perp^H \simeq p_\perp^{J,v}$

LHC results: Higgs transverse momentum with a jet veto

Multiplicative matching to fixed order (NLO H+j from MCFM, NNLO H from ggHiggs)

[Campbell, Ellis, Giele,'15]

[Bonvini et al '13]



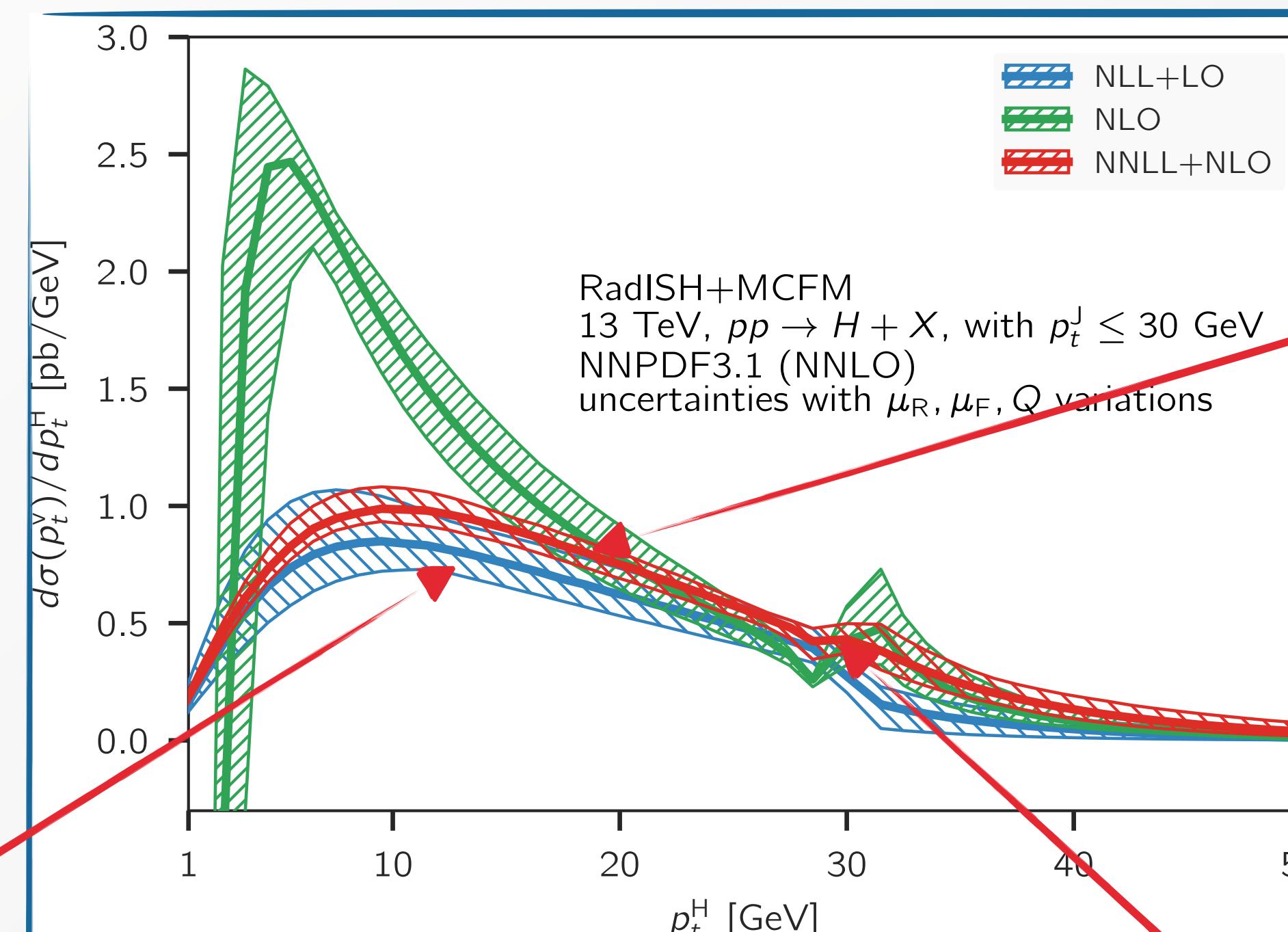
LHC results: Higgs transverse momentum with a jet veto

Multiplicative matching to fixed order (NLO H+j from MCFM, NNLO H from ggHiggs)

[Campbell, Ellis, Giele,'15]

[Bonvini et al '13]

residual uncertainties at
NNLL+NLO at the 10% level



good perturbative convergence to the left of the shoulder
above, multi-particle configurations play a substantial role

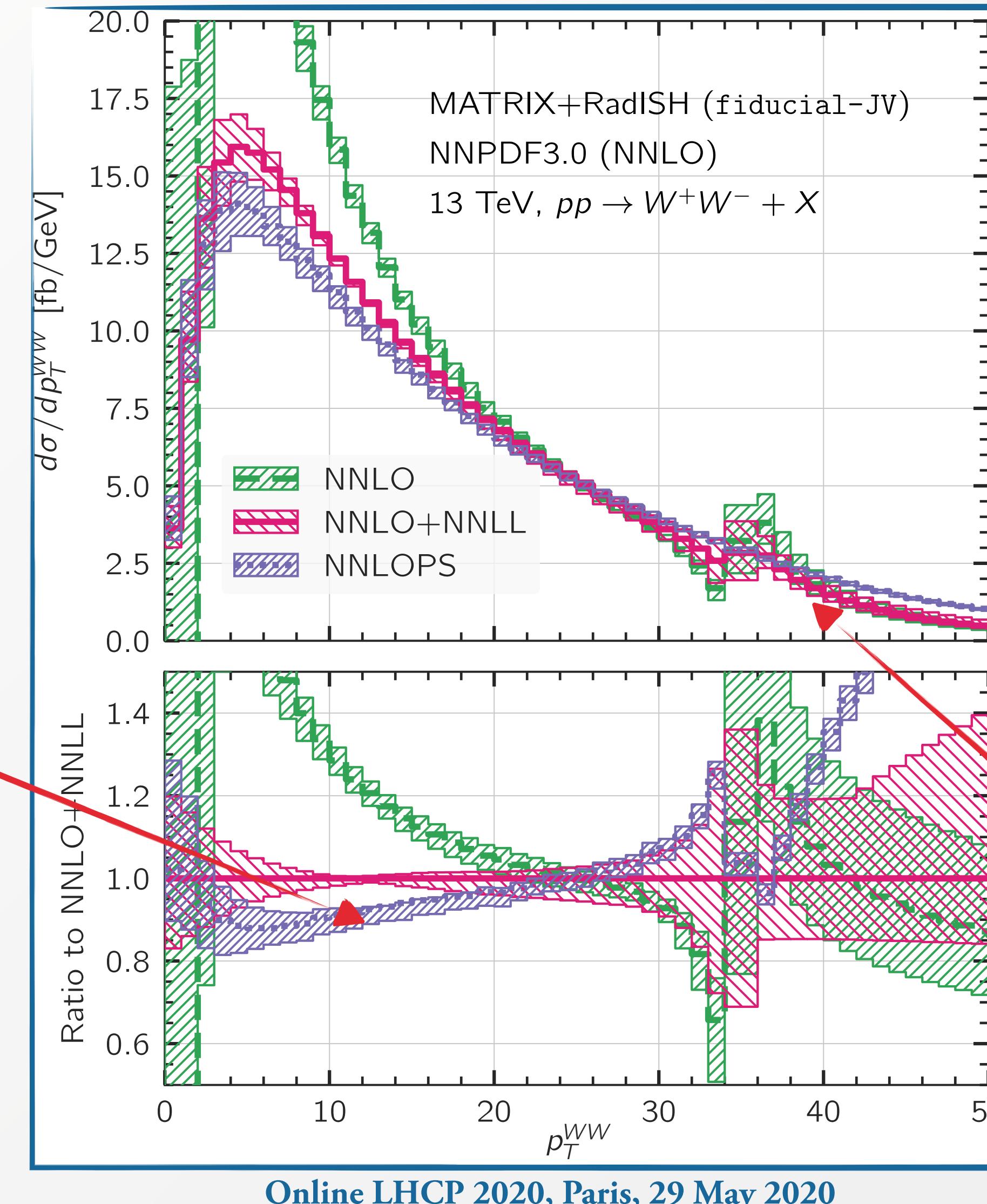
large K-factor becomes relevant
at larger p_t^H

much reduced sensitivity
to the Sudakov shoulder
with respect to NLO
spectrum

LHC applications: W^+W^- production

Fiducial region defined by a rather stringent jet veto

NNLL+NLO spectrum obtained by interfacing RadISH with MATRIX [Grazzini, Kallweit, Rathlev, Wiesemann '15, '17]



[Wiesemann, Re, Zanderighi '18]
Comparison with NNLOPS
result (much lower log
accuracy) shows differences
at the $\mathcal{O}(10\%)$ level

RadISH+MATRIX interface
for generic $2 \rightarrow 1$ and $2 \rightarrow 2$
colour singlet processes

[Kallweit, Wiesemann, Re, LR 2004.07720]

reduced sensitivity to the
Sudakov shoulder with
respect to NLO spectrum

Summary

- Precision of the data demands an increasing theoretical accuracy at the **multi-differential level** to fully exploit LHC potential
- First **joint resummation** for a **double-differential** kinematic observable involving a **jet algorithm** in hadronic collisions
- Formalism has been extended to **more complex final states**; $2 \rightarrow 1$ and $2 \rightarrow 2$ colour singlet processes available via upcoming MATRIX+RadISH interface

Backup

All-order structure of the matrix element

$$v = p_t/M$$

single-particle phase space

$$\Sigma(v) = \int d\Phi_B \mathcal{V}(\Phi_B) \sum_{n=0}^{\infty} \int \prod_{i=1}^n [dk_i] |\mathcal{M}(\Phi_B, k_1, \dots, k_n)|^2 \Theta(v - V(\{\Phi_B\}, k_1, \dots, k_n))$$

all-order form factor
(virtuals)

e.g. [Dixon, Magnea, Sterman '08]

matrix element for n real emissions

$\left| \begin{array}{c} \diagdown \\ \diagup \end{array} \right|^2 + \left| \begin{array}{c} \diagup \\ \diagdown \end{array} \right|^2 + \dots \right| ^2$

Transverse observable resummation with RadISH

- Establish a **logarithmic counting** for the squared matrix element $|\mathcal{M}(\Phi_B, k_1, \dots, k_n)|^2$

Decompose the squared amplitude in terms of **n -particle correlated blocks**, denoted by $|\tilde{\mathcal{M}}(k_1, \dots, k_n)|^2$
 $(|\tilde{\mathcal{M}}(k_1)|^2 = |\mathcal{M}(k_1)|^2)$

$$\sum_{n=0}^{\infty} |\mathcal{M}(\Phi_B, k_1, \dots, k_n)|^2 = |\mathcal{M}_B(\Phi_B)|^2$$

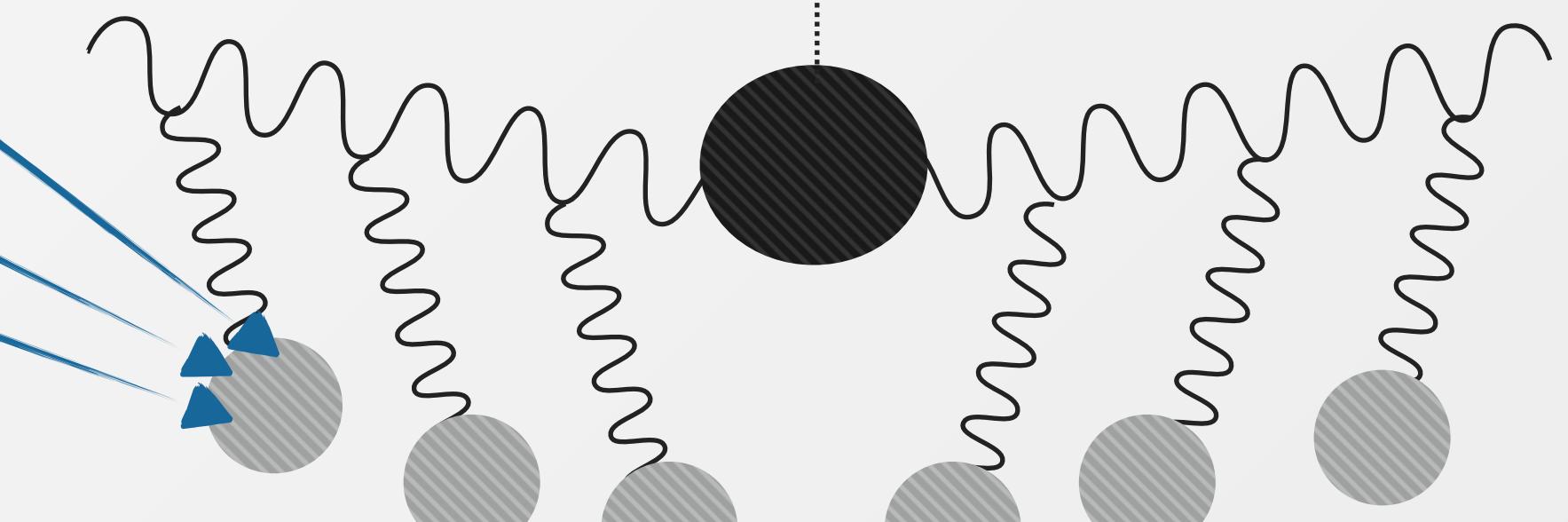
$$\times \sum_{n=0}^{\infty} \frac{1}{n!} \left\{ \prod_{i=1}^n \left(|\mathcal{M}(k_i)|^2 + \int [dk_a][dk_b] |\tilde{\mathcal{M}}(k_a, k_b)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} - \vec{k}_{ti}) \delta(Y_{ab} - Y_i) \right. \right.$$

$$\left. \left. + \int [dk_a][dk_b][dk_c] |\tilde{\mathcal{M}}(k_a, k_b, k_c)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} + \vec{k}_{tc} - \vec{k}_{ti}) \delta(Y_{abc} - Y_i) + \dots \right) \right\} \equiv |\mathcal{M}_B(\Phi_B)|^2 \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=1}^n |\mathcal{M}(k_i)|_{\text{inc}}^2$$

$$|\tilde{M}(k_1)|^2 = \frac{|M(k_1)|^2}{|M_B|^2} = |M(k_1)|^2$$

$$|\tilde{M}(k_1, k_2)|^2 = \frac{|M(k_1, k_2)|^2}{|M_B|^2} - \frac{1}{2!} |M(k_1)|^2 M(k_2)|^2$$

*expression valid for inclusive observables



Upon integration over the phase space, the expansion can be put in a **one to one correspondence** with the logarithmic structure

Systematic recipe to include terms up to the desired logarithmic accuracy

Resummation in direct space: the p_t case

2. Exploit rIRC safety to single out the IRC singularities of the real matrix element and achieve the **cancellation of the exponentiated divergences** of virtual origin

Introduce a slicing parameter $\epsilon \ll 1$ such that all inclusive blocks with $k_{t,i} < \epsilon k_{t,1}$, with $k_{t,1}$ hardest emission, can be neglected in the computation of the observable

$$\Sigma(v) = \int d\Phi_B |\mathcal{M}_B(\Phi_B)|^2 \mathcal{V}(\Phi_B)$$

unresolved emissions

$$\times \int [dk_1] |\mathcal{M}(k_1)|_{\text{inc}}^2 \left(\sum_{l=0}^{\infty} \frac{1}{l!} \int \prod_{i=2}^{l+1} [dk_i] |\mathcal{M}(k_i)|_{\text{inc}}^2 \Theta(\epsilon V(k_1) - V(k_i)) \right)$$

resolved emissions

$$\times \left(\sum_{m=0}^{\infty} \frac{1}{m!} \int \prod_{i=2}^{m+1} [dk_i] |\mathcal{M}(k_i)|_{\text{inc}}^2 \Theta(V(k_i) - \epsilon V(k_1)) \Theta(v - V(\Phi_B, k_1, \dots, k_{m+1})) \right)$$

Unresolved emission doesn't contribute to the evaluation of the observable: it can be exponentiated directly and employed to cancel the virtual divergences, giving rise to a Sudakov radiator

$$\mathcal{V}(\Phi_B) \exp \left\{ \int [dk] |\mathcal{M}(k)|_{\text{inc}}^2 \Theta(\epsilon V(k_1) - V(k)) \right\} \simeq e^{-R(\epsilon V(k_1))}$$

Resummation in direct space: the p_t case

Result at NLL accuracy can be written as

$$\Sigma(v) = \sigma^{(0)} \int \frac{dv_1}{v_1} \int_0^{2\pi} \frac{d\phi_1}{2\pi} e^{-R(\epsilon v_1)} R'(v_1)$$
$$\times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R'(\zeta_i v_1) \Theta(v - V(\Phi_B, k_1, \dots, k_{n+1}))$$
$$v_i = V(k_i), \quad \zeta_i = v_i/v_1$$

Formula can be evaluated with Monte Carlo method; dependence on ϵ vanishes exactly and result is finite in four dimensions

