# Transverse observable resummation at the LHC 

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## After Higgs



- The Higgs discovery in 2012 completed the Standard Model (SM) puzzle
- We do know that the picture is not yet complete: there are various phenomena which call for physics beyond the SM (neutrino masses, dark matter, baryon asymmetry...)
- Naturalness: since $m_{H}=125 \mathrm{GeV}$, to avoid fine-tuning scale of new physics $(\mathrm{NP}) \Lambda_{\mathrm{NP}}$ should be $O(\mathrm{TeV})$
- Ongoing direct and indirect searches for NP signatures at the LHC


## BSM searches at the LHC

[ATLAS-CONF-2016-018]


## Bump hunting

- Little or no theoretical input
- Reconstructed signal over a smooth and well understood background
- Besides Higgs, not very
[Alioli et al '17]


Deviations from SM predictions

- Look for deviations from SM
- Need an accurate theoretical description of the kinematic distributions


## Quest for precision

If the scale for NP $\Lambda_{N P}$ is a few $T e V$, expected deviations from the SM behaviour are


## LHC in the precision era

$1-5 \%$ level of precision is within reach at the (HL)-LHC


- Luminosity reached $100 \mathrm{fb}^{-1}$ at 13 TeV
- Increase in statistics enables study of differential distributions in detail
- Measurements at \% level (or even smaller) are available for several processes

Astonishing level of precision reached in e.g. Z transverse momentum: luminosity and other systematics are cancelled or reduced if results are normalized by fiducial cross section

Very accurate theoretical predictions needed

## Precision physics at the LHC: theory



Parton Distribution Functions (PDFs)
Long-distance, non-perturbative, universal objects

## Precision physics at the LHC: theory

Key concept: collinear factorization
$\sqrt{s}$ centre-of-mass energy
$X \quad Q \quad$ energy scale of the process

$$
\sigma\left(s, Q^{2}\right)=\sum_{a, b} \int d x_{1} d x_{2} f_{a / h_{1}}\left(x_{1}, Q^{2}\right) f_{b / h_{2}}\left(x_{2}, Q^{2}\right) \hat{\sigma}_{a b \rightarrow X}\left(Q^{2}, x_{1} x_{2} s\right)+\mathcal{O}\left(\Lambda_{\mathrm{QCD}}^{p} / Q^{p}\right)
$$

Hard-scattering matrix element Short-distance, perturbative, process-dependent

## Precision physics at the LHC: theory

$$
\begin{aligned}
& \sigma\left(s, Q^{2}\right)=\sum_{a, b} \int d x_{1} d x_{2} f_{a / h_{1}}\left(x_{1}, Q^{2}\right) f_{b / h_{2}}\left(x_{2}, Q^{2}\right) \hat{\sigma}_{a b \rightarrow X}\left(Q^{2}, x_{1} x_{2} s\right)+\mathcal{O}\left(\Lambda_{\mathrm{QCD}}^{p} / Q^{p}\right) \\
& \text { Input parameters: } \\
& \text { strong coupling } \\
& \text { PDFs } \\
& \text { few percent } \\
& \text { uncertainty; } \\
& \text { improvable } \\
& \text { Non-perturbative effects } \\
& \text { percent effect; } \\
& \text { not yet under } \\
& \text { control }
\end{aligned}
$$

## Precision physics at the LHC: theory

$$
\begin{gathered}
\sigma\left(s, Q^{2}\right)=\sum_{a, b} \int d x_{1} d x_{2} f_{a / h_{1}}\left(x_{1}, Q^{2}\right) f_{b / h_{2}}\left(x_{2}, Q^{2}\right) \hat{\sigma}_{a b \rightarrow X}\left(Q^{2}, x_{1} x_{2} s\right)+\mathcal{O}\left(\Lambda_{\mathrm{QCD}}^{p} / Q^{p}\right) \\
\hat{\sigma}=\hat{\sigma}_{0}\left(1+\alpha_{s} C_{1}+\alpha_{s}^{2} C_{2}+\alpha_{s}^{3} C_{3}+\ldots\right)
\end{gathered}
$$

LO NLO NNLO N3LO

$$
\alpha_{s} \sim 0.1
$$

$$
\begin{aligned}
& \delta \sim 10-20 \% \\
& \delta \sim 1-5 \%
\end{aligned}
$$

NLO
NNLO (or even N3LO)

## Transverse observables

Particularly clean experimental and theoretical environment for precision physics

Parameterized as

$$
V(k)=\left(\frac{k_{t}}{M}\right)^{a} f(\phi)
$$

for a single soft QCD emission $k$ collinear to incoming leg. Independent of the rapidity of radiation. $V \rightarrow 0$ for soft/collinear radiation.