It contains **subleading effect** which in the original CAESAR approach are disposed of by expanding R and R' around v

$$\cancel{R(\epsilon v_1) = R(v) + \frac{dR(v)}{d\ln(1/v)} \ln \frac{v}{\epsilon v_1} + \mathcal{O}\left(\ln^2 \frac{v}{\epsilon v_1}\right)}$$
$$\cancel{R'(v_i) = R'(v) + \mathcal{O}\left(\ln \frac{v}{v_i}\right)}$$

Not possible! valid only if the ratio v_i/v remains of order one in the whole emission phase space, but for observables which feature kinematic cancellations there are configurations with $v_i \gg v$. **Subleading effects necessary**

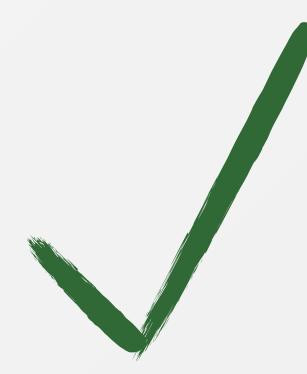
Resummation in direct space: the p_t case

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$$\times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R'(\zeta_i v_1) \Theta(v - V(\Phi_B, k_1, \dots, k_{n+1}))$$
$$v_i = V(k_i), \quad \zeta_i = v_i/v_1$$

Formula can be evaluated with Monte Carlo method; dependence on ϵ vanishes exactly and result is finite in four dimensions

Convenient to perform an expansion around k_{t1} (more efficient and simpler implementation)

$$R(\epsilon k_{t1}) = R(k_{t1}) + \frac{dR(k_{t1})}{d \ln(1/k_{t1})} \ln \frac{1}{\epsilon} + \mathcal{O}\left(\ln^2 \frac{1}{\epsilon}\right)$$
$$R'(k_{ti}) = R'(k_{t1}) + \mathcal{O}\left(\ln \frac{k_{t1}}{k_{ti}}\right)$$


Subleading effects retained: no divergence at small v , power-like behaviour respected

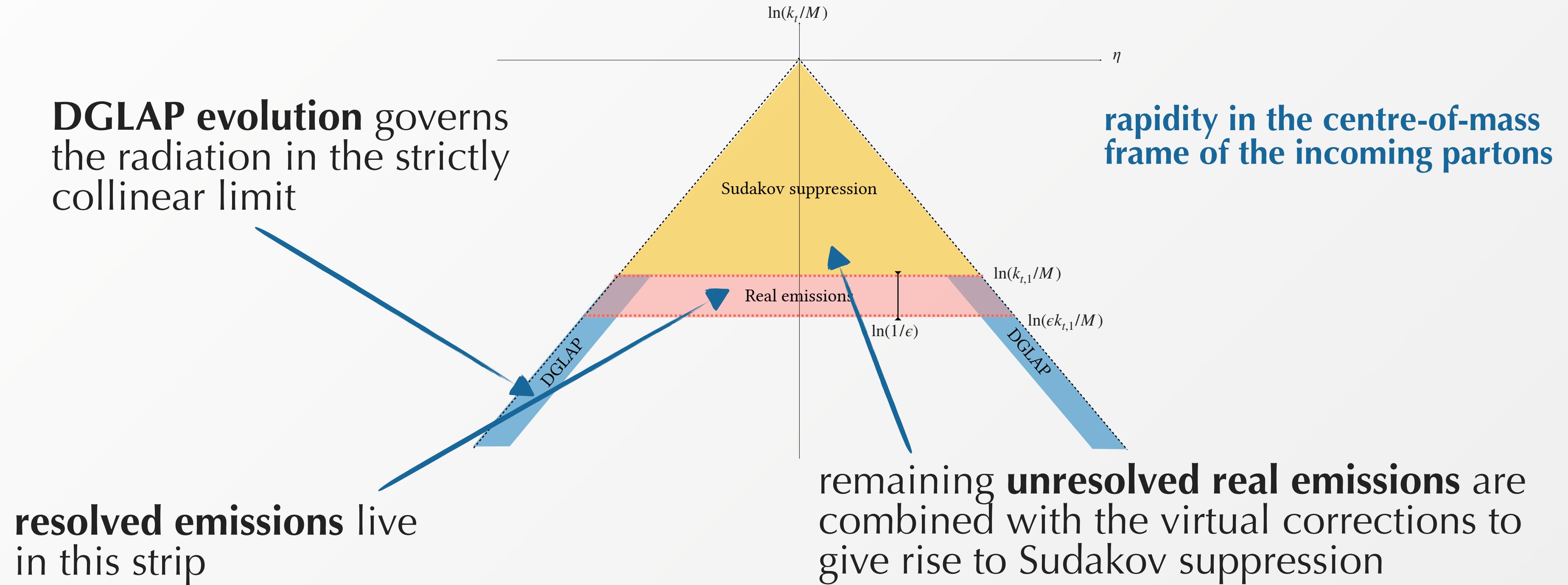
Logarithmic accuracy defined in terms of $\ln(M/k_{t1})$

Result formally equivalent to the b -space formulation

Parton luminosities

Consider configurations in which emissions are ordered in $k_{t,i}$, $k_{t,1}$ hardest emission

Phase space for each secondary emission can be depicted in the Lund diagram



- DGLAP evolution can be performed **inclusively** up to $\epsilon k_{t,1}$ thanks to rIRC safety
- In the **overlapping region** hard-collinear emissions modify the observable's value: the evolution should be performed exclusively (unintegrated in k_t)
- At NLL the real radiation can be approximated with its soft limit: DGLAP can be performed inclusively up to $k_{t,1}$ (i.e. one can evaluate $\mu_F = k_{t,1}$)

Beyond NLL

Extension to NNLL and beyond requires the systematic inclusion of the correlated blocks necessary to achieve the desired logarithmic accuracy

Moreover, one needs to **relax a series of assumptions** which give rise to subleading corrections neglected at NLL (for instance, exact rapidity bounds). These corrections can be included systematically by including additional terms in the expansion

$$R(\epsilon v_1) = R(v_1) + \frac{dR(v_1)}{d \ln(1/v_1)} \ln \frac{1}{\epsilon} + \mathcal{O}\left(\ln^2 \frac{1}{\epsilon}\right)$$

Finally, one needs to specify a complete treatment for **hard-collinear radiation**. Starting at NNLL one or more real emissions can be hard and collinear to the emitting leg, and the available phase space for subsequent real emissions changes

Two classes of contributions:

- one soft by construction and which is analogous to the R' contribution

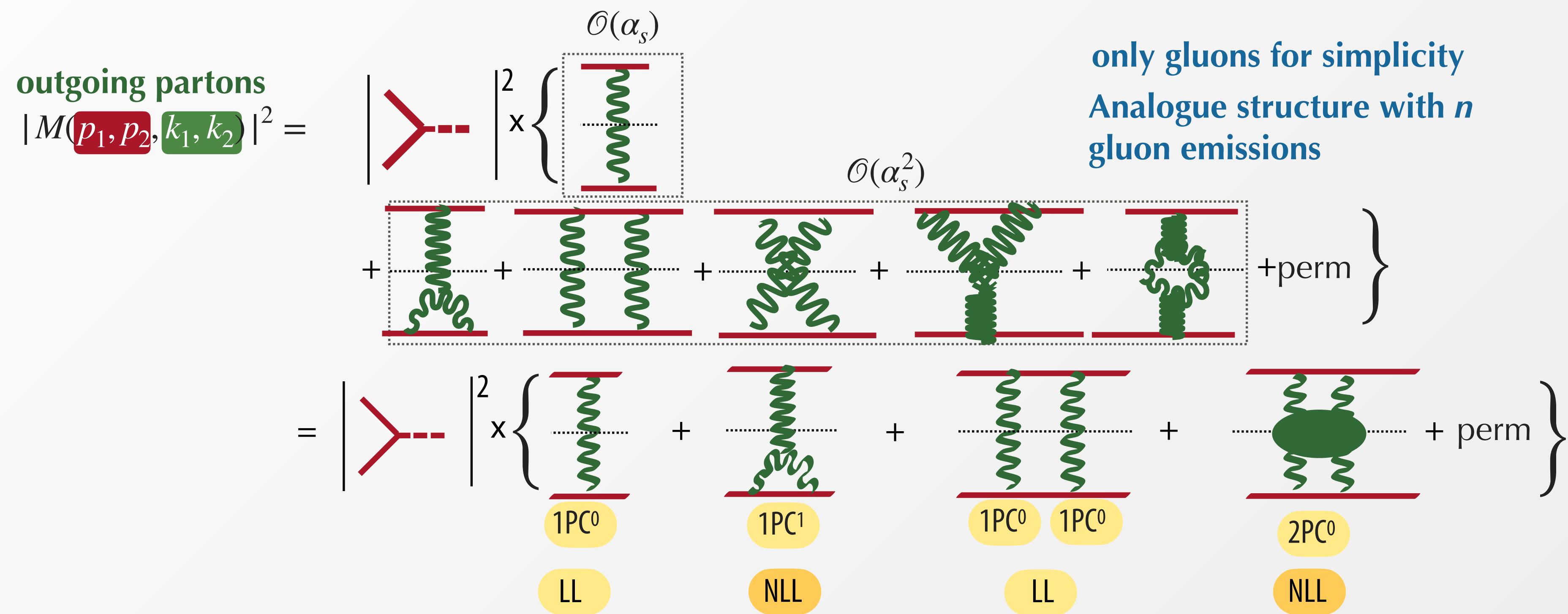
$$R'(v_i) = R'(v_1) + \mathcal{O}\left(\ln \frac{v_1}{v_i}\right)$$

- another hard and collinear (exclusive DGLAP step): last step of DGLAP evolution must be performed unintegrated in k_t

Logarithmic counting

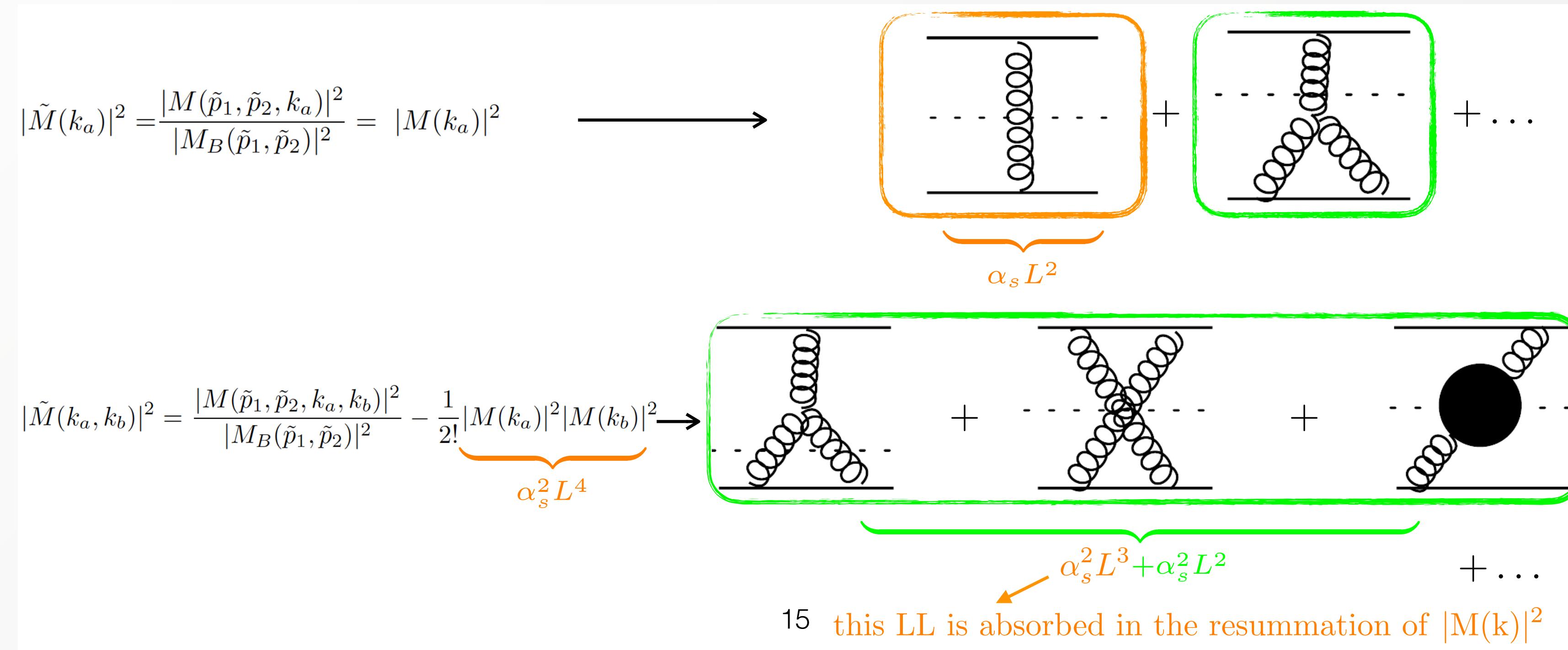
Necessary to establish a **well defined logarithmic counting**: possible to do that by decomposing the squared amplitude in terms of n-particle correlated blocks (**nPC**)

e.g. $pp \rightarrow H +$ emission of up to 2 (soft) gluons $\mathcal{O}(\alpha_s^2)$



Logarithmic counting defined in terms of **nPC blocks** (owing to rIRC safety of the observable)

Logarithmic counting: correlated blocks



Thanks to P. Monni

Resummation at NLL accuracy

Final result at NLL

$$\begin{aligned} \frac{d\Sigma(v)}{d\Phi_B} &= \int \frac{dk_{t,1}}{k_{t,1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi} e^{-R(k_{t,1})} e^{R'(k_{t,1})} \mathcal{L}_{\text{NLL}}(k_{t,1}) R'(k_{t,1}) \\ &\times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R'(\zeta_i k_{t,1}) \Theta(v - V(\Phi_B, k_1, \dots, k_{n+1})) \end{aligned}$$

This formula can be evaluated by means of fast Monte Carlo methods **RadISH** (Radiation off Initial State Hadrons)

Parton luminosity at NLL reads

$$\mathcal{L}_{\text{NLL}}(k_{t,1}) = \sum_c \frac{d|M_B|_{c\bar{c}}^2}{d\Phi_B} f_c(x_1, k_{t,1}^2) f_{\bar{c}}(x_2, k_{t,1}^2)$$

At higher logarithmic accuracy, it includes **coefficient functions** and **hard-virtual** corrections

Result at **N³LL accuracy**

$$\begin{aligned} \frac{d\Sigma(v)}{d\Phi_B} = & \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} \partial_L \left(-e^{-R(k_{t1})} \mathcal{L}_{\text{N}^3\text{LL}}(k_{t1}) \right) \int d\mathcal{Z}[\{R', k_i\}] \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \\ & + \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(k_{t1})} \int d\mathcal{Z}[\{R', k_i\}] \int_0^1 \frac{d\zeta_s}{\zeta_s} \frac{d\phi_s}{2\pi} \left\{ \left(R'(k_{t1}) \mathcal{L}_{\text{NNLL}}(k_{t1}) - \partial_L \mathcal{L}_{\text{NNLL}}(k_{t1}) \right) \right. \\ & \times \left(R''(k_{t1}) \ln \frac{1}{\zeta_s} + \frac{1}{2} R'''(k_{t1}) \ln^2 \frac{1}{\zeta_s} \right) - R'(k_{t1}) \left(\partial_L \mathcal{L}_{\text{NNLL}}(k_{t1}) - 2 \frac{\beta_0}{\pi} \alpha_s^2(k_{t1}) \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \ln \frac{1}{\zeta_s} \right) \\ & \left. + \frac{\alpha_s^2(k_{t1})}{\pi^2} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \right\} \left\{ \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_s)) - \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \right\} \\ & + \frac{1}{2} \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(k_{t1})} \int d\mathcal{Z}[\{R', k_i\}] \int_0^1 \frac{d\zeta_{s1}}{\zeta_{s1}} \frac{d\phi_{s1}}{2\pi} \int_0^1 \frac{d\zeta_{s2}}{\zeta_{s2}} \frac{d\phi_{s2}}{2\pi} R'(k_{t1}) \\ & \times \left\{ \mathcal{L}_{\text{NLL}}(k_{t1}) (R''(k_{t1}))^2 \ln \frac{1}{\zeta_{s1}} \ln \frac{1}{\zeta_{s2}} - \partial_L \mathcal{L}_{\text{NLL}}(k_{t1}) R''(k_{t1}) \left(\ln \frac{1}{\zeta_{s1}} + \ln \frac{1}{\zeta_{s2}} \right) \right. \\ & \left. + \frac{\alpha_s^2(k_{t1})}{\pi^2} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \right\} \\ & \times \left\{ \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_{s1}, k_{s2})) - \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_{s1})) - \right. \\ & \left. \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_{s2})) + \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \right\} + \mathcal{O}\left(\alpha_s^n \ln^{2n-6} \frac{1}{v}\right), \quad (3.18) \end{aligned}$$

[Bizon, Monni, Re, LR, Torrielli '17]

All ingredients to perform resummation at **N³LL accuracy** are now available

[Catani et al. '11, '12][Gehrmann et al. '14][Li, Zhu '16, Vladimirov '16][Moch et al. '18, Lee et al. '19]

Fixed-order predictions now available at **NNLO**

[A. Gehrmann-De Ridder et al. '15, 16, '17][Boughezal et al. '15, 16]

Online LHCP 2020, Paris, 29 May 2020

Matching with fixed order

Multiplicative matching performed at the **double-cumulant level**

fixed-order double-cumulative result at NNLO

double-cumulative result at NNLL

$$\sigma_{\text{NNLO}}(p_{\perp}^H < p_{\perp}^{H,v}, p_{\perp}^J < p_{\perp}^{J,v}) = \sigma_{\text{NNLO}} - \int \Theta(p_{\perp}^H > p_{\perp}^{H,v}) \vee \Theta(p_{\perp}^J > p_{\perp}^{J,v}) d\sigma_{H+J,\text{NLO}}$$

$$\sigma_{\text{match}}(p_{\perp}^H < p_{\perp}^{H,v}, p_{\perp}^J < p_{\perp}^{J,v}) = \frac{\sigma_{\text{NNLL}}(p_{\perp}^H < p_{\perp}^{H,v}, p_{\perp}^J < p_{\perp}^{J,v})}{\sigma_{\text{NNLL}}(\{p_{\perp}^{J,v}, p_{\perp}^{H,v}\} \rightarrow \infty)} \left[\sigma_{\text{NNLL}}(\{p_{\perp}^{J,v}, p_{\perp}^{H,v}\} \rightarrow \infty) \frac{\sigma_{\text{NNLO}}(p_{\perp}^H < p_{\perp}^{H,v}, p_{\perp}^J < p_{\perp}^{J,v})}{\sigma_{\text{NNLL,exp}}(p_{\perp}^H < p_{\perp}^{H,v}, p_{\perp}^J < p_{\perp}^{J,v})} \right] \mathcal{O}(\alpha_s^2)$$

asymptotic limit of the NNLL result

- NNLL+NNLO result for $p_{\perp}^{J,v}$ recovered for $p_{\perp}^{H,v} \rightarrow \infty$
- **NNLO constant** included through multiplicative matching (NNLL' accuracy)

Matching to fixed order: multiplicative matching

Cumulative cross section should reduce to the fixed order at large v

$$\Sigma_{\text{matched}}^{\text{mult}}(v) \sim \Sigma_{\text{res}}(v) \left[\frac{\Sigma_{\text{f.o.}}(v)}{\Sigma_{\text{res}}(v)} \right] \text{ expanded}$$

$$\Sigma_{\text{f.o.}}(v) = \sigma_{f.o.} - \int_v^\infty \frac{d\sigma}{dv} dv$$

- allows to include constant terms from NNLO (if N³LO total xs available)
- physical suppression at small v cures potential instabilities

To ensure that resummation does not affect the hard region of the spectrum when the matching is performed we introduce **modified logarithms**

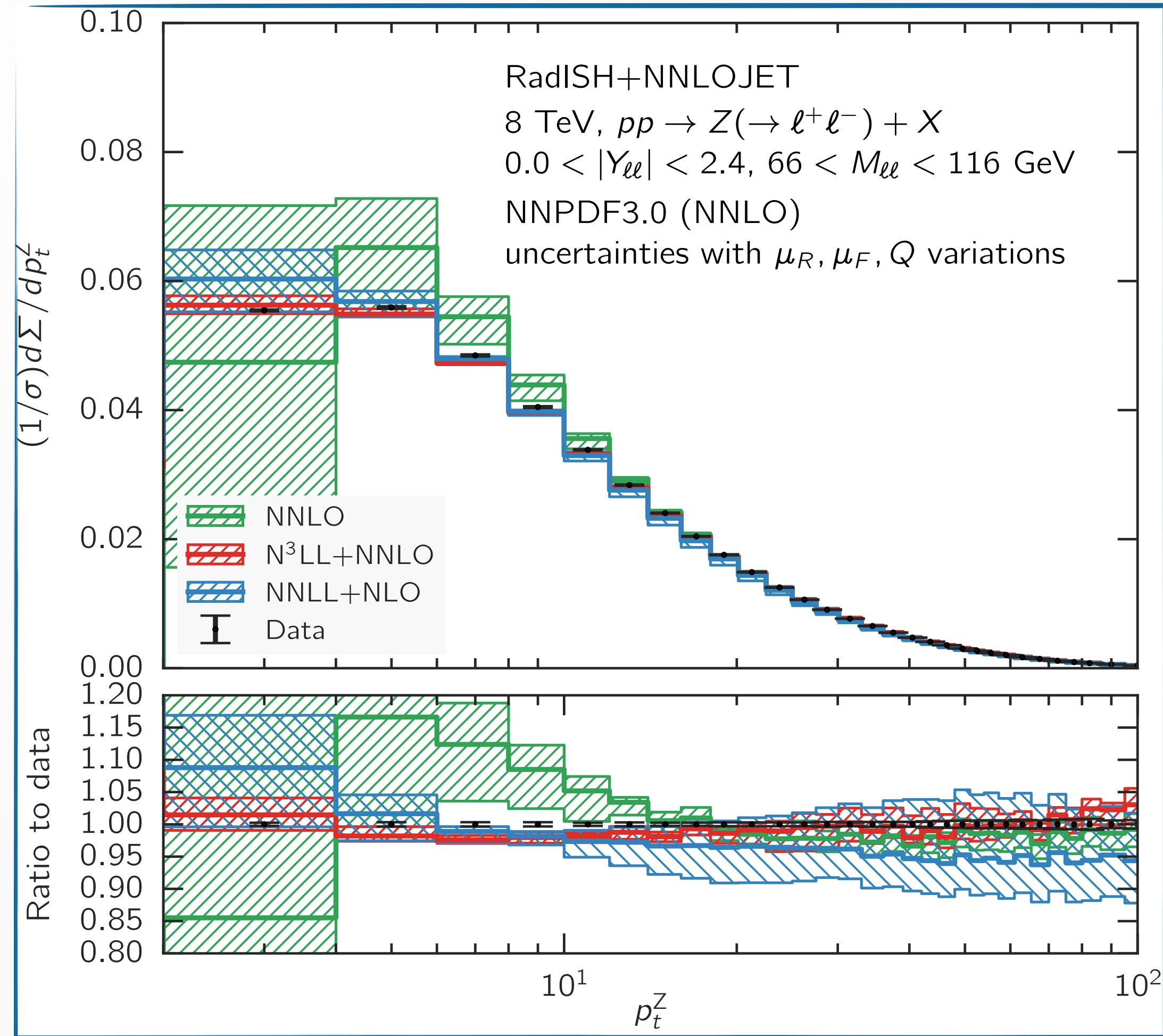
This corresponds to restrict the rapidity phase space at large k_t

$$\int_{-\ln Q/k_{t,i}}^{\ln Q/k_{t,i}} d\eta \rightarrow \int_{-\ln Q/k_{t,1}}^{\ln Q/k_{t,1}} d\eta \rightarrow \int_{-\epsilon}^{\epsilon} d\eta \rightarrow 0$$

$$\ln(Q/k_{t1}) \rightarrow \frac{1}{p} \ln \left(1 + \left(\frac{Q}{k_{t1}} \right)^p \right)$$

Q : **perturbative resummation scale**
used to probe the size of subleading logarithmic corrections
 p : arbitrary matching parameter

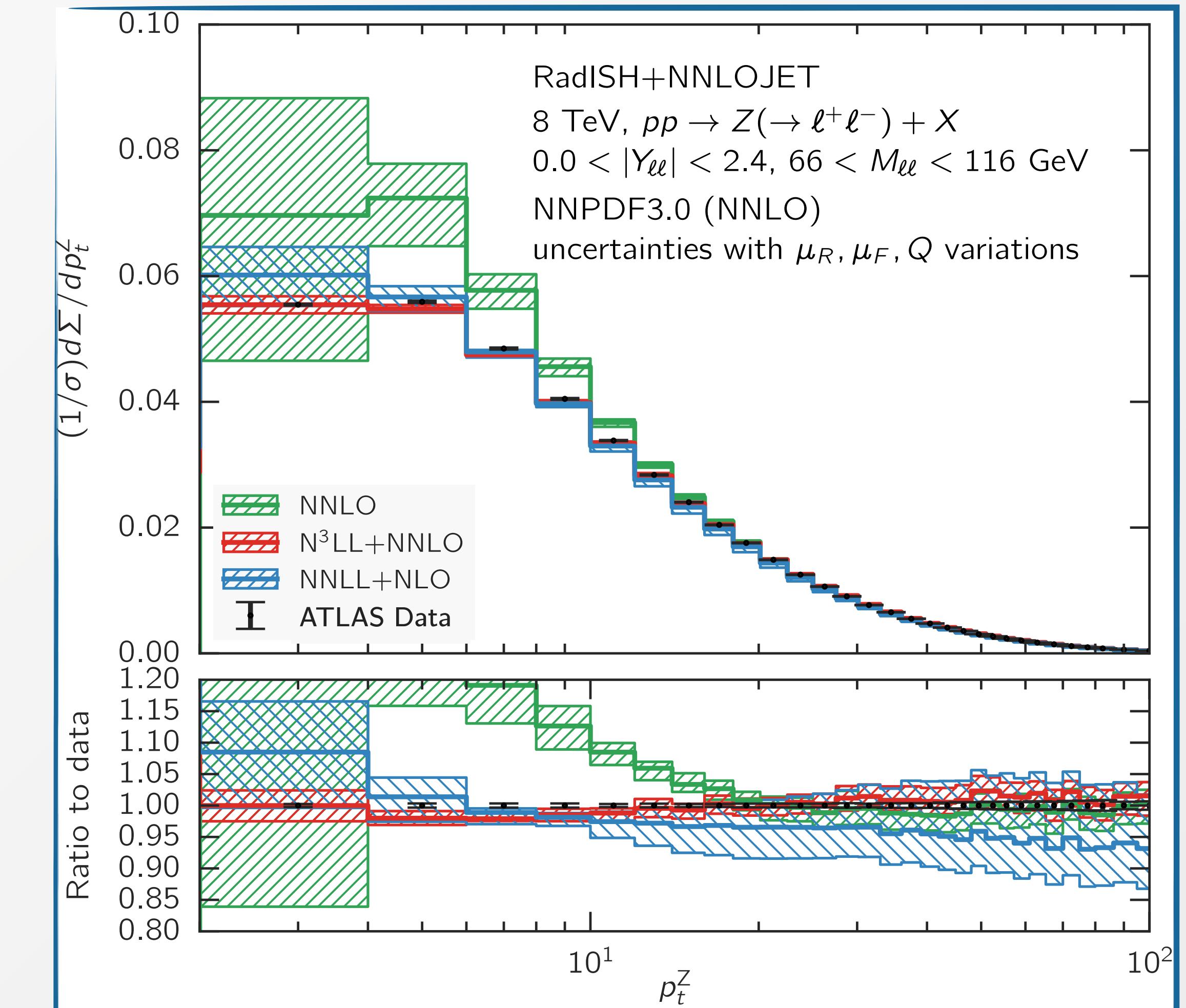
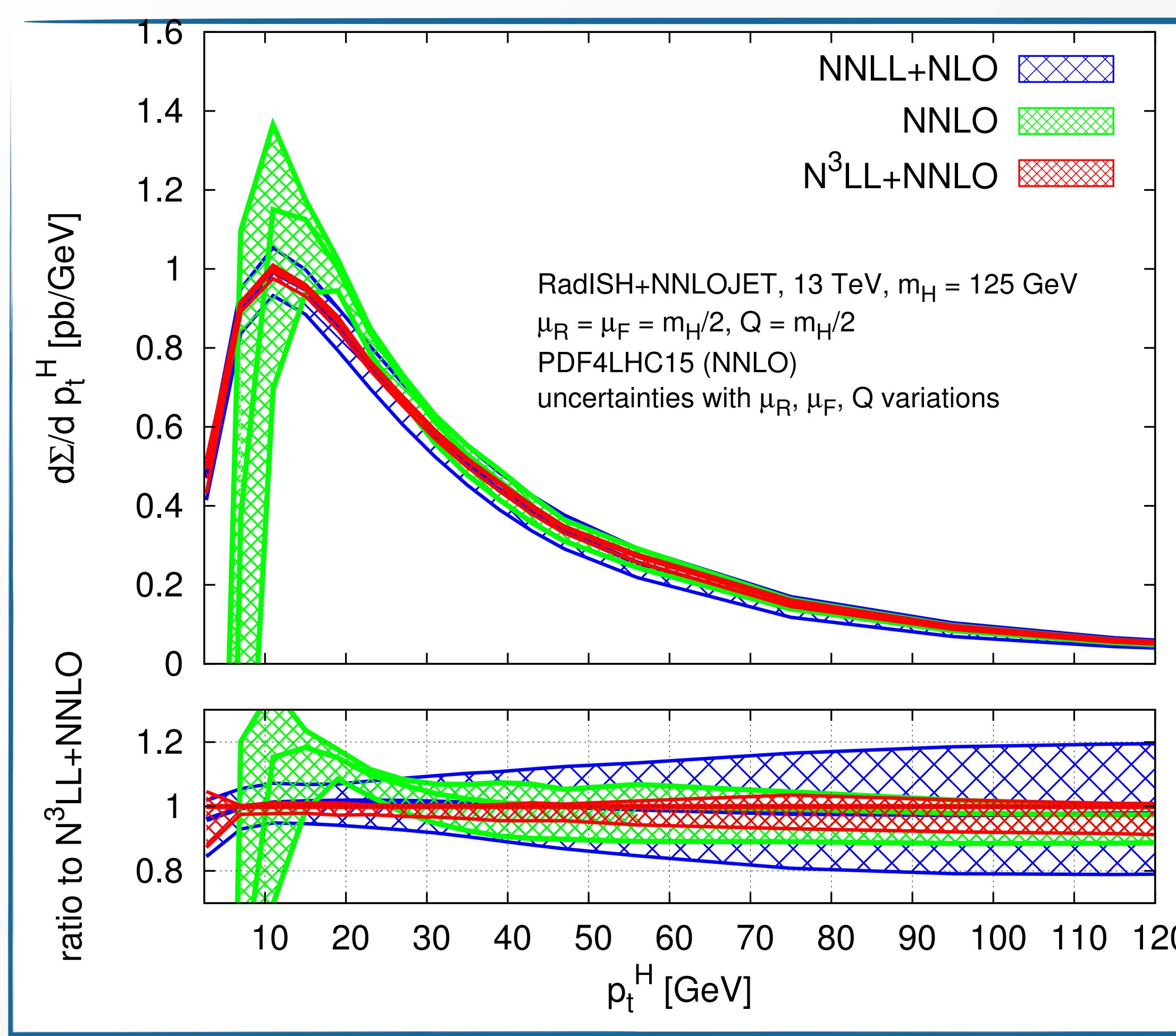
Predictions for the Z spectrum at 8 TeV



- Good description of the data in all fiducial regions
- Perturbative uncertainty at the few percent level, still does not match the precision of the ATLAS data

Resummation of the transverse momentum spectrum at N³LL+NNLO

N³LL result matched to NNLO H+j, Z+j, W[±]+j [Bizon, LR et al. '18, '19]



Theoretical predictions for Z and W observables at 13 TeV

Bizon, Gehrmann-De Ridder, Gehrmann, Glover, Huss, Monni, Re, LR, Walker, 190x.xxxx

Results obtained using the following fiducial cuts (agreed with ATLAS)

$$p_t^{\ell^\pm} > 25 \text{ GeV}, \quad |\eta^{\ell^\pm}| < 2.5, \quad 66 \text{ GeV} < M_{\ell\ell} < 116 \text{ GeV}$$

$$p_t^\ell > 25 \text{ GeV}, \quad |\eta^\ell| < 2.5, \quad E_T^{\nu_\ell} > 25 \text{ GeV}, \quad m_T > 50 \text{ GeV}$$

using NNPDF3.1 with $\alpha_s(M_Z)=0.118$ and setting the central scales to

$$\mu_R = \mu_F = M_T = \sqrt{M_{\ell\ell'}^2 + p_T^2}, \quad Q = \frac{M_{\ell\ell'}}{2}$$