Inclusive observables (e.g. transverse momentum $p_{t}$ ) probe directly the kinematics of the colour singlet

- negligible sensitivity to multi-parton interactions
- reduced sensitivity to non-perturbative effects
- measured extremely precisely at experiments (sub-percent in $Z$ differential)

$$
V\left(k_{1}, \ldots k_{n}\right)=V\left(k_{1}+\ldots+k_{n}\right)
$$

## Transverse observables at the LHC

Implications for indirect constraints on BSM physics

[Bishara et al '16]
Bound on light Yukawa couplings

[Grazzini et al '16]

Sensitivity to dimension-6 operators

## Transverse observables at the LHC

Also important implications for extraction of SM parameters (strong coupling and PDF determination, $W$ mass measurements...)

[NNPDF '18]

[Boughezal et al '17]

## All-order resummation

Cumulative cross section

$$
\Sigma(v)=\int_{0}^{v} d V \frac{d \sigma}{d V} \sim \sigma_{0}\left[1+\alpha_{s}+\alpha_{s}^{2}+\ldots\right]
$$

Fixed-order prediction: reliable for inclusive enough observables and in regions not marred by soft/collinear radiation

Real and virtual contributions can become however highly unbalanced in processes where the real radiation is strongly constrained by kinematics

Large logarithms appear at all order as a left-over of the real-virtual cancellation of IRC divergences

$$
\begin{array}{ccc}
\ln \Sigma(v)=\sum_{n}\left\{\mathcal{O}\left(\alpha_{s}^{n} L^{n+1}\right)+\mathcal{O}\left(\alpha_{s}^{n} L^{n}\right)+\mathcal{O}\left(\alpha_{s}^{n} L^{n-1}\right)+\ldots\right\} & L=\ln R \\
\text { LL } & \text { NLL } & \text { NNLL }
\end{array} \begin{aligned}
& R \text { : ratio of typical scales } \\
& \\
&
\end{aligned}
$$

> Fixed order predictions no longer reliable:
> all-order resummation of the perturbative series

## All-order resummation

$$
\ln \Sigma(v)=\sum\left\{\mathcal{O}\left(\alpha_{s}^{n} L^{n+1}\right)+\mathcal{O}\left(\alpha_{s}^{n} L^{n}\right)+\mathcal{O}\left(\alpha_{s}^{n} L^{n-1}\right)+\ldots\right\}
$$


*It's the sum that makes the total

## Example: transverse momentum spectrum

System with high invariant mass $M \gg p_{t}$, where the transverse momentum $p_{t}$ vanishes at Born level

If $p_{t} \ll M$, the emission of real radiation is strongly suppressed. Double logarithms of $p_{t} / M$ appear as a leftover of the real/virtual cancellation at all orders and spoil the perturbative convergence at small $p_{t}$


Logarithms of $p_{t} / M$ must be resummed to reliably describe the small $\boldsymbol{p}_{\boldsymbol{t}}$ region

## Case study: transverse momentum $p_{t}$

Resummation of transverse momentum is particularly delicate because $p_{t}$ is a vectorial quantity

Two concurring mechanisms leading to a system with small $p_{\mathrm{t}}$


$$
p_{t}^{2} \sim k_{t, i}^{2} \ll M^{2}
$$

cross section naturally suppressed as there is no phase space left for gluon emission
(Sudakov limit)


$$
\sum_{i=1}^{n} \vec{k}_{k_{i}} \approx 0
$$

Large kinematic cancellations
$p_{t} \sim 0$ far from the Sudakov limit

> Power suppression

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## Dominant at small $p_{t}$

$$
\begin{aligned}
& \sum_{5} 5^{5} \\
& \sum_{i=1}^{n} \vec{k}_{t, i} \simeq 0
\end{aligned}
$$

Large kinematic cancellations
$p_{t} \sim 0$ far from the Sudakov limit

> Power suppression

## Resummation in conjugate space

Resummation usually performed in impact-parameter (b) space where the two competing mechanisms are handled trough a Fourier transform

$$
\delta\left(\vec{p}_{t}-\sum_{i=1}^{n} \vec{k}_{t, i}\right)=\int d^{2} b \frac{1}{4 \pi^{2}} e^{i \vec{b} \cdot \vec{p}_{t}} \prod_{i=1}^{n} e^{-i \vec{b} \cdot \vec{k}_{t, i}}
$$

[Parisi, Petronzio '78; Collins, Soper, Sterman '85]
Transverse-momentum conservation is respected
All-order result
[Catani, Grazzini '11][Catani et al. '12]