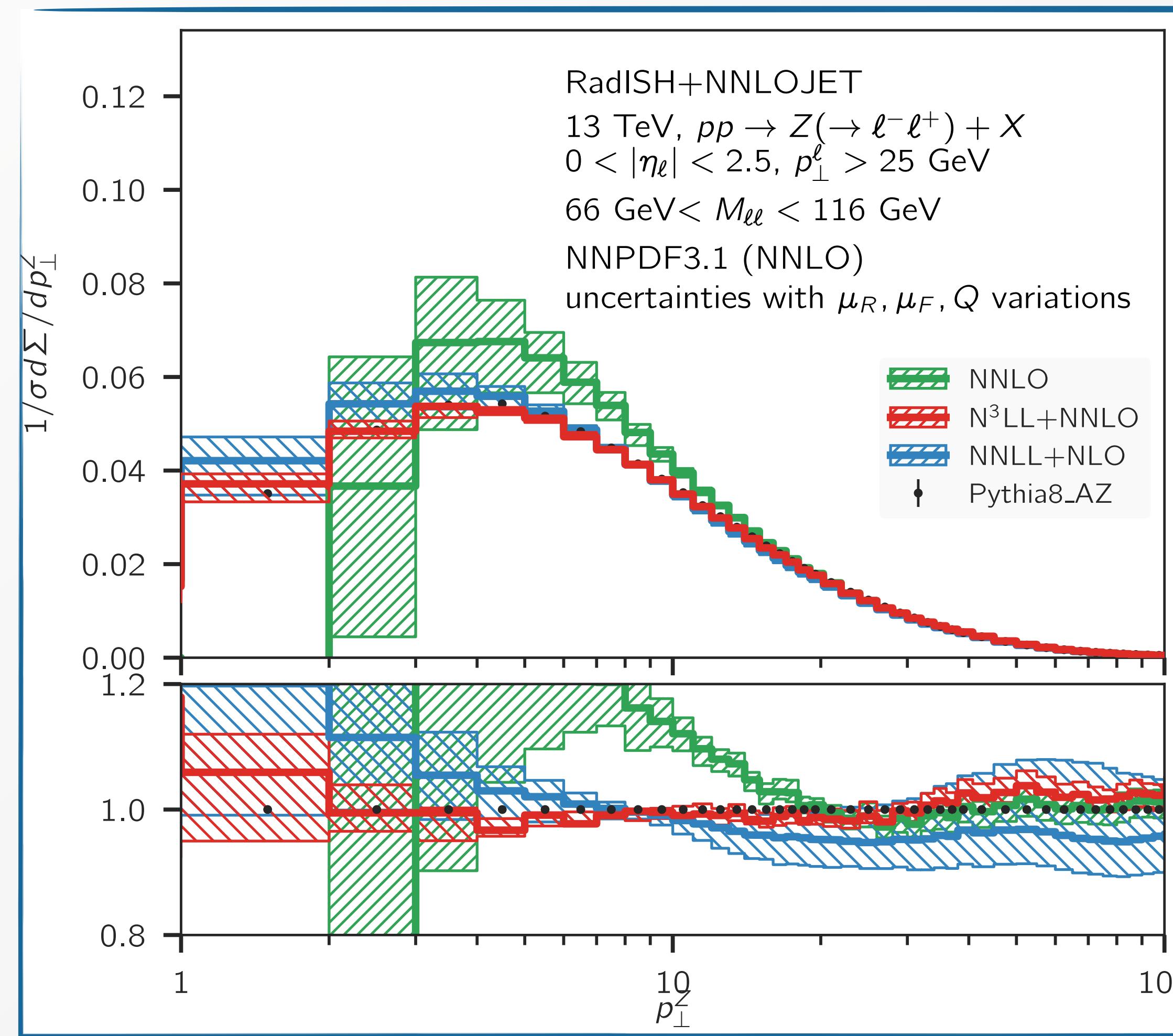
5 flavour (massless) scheme: no HQ effects, LHAPDF PDF thresholds

Scale uncertainties estimated by varying **renormalization** and **factorization** scale by a factor of two around their central value (**7 point variation**) and varying the **resummation** scale by a factor of 2 around its central value for factorization and renormalization scales set to their central value: **9 point envelope**

Matching parameter p set to 4 as a default

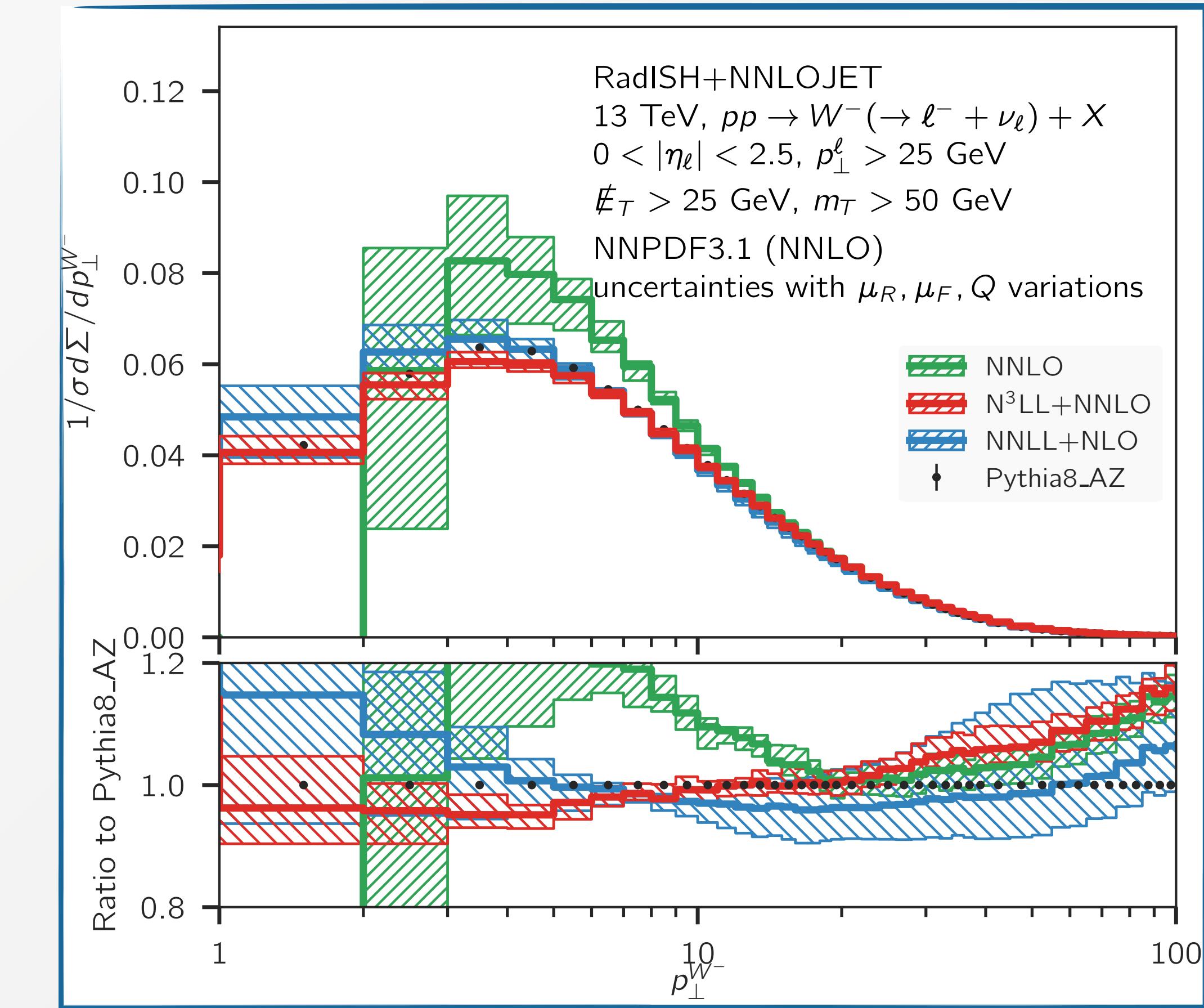
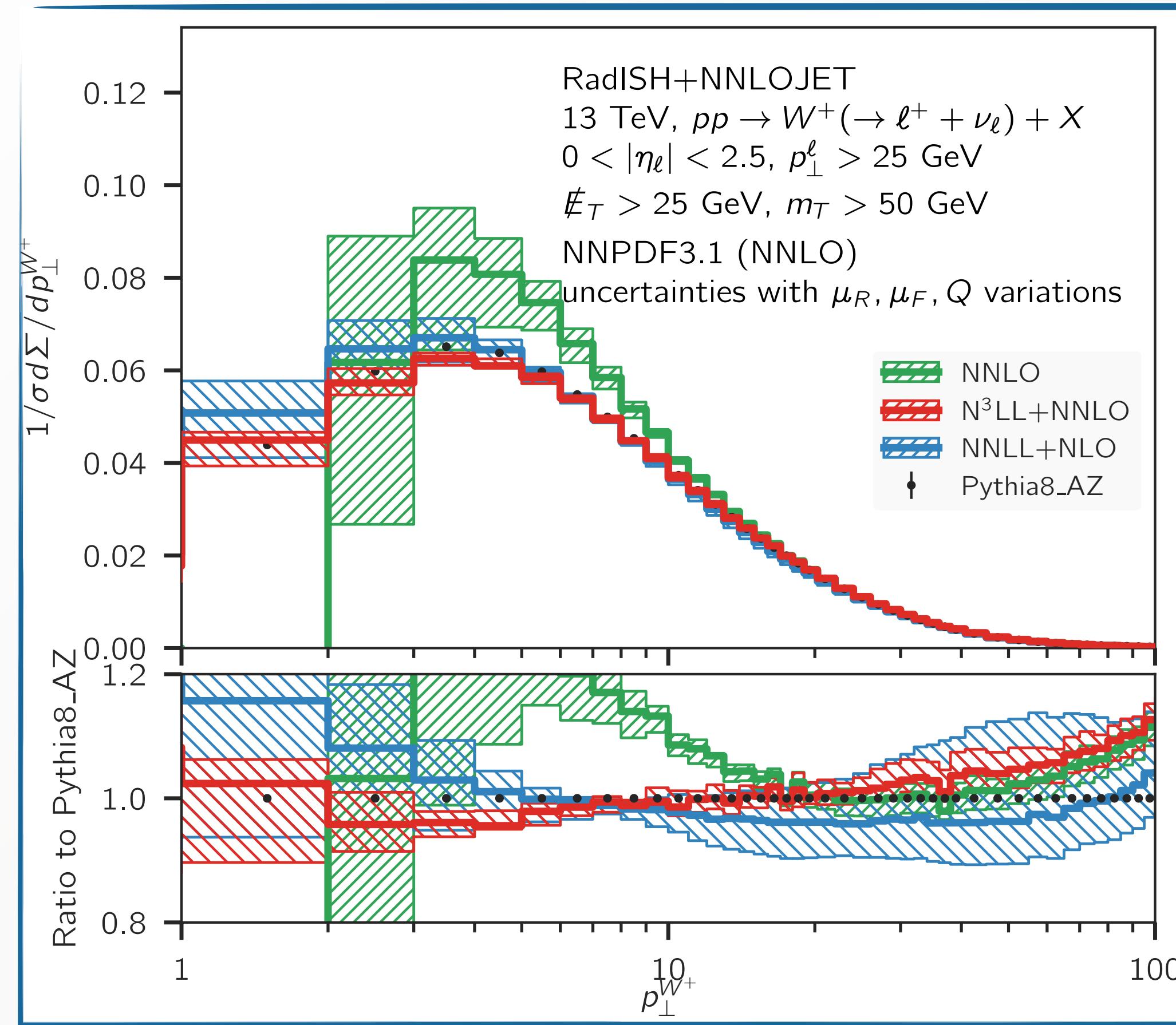
No non perturbative parameters included in the following

Predictions for the Z spectrum



Thanks to Jan Kretzschmar for providing the PYTHIA8 AZ tune results

Predictions for the W^+ and W^- spectra



Ratio of differential distributions

Z and W production share a similar pattern of QCD radiative corrections

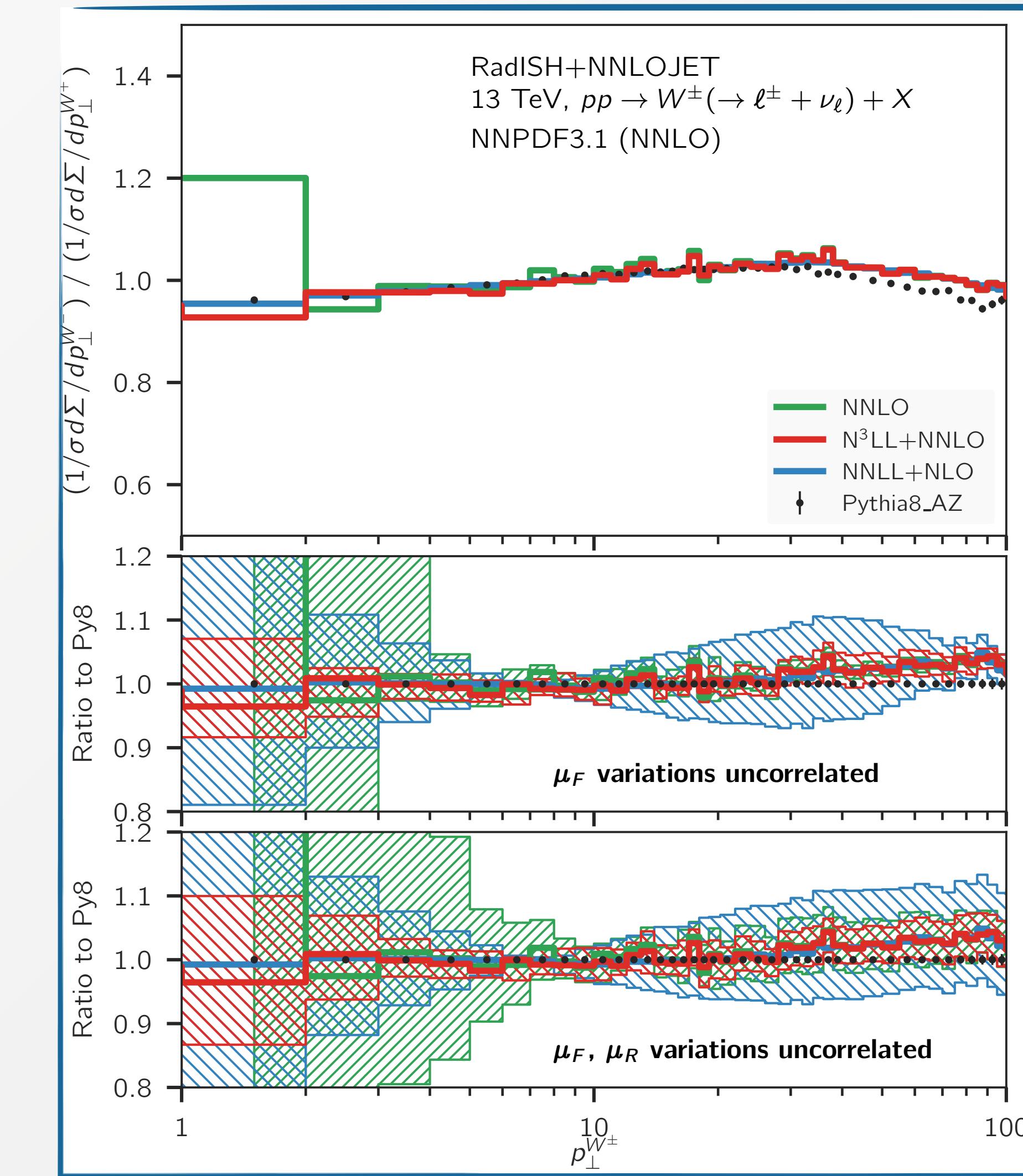
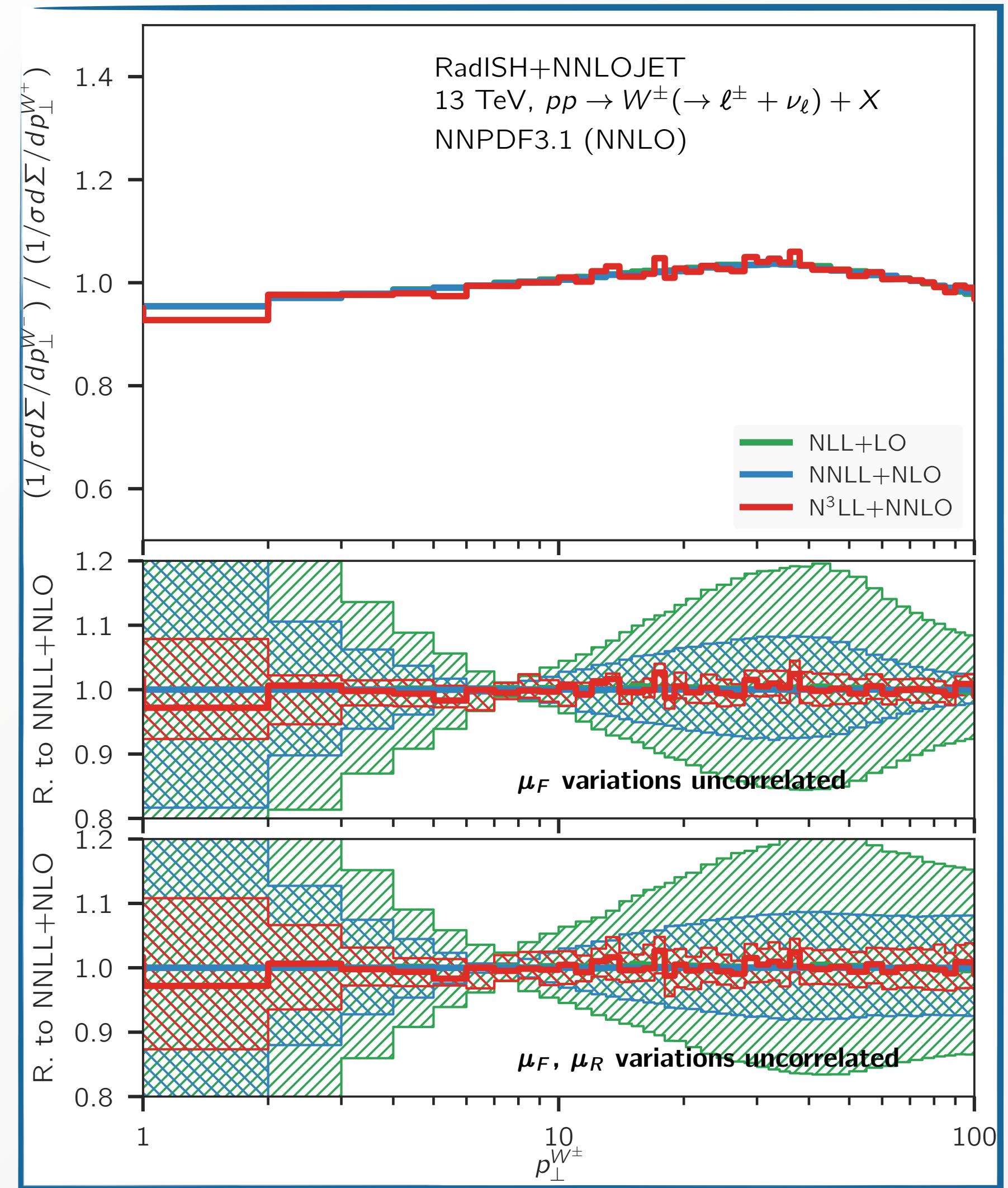
Crucial to understand correlation between Z and W spectra to exploit data-driven predictions

$$\frac{1}{\sigma^W} \frac{d\sigma^W}{p_\perp^W} \sim \frac{1}{\sigma_{\text{data}}^Z} \frac{d\sigma_{\text{data}}^Z}{p_\perp^Z} \frac{\frac{1}{\sigma_{\text{theory}}^W} \frac{d\sigma_{\text{theory}}^W}{p_\perp^W}}{\frac{1}{\sigma_{\text{theory}}^Z} \frac{d\sigma_{\text{theory}}^Z}{p_\perp^Z}}$$

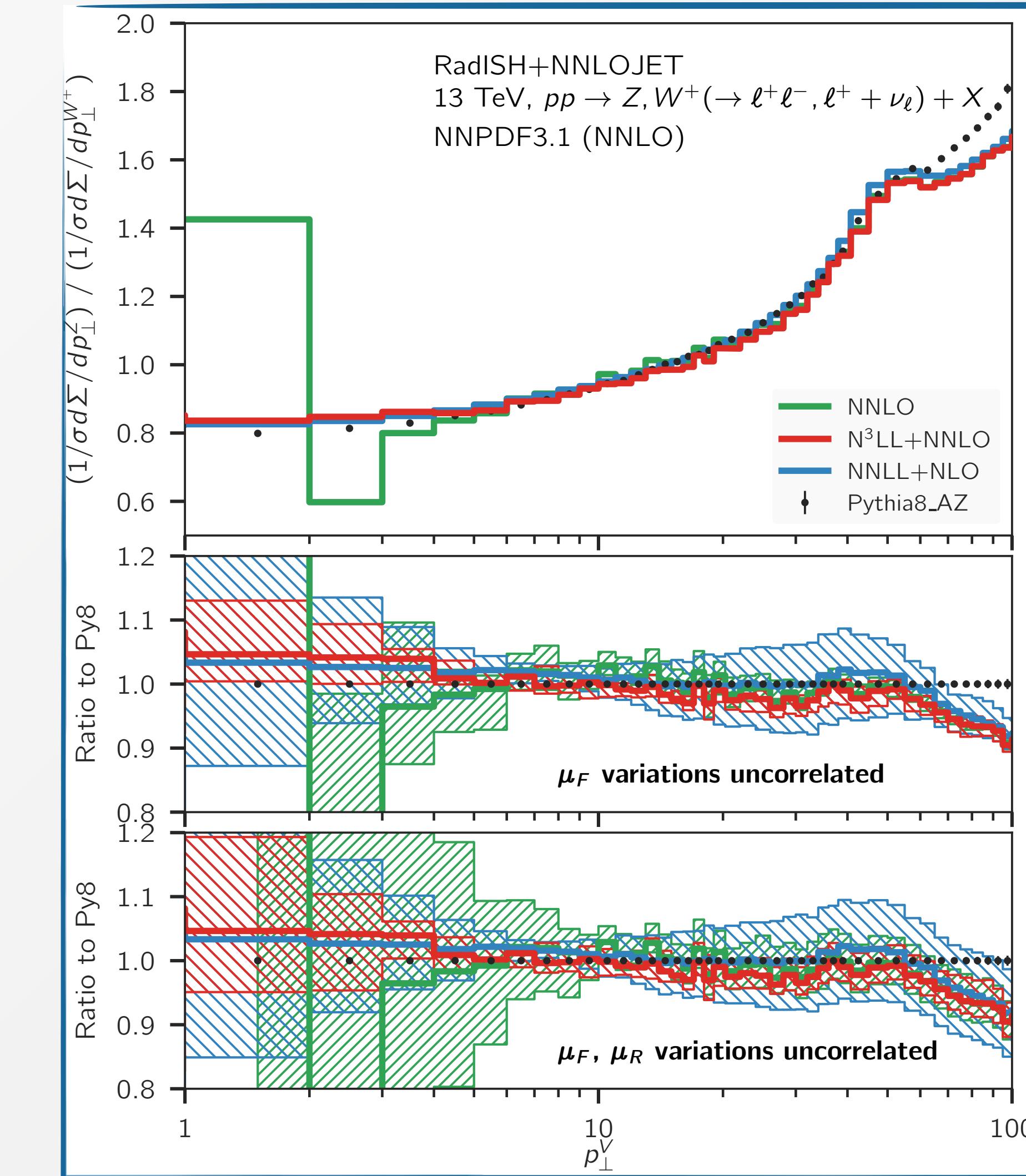
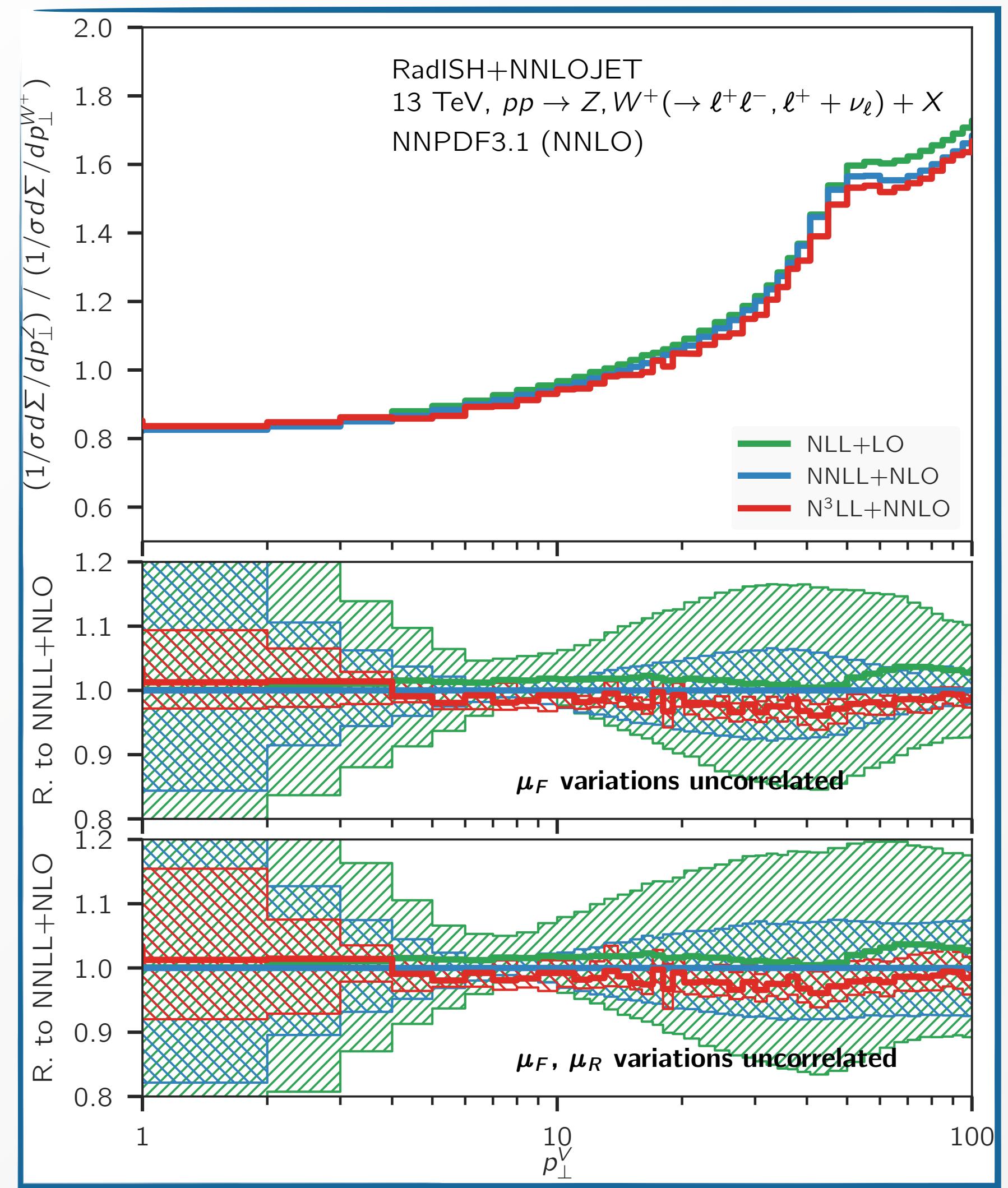
Several choices are possible:

- **Correlate resummation and renormalisation scale variations, keep factorisation scale uncorrelated**, while keeping
$$\frac{1}{2} \leq \frac{\mu_F^{\text{num}}}{\mu_F^{\text{den}}} \leq 2$$
- More **conservative** estimate: vary both **renormalisation and factorisation scales in an uncorrelated way** with
$$\frac{1}{2} \leq \frac{\mu^{\text{num}}}{\mu^{\text{den}}} \leq 2$$

Results for W/W^+ ratio



Results for Z/W^+ ratio



Equivalence with b -space formulation

$$\frac{d\Sigma(v)}{d\Phi_B} = \int_{\mathcal{C}_1} \frac{dN_1}{2\pi i} \int_{\mathcal{C}_2} \frac{dN_2}{2\pi i} x_1^{-N_1} x_2^{-N_2} \sum_{c_1, c_2} \frac{d|M_B|_{c_1 c_2}^2}{d\Phi_B} \mathbf{f}_{N_1}^T(\mu_0) \hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v) \mathbf{f}_{N_2}(\mu_0)$$

**unresolved
emission + virtual
corrections**

Result valid for
all inclusive
observables (e.g.
 p_t, φ^*)

**resolved
emission**

$$\begin{aligned} \hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v) &= \left[\mathbf{C}_{N_1}^{c_1; T}(\alpha_s(\mu_0)) H(\mu_R) \mathbf{C}_{N_2}^{c_2}(\alpha_s(\mu_0)) \right] \int_0^M \frac{dk_{t1}}{k_{t1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi} \\ &\times e^{-\mathbf{R}(\epsilon k_{t1})} \exp \left\{ - \sum_{\ell=1}^2 \left(\int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \frac{\alpha_s(k_t)}{\pi} \Gamma_{N_\ell}(\alpha_s(k_t)) + \int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \Gamma_{N_\ell}^{(C)}(\alpha_s(k_t)) \right) \right\} \\ &\quad \sum_{\ell_1=1}^2 \left(\mathbf{R}'_{\ell_1}(k_{t1}) + \frac{\alpha_s(k_{t1})}{\pi} \Gamma_{N_{\ell_1}}(\alpha_s(k_{t1})) + \Gamma_{N_{\ell_1}}^{(C)}(\alpha_s(k_{t1})) \right) \\ &\quad \times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} \sum_{\ell_i=1}^2 \left(\mathbf{R}'_{\ell_i}(k_{ti}) + \frac{\alpha_s(k_{ti})}{\pi} \Gamma_{N_{\ell_i}}(\alpha_s(k_{ti})) + \Gamma_{N_{\ell_i}}^{(C)}(\alpha_s(k_{ti})) \right) \\ &\quad \times \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \end{aligned}$$

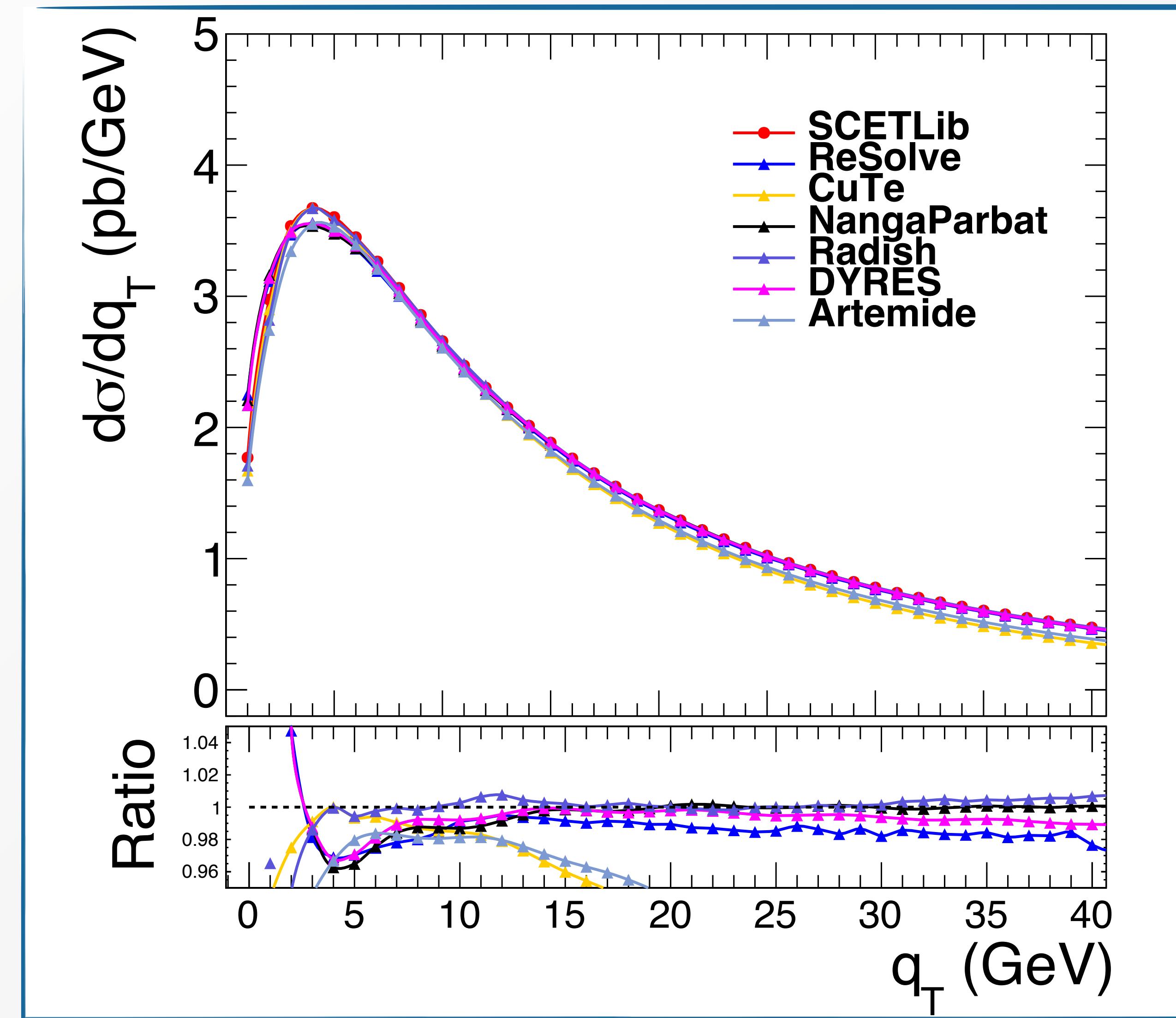
Formulation **equivalent to b -space result** (up to a **scheme change** in the anomalous dimensions)

$$\begin{aligned} \frac{d^2\Sigma(v)}{d\Phi_B dp_t} &= \sum_{c_1, c_2} \frac{d|M_B|_{c_1 c_2}^2}{d\Phi_B} \int b db p_t J_0(p_t b) \mathbf{f}^T(b/b) \mathbf{C}_{N_1}^{c_1; T}(\alpha_s(b/b)) H(M) \mathbf{C}_{N_2}^{c_2}(\alpha_s(b/b)) \mathbf{f}(b/b) \\ &\quad \times \exp \left\{ - \sum_{\ell=1}^2 \int_0^M \frac{dk_t}{k_t} \mathbf{R}'_\ell(k_t) (1 - J_0(bk_t)) \right\} \end{aligned}$$

$$(1 - J_0(bk_t)) \simeq \Theta(k_t - \frac{b_0}{b}) + \frac{\zeta_3}{12} \frac{\partial^3}{\partial \ln(Mb/b_0)^3} \Theta(k_t - \frac{b_0}{b})$$

**N³LL effect: absorbed in the definition
of H_2, B_3, A_4 coefficients wrt to CSS**

Equivalence with *b*-space formulation



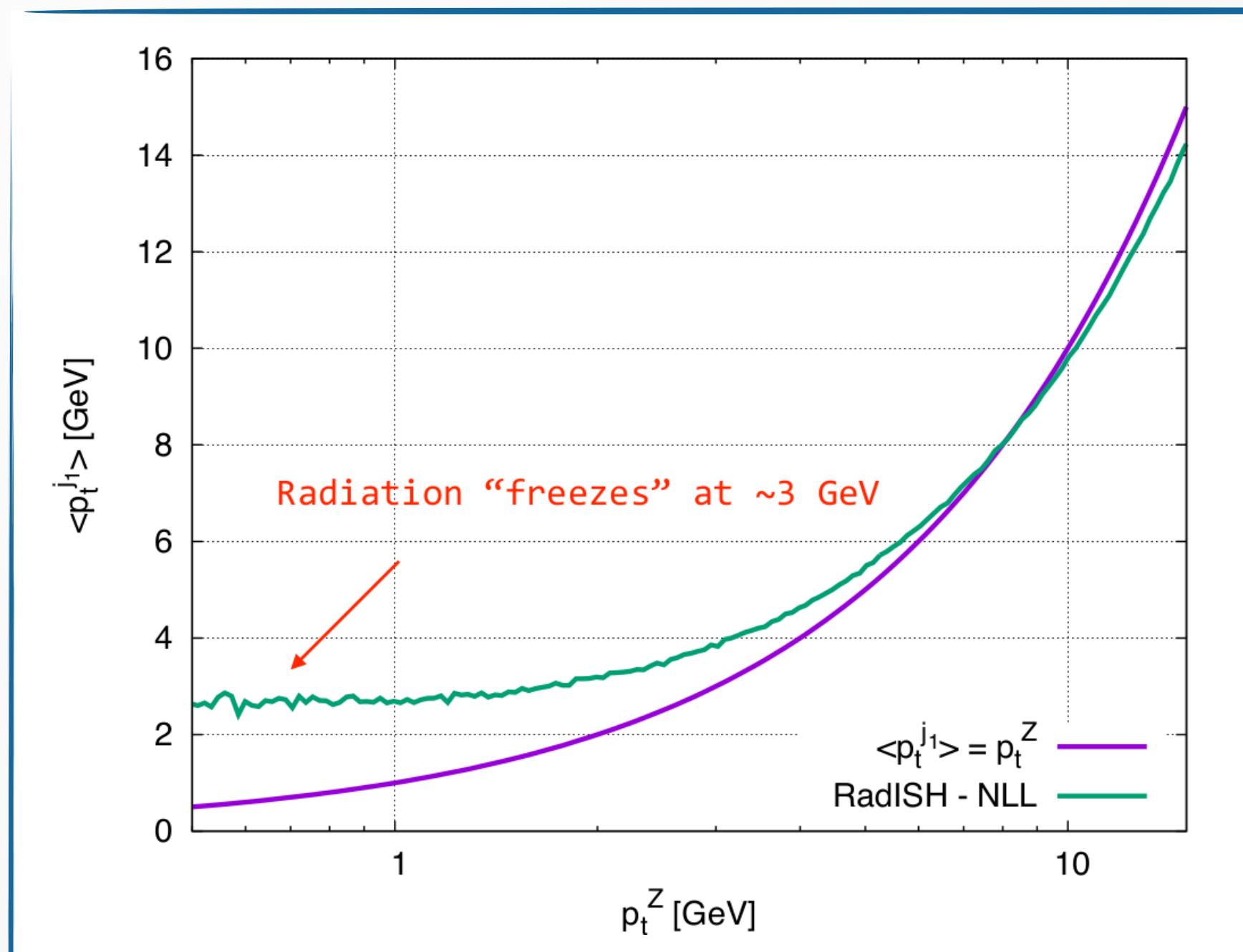
The Landau pole and the small p_T limit

Running coupling $\alpha_s(k_{t1}^2)$ and Sudakov radiator hit Landau pole at

$$\alpha_s(\mu_R^2)\beta_0 \ln Q/k_{t1} = \frac{1}{2}$$

$$k_{t1} \sim 0.01 \text{ GeV}, \quad \mu_R = Q = m_Z$$

Only real cutoff in the calculation: emission probability is set to zero below this scale and parton densities are frozen.



At small p_t the large azimuthal cancellations dominate over the Sudakov suppression: the cutoff is never an issue in practice

$$\frac{d^2\Sigma(v)}{dp_t d\Phi_B} \simeq 2\sigma^{(0)}(\Phi_B)p_t \left(\frac{\Lambda_{\text{QCD}}^2}{M^2} \right)^{\frac{16}{25} \ln \frac{41}{16}}$$

Thanks to P. Monni

Online LHCP 2020, Paris, 29 May 2020

Behaviour at small p_t

Explicit evaluation shows that the Parisi-Petronzio perturbative scaling at small p_t is reproduced. At NLL, Drell-Yan pair production, $n_f=4$

$$\frac{d^2\Sigma(v)}{dp_t d\Phi_B} = 4 \sigma^{(0)}(\Phi_B) p_t \int_{\Lambda_{\text{QCD}}}^M \frac{dk_{t1}}{k_{t1}^3} e^{-R(k_{t1})} \simeq 2\sigma^{(0)}(\Phi_B) p_t \left(\frac{\Lambda_{\text{QCD}}^2}{M^2} \right)^{\frac{16}{25} \ln \frac{41}{16}}$$

As now higher logarithmic terms (up to N³LL) are under control, the coefficient of this scaling can be systematically improved in *perturbation theory* (non-perturbative effects – of the same order – not considered)

N³LL calculation allows one to have control over the terms of relative order $O(\alpha_s^2)$. Scaling $L \sim 1/\alpha_s$ valid in the deep infrared regime.

Numerical implementation

$$\begin{aligned} \frac{d\Sigma(p_t)}{d\Phi_B} &= \int_0^M \frac{dk_{t1}}{k_{t1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi} \partial_L \left(-e^{-R'(k_{t1})} \mathcal{L}_{\text{NLL}}(k_{t1}) \right) \times \\ &\quad \times \underbrace{\epsilon^{R'(k_{t1})} \sum_{n=0}^{\infty} \frac{1}{n!} \left(\prod_{i=2}^{n+1} \int_{\epsilon k_{t1}}^{k_{t1}} \frac{dk_{ti}}{k_{ti}} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R'(k_{ti}) \right) \Theta(p_t - |\vec{k}_{t1} + \dots + \vec{k}_{t(n+1)}|)}_{\equiv \int d\mathcal{Z}[\{R', k_i\}] \Theta(p_t - |\vec{k}_{t1} + \dots + \vec{k}_{t(n+1)}|)}. \end{aligned}$$

- $L = \ln(M/k_{t1})$; luminosity $\mathcal{L}_{\text{NLL}}(k_{t1}) = \sum_{c_1, c_2} \frac{d|M_B|^2_{c_1 c_2}}{d\Phi_B} f_{c_1}(x_1, k_{t1}) f_{c_2}(x_2, k_{t1})$.
- $\int d\mathcal{Z}[\{R', k_i\}] \Theta$ finite as $\epsilon \rightarrow 0$:

$$\begin{aligned} \epsilon^{R'(k_{t1})} &= 1 - R'(k_{t1}) \ln(1/\epsilon) + \dots = 1 - \int_{\epsilon k_{t1}}^{k_{t1}} R'(k_{t1}) + \dots, \\ \int d\mathcal{Z}[\{R', k_i\}] \Theta &= \left[1 - \int_{\epsilon k_{t1}}^{k_{t1}} R'(k_{t1}) + \dots \right] \left[\Theta(p_t - |\vec{k}_{t1}|) + \int_{\epsilon k_{t1}}^{k_{t1}} R'(k_{t1}) \Theta(p_t - |\vec{k}_{t1} + \vec{k}_{t2}|) + \dots \right] \\ &= \Theta(p_t - |\vec{k}_{t1}|) + \underbrace{\int_0^{k_{t1}} R'(k_{t1})}_{\epsilon \rightarrow 0} \underbrace{\left[\Theta(p_t - |\vec{k}_{t1} + \vec{k}_{t2}|) - \Theta(p_t - |\vec{k}_{t1}|) \right]}_{\text{finite: real-virtual cancellation}} + \dots \end{aligned}$$

- Evaluated with Monte Carlo techniques: $\int d\mathcal{Z}[\{R', k_i\}]$ is generated as a parton shower over secondary emissions.

Numerical implementation

- ▶ Secondary radiation:

$$\begin{aligned} d\mathcal{Z}[\{R', k_i\}] &= \sum_{n=0}^{\infty} \frac{1}{n!} \left(\prod_{i=2}^{n+1} \int_0^{2\pi} \frac{d\phi_i}{2\pi} \int_{\epsilon k_{t1}}^{k_{t1}} \frac{dk_{ti}}{k_{ti}} R'(k_{t1}) \right) \epsilon^{R'(k_{t1})} \\ &= \sum_{n=0}^{\infty} \left(\prod_{i=2}^{n+1} \int_0^{2\pi} \frac{d\phi_i}{2\pi} \int_{\epsilon k_{t1}}^{k_{t(i-1)}} \frac{dk_{ti}}{k_{ti}} R'(k_{t1}) \right) \epsilon^{R'(k_{t1})}, \\ \epsilon^{R'(k_{t1})} &= e^{-R'(k_{t1}) \ln 1/\epsilon} = \prod_{i=2}^{n+2} e^{-R'(k_{t1}) \ln k_{t(i-1)}/k_{ti}}, \end{aligned}$$

with $k_{t(n+2)} = \epsilon k_{t1}$.

- ▶ Each secondary emissions has differential probability

$$dw_i = \frac{d\phi_i}{2\pi} \frac{dk_{ti}}{k_{ti}} R'(k_{t1}) e^{-R'(k_{t1}) \ln k_{t(i-1)}/k_{ti}} = \frac{d\phi_i}{2\pi} d \left(e^{-R'(k_{t1}) \ln k_{t(i-1)}/k_{ti}} \right).$$

- ▶ $k_{t(i-1)} \geq k_{ti}$. Scale k_{ti} extracted by solving $e^{-R'(k_{t1}) \ln k_{t(i-1)}/k_{ti}} = r$, with r random number extracted uniformly in $[0, 1]$. **Shower ordered in k_{ti}** .
- ▶ Extract ϕ_i randomly in $[0, 2\pi]$.

Joint resummation in direct space

$$\begin{aligned}
\sigma_{\text{incl}}^{\text{NNLL}}(p_t^{\text{J},\text{v}}, p_t^{\text{H},\text{v}}) = & \int_0^{p_t^{\text{J},\text{v}}} \frac{dk_{t,1}}{k_{t,1}} \frac{d\phi_1}{2\pi} \int d\mathcal{Z} \left\{ \frac{d}{dL_{t,1}} \left[-e^{-R_{\text{NNLL}}(L_{t,1})} \mathcal{L}_{\text{NNLL}}(\mu_F e^{-L_{t,1}}) \right] \Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1}|) \right. \\
& + e^{-R_{\text{NLL}}(L_{t,1})} \hat{R}'(k_{t,1}) \int_0^{k_{t,1}} \frac{dk_{t,s_1}}{k_{t,s_1}} \frac{d\phi_{s_1}}{2\pi} \left[\left(\delta \hat{R}'(k_{t,1}) + \hat{R}''(k_{t,1}) \ln \frac{k_{t,1}}{k_{t,s_1}} \right) \mathcal{L}_{\text{NLL}}(\mu_F e^{-L_{t,1}}) - \frac{d}{dL_{t,1}} \mathcal{L}_{\text{NLL}}(\mu_F e^{-L_{t,1}}) \right] \\
& \times \left. \left[\Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1} + \vec{k}_{t,s_1}|) - \Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1}|) \right] \right\}, \tag{38}
\end{aligned}$$

$$\begin{aligned}
\sigma_{\text{clust}}^{\text{NNLL}}(p_t^{\text{J},\text{v}}, p_t^{\text{H},\text{v}}) = & \int_0^\infty \frac{dk_{t,1}}{k_{t,1}} \frac{d\phi_1}{2\pi} \int d\mathcal{Z} e^{-R_{\text{NLL}}(L_{t,1})} \mathcal{L}_{\text{NLL}}(\mu_F e^{-L_{t,1}}) 8 C_A^2 \frac{\alpha_s^2}{\pi^2} \frac{L_{t,1}}{(1 - 2\beta_0 \alpha_s L_{t,1})^2} \Theta(p_t^{\text{J},\text{v}} - \max_{i>1} \{k_{t,i}\}) \\
& \times \left\{ \int_0^{k_{t,1}} \frac{dk_{t,s_1}}{k_{t,s_1}} \frac{d\phi_{s_1}}{2\pi} \int_{-\infty}^\infty d\Delta\eta_{1s_1} J_{1s_1}(R) \left[\Theta(p_t^{\text{J},\text{v}} - |\vec{k}_{t,1} + \vec{k}_{t,s_1}|) - \Theta(p_t^{\text{J},\text{v}} - k_{t,1}) \right] \Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1} + \vec{k}_{t,s_1}|) \right. \\
& + \frac{1}{2!} \hat{R}'(k_{t,1}) \int_0^{k_{t,1}} \frac{dk_{t,s_1}}{k_{t,s_1}} \frac{dk_{t,s_2}}{k_{t,s_2}} \frac{d\phi_{s_1}}{2\pi} \frac{d\phi_{s_2}}{2\pi} \int_{-\infty}^\infty d\Delta\eta_{s_1s_2} J_{s_1s_2}(R) \left[\Theta(p_t^{\text{J},\text{v}} - |\vec{k}_{t,s_1} + \vec{k}_{t,s_2}|) - \Theta(p_t^{\text{J},\text{v}} - \max\{k_{t,s_1}, k_{t,s_2}\}) \right] \\
& \times \left. \Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1} + \vec{k}_{t,s_1} + \vec{k}_{t,s_2}|) \Theta(p_t^{\text{J},\text{v}} - k_{t,1}) \right\}, \tag{42}
\end{aligned}$$

$$\begin{aligned}
\sigma_{\text{correl}}^{\text{NNLL}}(p_t^{\text{J},\text{v}}, p_t^{\text{H},\text{v}}) = & \int_0^\infty \frac{dk_{t,1}}{k_{t,1}} \frac{d\phi_1}{2\pi} \int d\mathcal{Z} e^{-R_{\text{NLL}}(L_{t,1})} \mathcal{L}_{\text{NLL}}(\mu_F e^{-L_{t,1}}) 8 C_A^2 \frac{\alpha_s^2}{\pi^2} \frac{L_{t,1}}{(1 - 2\beta_0 \alpha_s L_{t,1})^2} \Theta(p_t^{\text{J},\text{v}} - \max_{i>1} \{k_{t,i}\}) \\
& \times \left\{ \int_0^{k_{t,1}} \frac{dk_{t,s_1}}{k_{t,s_1}} \frac{d\phi_{s_1}}{2\pi} \int_{-\infty}^\infty d\Delta\eta_{1s_1} \mathcal{C} \left(\Delta\eta_{1s_1}, \Delta\phi_{1s_1}, \frac{k_{t,1}}{k_{t,s_1}} \right) (1 - J_{1s_1}(R)) \right. \\
& \times \left. \left[\Theta(p_t^{\text{J},\text{v}} - k_{t,1}) - \Theta(p_t^{\text{J},\text{v}} - |\vec{k}_{t,1} + \vec{k}_{t,s_1}|) \right] \Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1} + \vec{k}_{t,s_1}|) \right. \\
& + \frac{1}{2!} \hat{R}'(k_{t,1}) \int_0^{k_{t,1}} \frac{dk_{t,s_1}}{k_{t,s_1}} \frac{dk_{t,s_2}}{k_{t,s_2}} \frac{d\phi_{s_1}}{2\pi} \frac{d\phi_{s_2}}{2\pi} \int_{-\infty}^\infty d\Delta\eta_{s_1s_2} \mathcal{C} \left(\Delta\eta_{s_1s_2}, \Delta\phi_{s_1s_2}, \frac{k_{t,s_2}}{k_{t,s_1}} \right) (1 - J_{s_1s_2}(R)) \Theta(p_t^{\text{J},\text{v}} - k_{t,1}) \\
& \times \left. \left[\Theta(p_t^{\text{J},\text{v}} - \max\{k_{t,s_1}, k_{t,s_2}\}) - \Theta(p_t^{\text{J},\text{v}} - |\vec{k}_{t,s_1} + \vec{k}_{t,s_2}|) \right] \Theta(p_t^{\text{H},\text{v}} - |\vec{k}_{t,1} + \dots + \vec{k}_{t,n+1} + \vec{k}_{t,s_1} + \vec{k}_{t,s_2}|) \right\}. \tag{43}
\end{aligned}$$

All-order resummation: CAESAR/ARES approach

Solution 2:

Translate the resummability into properties of the observable in the presence of multiple radiation:
recursive infrared and collinear (rIRC) safety

[Banfi, Salam, Zanderighi '01, '03, '04]

[Banfi, McAslan, Monni, Zanderighi, El-Menoufi '14, '18]

Simple observable easy to calculate

$$\tilde{\sigma} \sim \int \frac{d\nu_1}{\nu_1} \boxed{\Sigma_s(\nu_1)} \boxed{\mathcal{F}(\nu, \nu_1)}$$

Transfer function relates the resummation of the full observable to the one of the simple observable.

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Transfer function relates the resummation of the full observable to the one of the simple observable.
i.e. conditional probability

Separation obtained by introducing a **resolution scale** $q_0 = \epsilon k_{t,1}$

$$\tilde{\sigma} \sim \int [dk_1] e^{-R(q_0)}$$

Unresolved emission can be treated as **totally unconstrained**
→ **exponentiation**

$$\times |\mathcal{M}(k_1)|^2 \left(\sum_{m=0}^{\infty} \frac{1}{m!} \int \prod_{i=2}^{m+1} [dk_i] |\mathcal{M}(k_i)|^2 \Theta(V(k_i) - q_0) \Theta(v - V(k_1, \dots, k_{m+1})) \right)$$

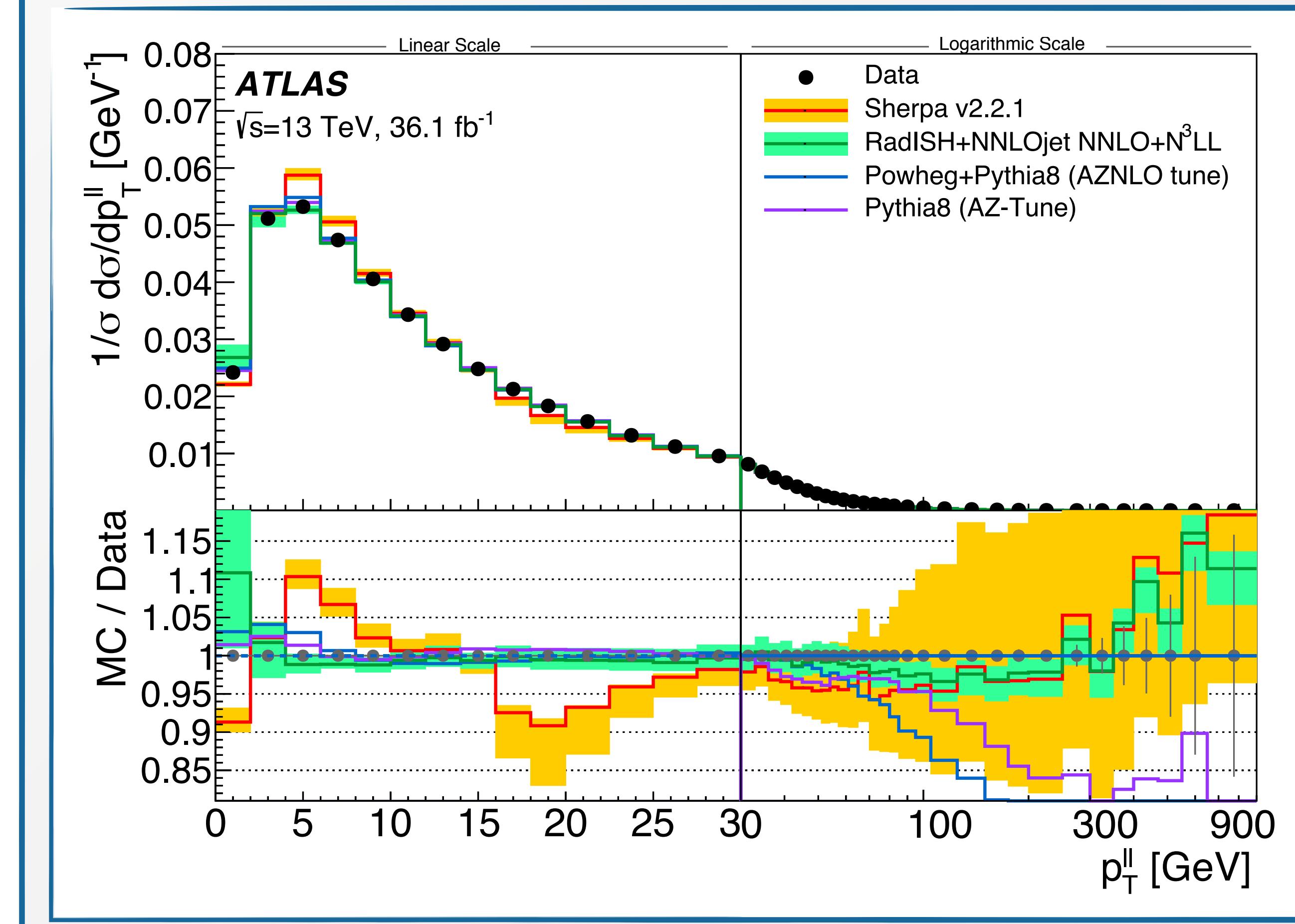
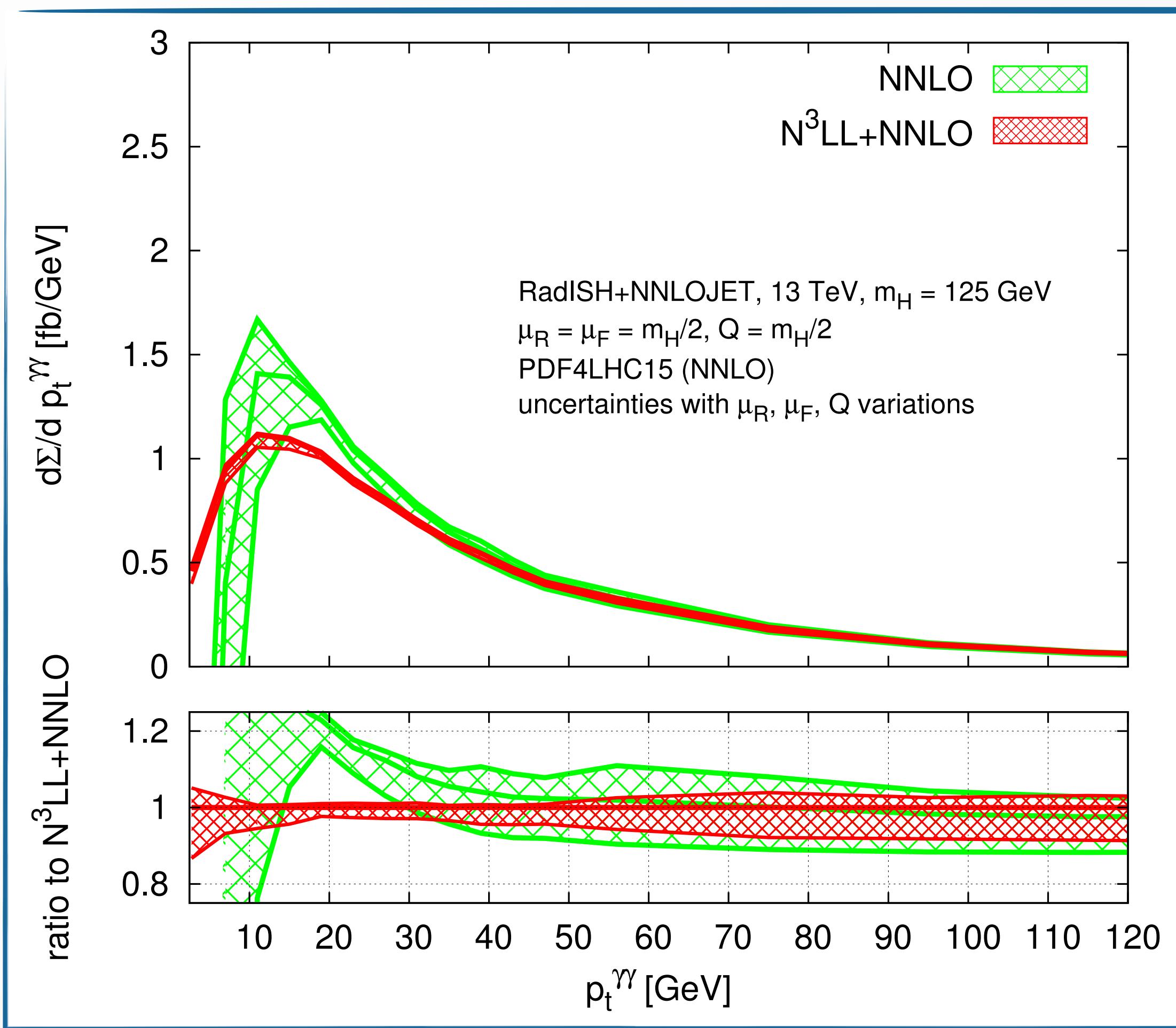
Resolved emission treated exclusively with **Monte Carlo methods**. Integral is finite, can be integrated in d=4 with a computer

Approach recently formulated within SCET language [Bauer, Monni '18, '19 + ongoing work]

Method entirely formulated in **direct space**

Resummation of the transverse momentum spectrum at N³LL+NNLO

N³LL result matched to NNLO H+j, Z+j, W[±]+j [Bizon, LR et al. '17, '18, '19]



[ATLAS 1912.02844]

H+j at same accuracy also in SCET [Chen et al. '18]

Resummation of the transverse momentum spectrum in b space

Solution 1:

move to **conjugate space** where phase space factorization is manifest

p_\perp resummation [Parisi, Petronzio '79; Collins, Soper, Sterman '85]

$$\delta^{(2)}\left(\vec{p}_t - \sum_{i=1}^n \vec{k}_{t,i}\right) = \int d^2 b \frac{1}{4\pi^2} e^{i \vec{b} \cdot \vec{p}_t} \prod_{i=1}^n e^{-i \vec{b} \cdot \vec{k}_{t,i}}$$

two-dimensional momentum conservation

Exponentiation in conjugate space

$$\sigma = \sigma_0 \int d^2 \vec{p}_\perp^H \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i \vec{b} \cdot \vec{p}_\perp^H} \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=1}^n \int [dk_i] |M(k_i)|^2$$

virtual corrections

$$\left(e^{i \vec{b} \cdot \vec{k}_{t,i}} - 1 \right) = \sigma_0 \int d^2 \vec{p}_\perp^H \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i \vec{b} \cdot \vec{p}_\perp^H} e^{-R_{\text{NLL}}(L)}$$

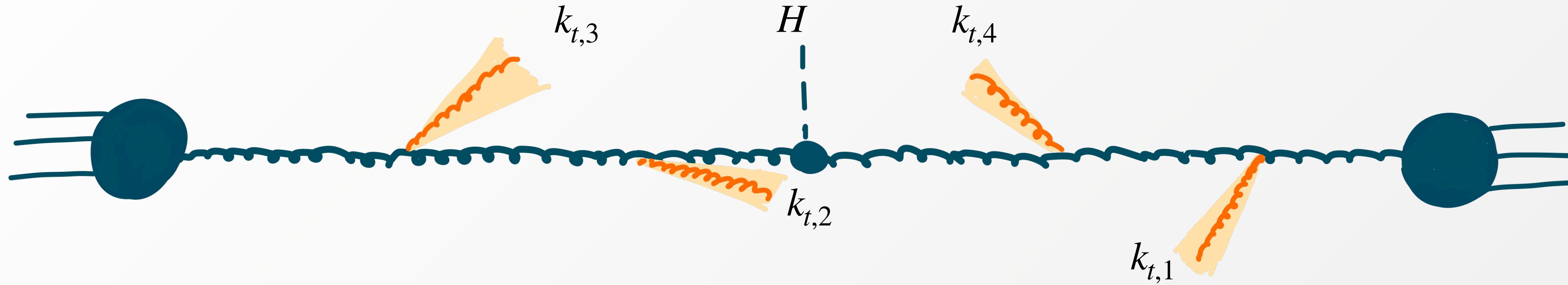
NLL formula with scale-independent PDFs

$$R_{\text{NLL}}(L) = -L g_1(\alpha_s L) - g_2(\alpha_s L)$$

$$L = \ln(m_H b / b_0)$$

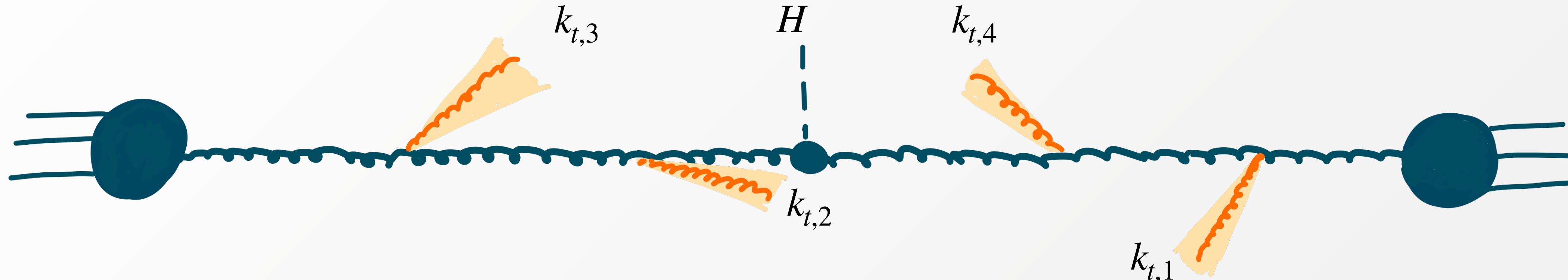
Double-differential resummation at NLL in b space

At NLL, emissions are **strongly ordered** in angle. Clustering algorithms will associate **each emission** to a **different** jet



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Additional constraint on **real radiation**

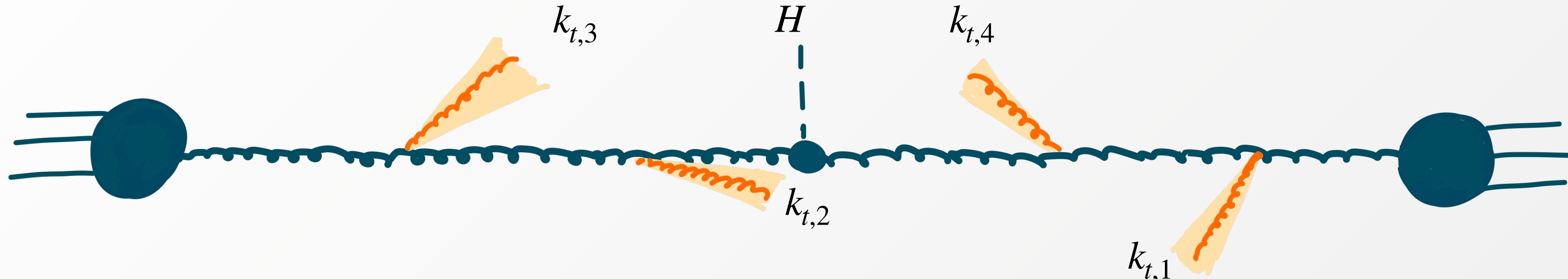
$$\Theta(p_{\perp}^{\text{J,v}} - \max\{k_{t,1}, \dots, k_{t,n}\}) = \prod_{i=1}^n \Theta(p_{\perp}^{\text{J,v}} - k_{t,i})$$

p_{\perp}^H resummation formula

$$\frac{d\sigma}{d^2\vec{p}_{\perp}^H} = \sigma_0 \int \frac{d^2\vec{b}}{4\pi^2} e^{-i\vec{b}\cdot\vec{p}_{\perp}^H} e^{-R_{\text{NLL}}(L)}$$

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Joint $p_\perp^H, p_\perp^{\text{J,v}}$ resummation formula

$$\frac{d\sigma(p_\perp^{\text{J,v}})}{d^2\vec{p}_\perp^H} = \sigma_0 \int \frac{d^2\vec{b}}{4\pi^2} e^{-i\vec{b}\cdot\vec{p}_\perp^H} e^{-S_{\text{NLL}}(L)}$$

CMW scheme [Catani, Marchesini, Webber '91]

$$S_{\text{NLL}}(L) = -Lg_1(\alpha_s L) - g_2(\alpha_s L) + \int_0^{m_H} \frac{dk_t}{k_t} R'_{\text{NLL}}(k_t) J_0(bk_t) \Theta(k_t - p_\perp^{\text{J,v}})$$

$$R'_{\text{NLL}}(k_t) = 4 \left(\frac{\alpha_s^{\text{CMW}}(k_t)}{\pi} C_A \ln \frac{m_H}{k_t} - \alpha_s(k_t) \beta_0 \right)$$

Double-differential resummation at NNLL in b space

Crucial observation: in b space the phase space constraints entirely factorize  $e^{i\vec{b}\cdot\vec{k}_{t,i}}$

The jet veto constraint can be included by implementing the jet veto resummation at the b -space integrand level
directly in impact-parameter space

Inclusive contribution: phase space constraint of the form

$$\Theta(p_{\perp}^{\text{J,v}} - \max\{\mathbf{k}_{t,1}, \dots, \mathbf{k}_{t,n}\}) = \prod_{i=1}^n \Theta(p_{\perp}^{\text{J,v}} - \mathbf{k}_{t,i})$$

Promote radiator at NNLL

$$S_{\text{NNLL}} = -Lg_1(\alpha_s L) - g_2(\alpha_s L) - \alpha_s g_3(\alpha_s L) + \int_0^{m_H} \frac{dk_t}{k_t} R'_{\text{NNLL}}(k_t) J_0(bk_t) \Theta(k_t - p_{\perp}^{\text{J,v}})$$

Double-differential resummation at NNLL in b space

Resummation formula must be amended at NNLL [Banfi et al. '12][Becher et al. '13][Stewart et al. '14]

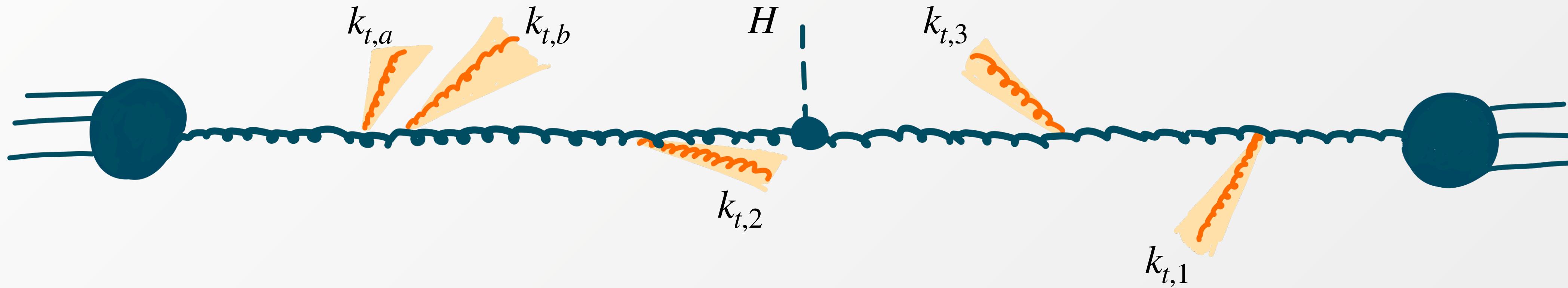
$$\frac{d\sigma(p_{\perp}^{\text{J,v}})}{d^2 \vec{p}_{\perp}^H} = \sigma_0 \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i \vec{b} \cdot \vec{p}_{\perp}^H} e^{-S_{\text{NNLL}}(L)} (1 + \mathcal{F}_{\text{clust}} + \mathcal{F}_{\text{correl}})$$

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- **clustering correction:** jet algorithm can cluster two **independent** emissions into the same jet

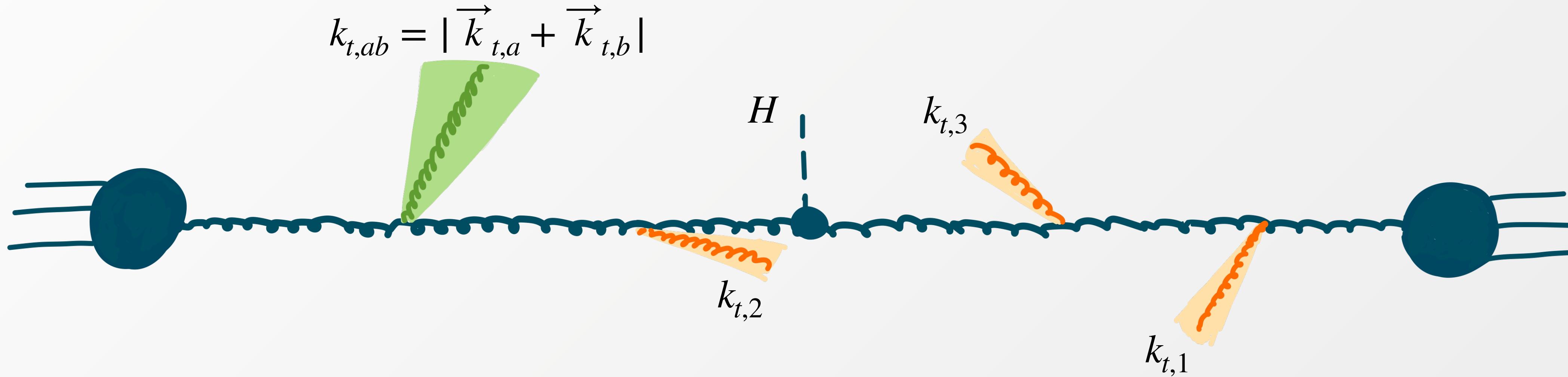


Double-differential resummation at NNLL in b space

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$$\frac{d\sigma(p_{\perp}^{\text{J,v}})}{d^2 \vec{p}_{\perp}^H} = \sigma_0 \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i \vec{b} \cdot \vec{p}_{\perp}^H} e^{-S_{\text{NNLL}}(L)} (1 + \mathcal{F}_{\text{clust}} + \mathcal{F}_{\text{correl}})$$

- **clustering correction:** jet algorithm can cluster two **independent** emissions into the same jet

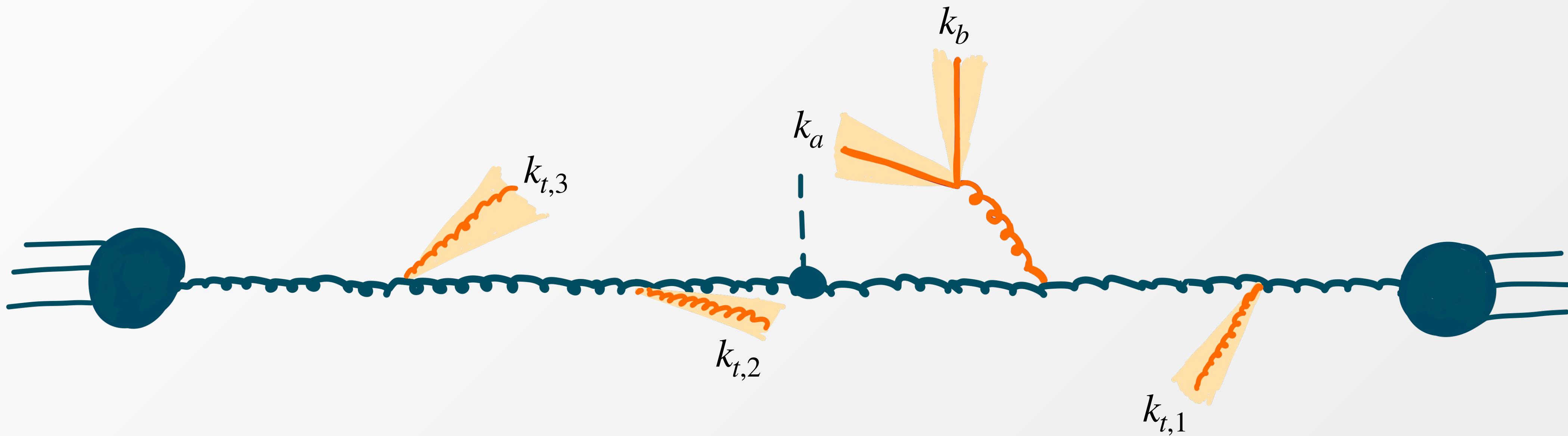


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- **correlated correction:** amends the inclusive treatment of the correlated squared amplitude for two emission accounting for configurations where the two **correlated** emissions (non abelian) are not clustered in the same jet

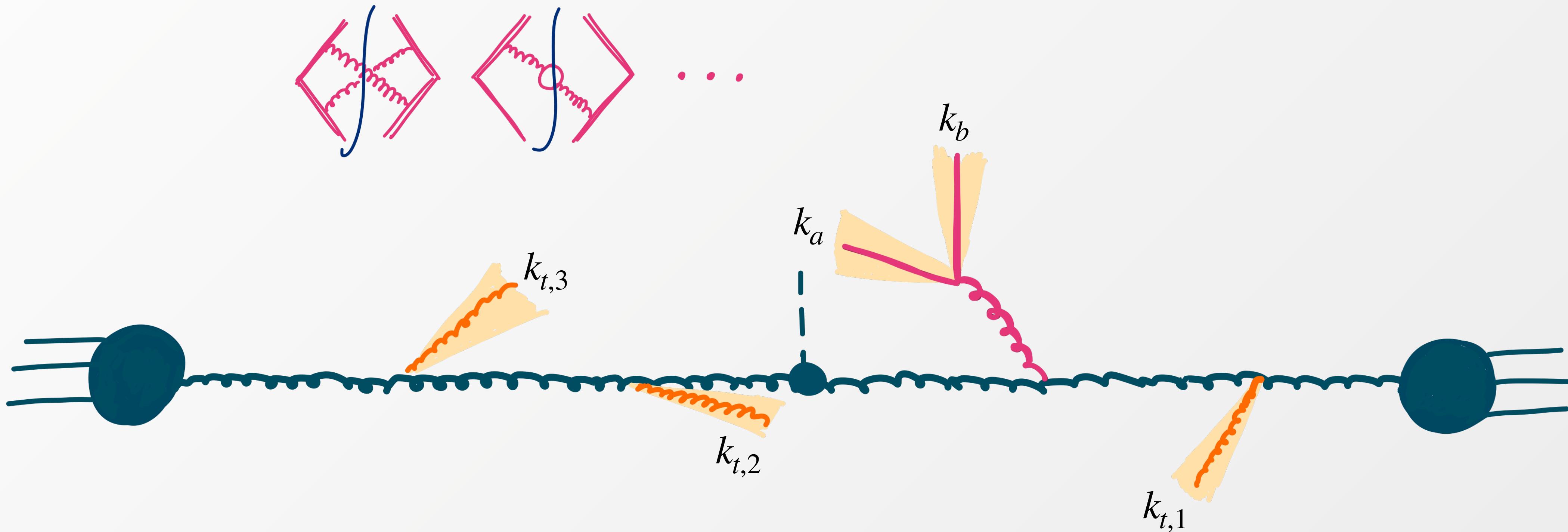


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Double-differential resummation at NNLL in b space

NNLL prediction finally requires the consistent treatment of non-soft collinear emissions off the initial state particles

Soft and non-soft emission cannot be clustered by a k_t -type jet algorithm. Non-soft collinear radiation can be handled by taking a Mellin transform of the resummed cross section, giving rise to **scale evolution of PDFs** and of the $\mathcal{O}(\alpha_s)$ **collinear coefficient functions**

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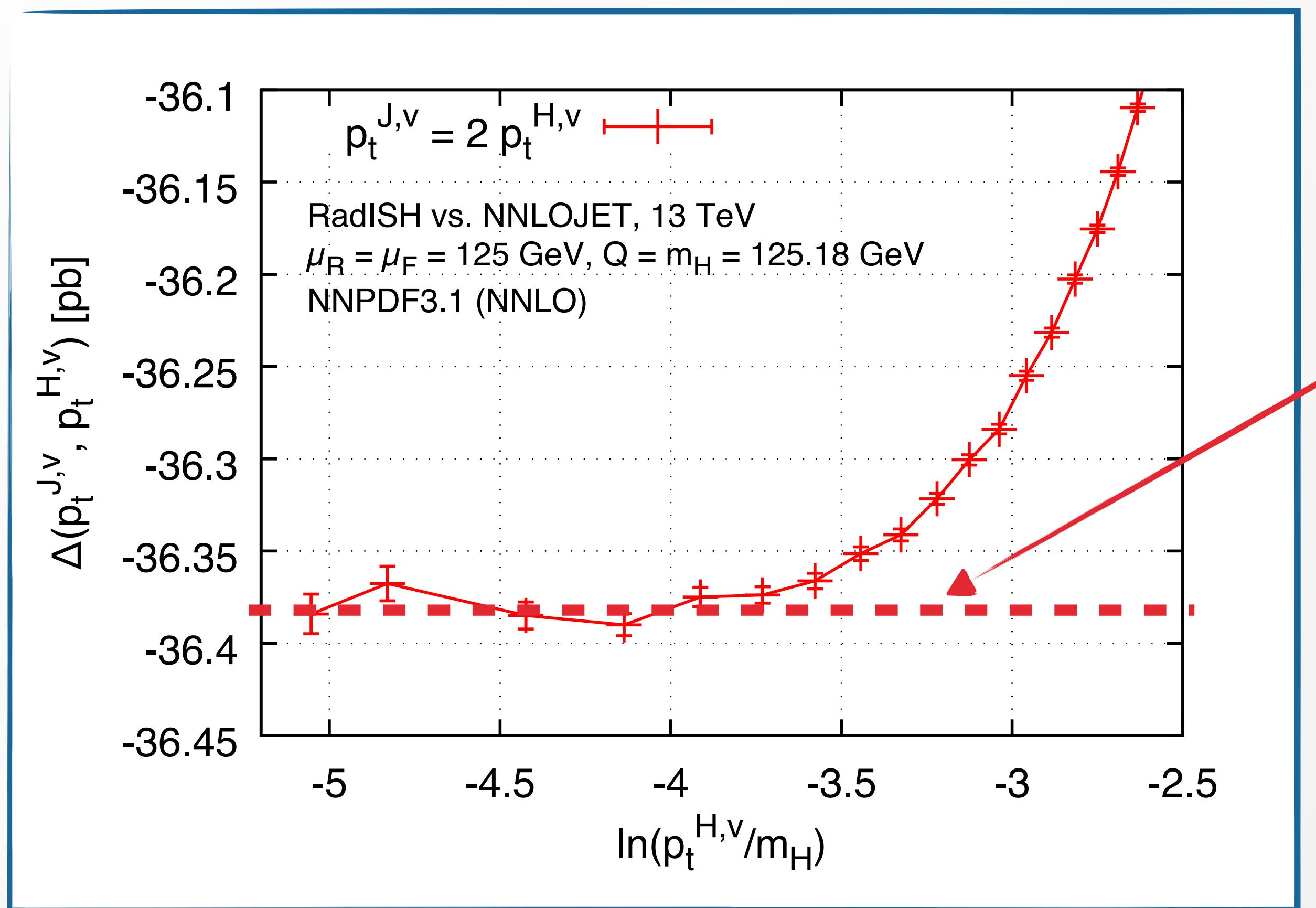
Final result at NNLL, including **hard-virtual corrections** at $\mathcal{O}(\alpha_s)$

$$\frac{d\sigma(p_{\perp}^{\text{J,v}})}{dy_H d^2 \vec{p}_{\perp}^H} = M_{\text{gg} \rightarrow H}^2 \mathcal{H}(\alpha_s(m_H)) \int_{\mathcal{C}_1} \frac{d\nu_1}{2\pi i} \int_{\mathcal{C}_2} \frac{d\nu_2}{2\pi i} x_1^{-\nu_1} x_2^{-\nu_2} \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i \vec{b} \cdot \vec{p}_{\perp}^H} e^{-S_{\text{NNLL}}} (1 + \mathcal{F}_{\text{clust}} + \mathcal{F}_{\text{correl}})$$
$$\times f_{\nu_1, a_1}(b_0/b) f_{\nu_2, a_2}(b_0/b) \left[\mathcal{P} e^{-\int_{p_{\perp}^{\text{J,v}}}^{m_H} \frac{d\mu}{\mu} \Gamma_{\nu_1}(\alpha_s(\mu)) J_0(b\mu)} \right]_{c_1 a_1} \left[\mathcal{P} e^{-\int_{p_{\perp}^{\text{J,v}}}^{m_H} \frac{d\mu}{\mu} \Gamma_{\nu_2}(\alpha_s(\mu)) J_0(b\mu)} \right]_{c_2 a_2}$$
$$\times C_{\nu_1, g b_1}(\alpha_s(b_0/b)) C_{\nu_2, g b_2}(\alpha_s(b_0/b)) \left[\mathcal{P} e^{-\int_{p_{\perp}^{\text{J,v}}}^{m_H} \frac{d\mu}{\mu} \Gamma_{\nu_1}^{(C)}(\alpha_s(\mu)) J_0(b\mu)} \right]_{c_1 b_1} \left[\mathcal{P} e^{-\int_{p_{\perp}^{\text{J,v}}}^{m_H} \frac{d\mu}{\mu} \Gamma_{\nu_2}^{(C)}(\alpha_s(\mu)) J_0(b\mu)} \right]_{c_2 b_2}$$

Mellin moments

Flavour indices

Accuracy check at $\mathcal{O}(\alpha_s^2)$



$$\Delta(p_\perp^{J,v}, p_\perp^{H,v}) = \sigma^{\text{NNLO}}(p_\perp^{J,v}, p_\perp^{H,v}) - \sigma_{\text{exp.}}^{\text{NNLL}}(p_\perp^{J,v}, p_\perp^{H,v})$$

$$\sigma^{\text{NNLO}}(p_\perp^H < p_\perp^{H,v}, p_\perp^J < p_\perp^{J,v}) = \sigma^{\text{NNLO}} - \int \Theta(p_\perp^H > p_\perp^{H,v}) \vee \Theta(p_\perp^J > p_\perp^{J,v}) d\sigma_{H+J}^{\text{NLO}}$$

LHC results: Higgs transverse momentum with a jet veto