$$
\begin{aligned}
\frac{d^{2} \Sigma(v)}{d \Phi_{B} d p_{t}} & =\sum_{c_{1}, c_{2}} \frac{d\left|M_{B}\right|_{c_{1} c_{2}}^{2}}{d \Phi_{B}} \int b d b p_{t} J_{0}\left(p_{t} b\right) \mathbf{f}^{T}\left(b_{0} / b\right) \mathbf{C}_{N_{1}}^{c_{i} ; T}\left(\alpha_{S}\left(b_{0} / b\right)\right) H_{\operatorname{CSS}}(M) \mathbf{C}_{N_{2}}^{c_{2}}\left(\alpha _ { S } \left(b_{0} / l\right.\right. \\
& \times \exp \left\{-\sum_{\ell=1}^{2} \int_{0}^{M} \frac{d k_{t}}{k_{t}} \mathbf{R}_{\operatorname{CSS}, \ell}^{\prime}\left(k_{t}\right) \Theta\left(k_{t}-\frac{b_{0}}{b}\right)\right\} \\
R_{\mathrm{CSS}}(b) & =\sum_{l=1}^{2} \int_{b_{0} / b}^{M} \frac{d k_{T}}{k_{T}} R_{\operatorname{CSS}, l}^{\prime}\left(k_{T}\right)=\sum_{l=1}^{2} \int_{b_{0} / b}^{M} \frac{d k_{T}}{k_{T}}\binom{\left.A_{\operatorname{CS}, \ell}\left(\alpha_{s}\left(k_{T}\right)\right) \ln \frac{M^{2}}{k_{T}^{2}}+B_{\operatorname{CSS}, \ell}\left(\alpha_{s}\left(k_{T}\right)\right)\right)}{\text { anomalous dirmensions }}
\end{aligned}
$$

[Davies, Stirling '84] [De Florian, Grazzini '01]
[Becher, Neubert '10][Li, Zhu '16][Vladimirov '16]

## Resummation \& factorization

All-order resummation based on factorization properties

- Of the amplitudes: when radiation becomes soft and/or collinear amplitudes factorize up to regular terms

Necessary condition to establish an all-order formulation since the same structures must appear at all-orders

- Of the observable: in the presence of multiple emissions $k_{i}$, the observable is related to the radiation through phase-space constraints

$$
\Sigma(v) \sim \int\left[d k_{i}\right] M\left(k_{1}, \ldots, k_{n}\right) \Theta\left(v-V\left(k_{1}, \ldots k_{n}\right)\right)
$$

Factorization seems required to disentangle the phase-space constraints
Kinematic factorization is however process-dependent, and must be performed separately for each observable. Typically performed in a conjugate space where factorization is manifest, like for the $p_{t}$ case

## Resummation \& factorization

Resummation techniques based on observable factorization very successful for various observables

However, approach have some limitations

- only observables for which a factorization theorem is known can be resummed
- since factorization is usually achieved in a conjugate space, one has to compute an inverse transform, which sometime causes numerical instabilities.

Is it possible to achieve resummation without the need to establish factorization properties on a case-by-case basis?

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Wes!

## The CAESAR/ARES method: resummation in direct space

Translate the resummability of the observable into properties of the observable in the presence of multiple radiation: recursive infrared and collinear safety (rIRC)
[Banfi, Salam, Zanderighi '01, '03, '04]
a) in the presence of multiple soft and/or collinear emissions the observable has the same scaling properties as with just one of them
b) there exists a resolution scale $q_{0}$, independent of the observable, such that emissions below $q_{0}$ do not contribute significantly to the observable's value.

Unresolved emission can be treated as totally uncorrelated Conjugate space unnecessary as resolved emission can be treated exclusively in momentum space with Monte Carlo methods

Observable to resum at
NkLL accuracy

Only $\mathrm{N}^{k-1} \mathrm{LL}$ ingredients enter in $\mathscr{F}$ thanks to rIRC safety

"Simple" observable whose resummation is known analytically

## Resummation in direct space: the $\boldsymbol{p}_{\boldsymbol{t}}$ case

Non-trivial problem: not possible to find a closed analytic expression in direct space which is both
a) free of logarithmically subleading corrections
b) free of singularities at finite $p_{t}$ values

A naive logarithmic counting at small $p_{t}$ is not sensible, as one loses the correct power-suppressed scaling if only logarithms are retained

It is not possible to reproduce a power-like behaviour with logs of $p_{t} / M$

Can we apply the CAESAR method to transverse-momentum resummation?

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Can we apply the CAESAR method to transverse-momentum resummation?
[Bizon, Monni, Re, LR, Torrielli '17]

## All-order structure of the matrix element

All-order cumulative cross section can be written as

$$
v=p_{t} / M
$$

single-particle phase space


## All-order structure of the matrix element

To find resummed expression one needs to establish an explicit logarithmic counting for the squared matrix element $\left|\cdot M\left(\Phi_{B}, k_{1}, \ldots k_{n}\right)\right|^{2}$

Possibile to do that by decomposing the squared amplitude in terms of $\boldsymbol{n}$-particle correlated blocks, denoted by $\left|\tilde{M}\left(k_{1}, \ldots, k_{n}\right)\right|^{2}\left(\left|\cdot \tilde{M}\left(k_{1}\right)\right|^{2}=\left|\cdot \mathscr{M}\left(k_{1}\right)\right|^{2}\right)$

$$
\left.\begin{array}{l}
\sum_{n=0}^{\infty}\left|\mathscr{M}\left(\Phi_{B}, k_{1}, \ldots, k_{n}\right)\right|^{2}=\mid \mathscr{M}_{B}\left(\left.\Phi_{B}\right|^{2}\right. \\
\quad \times \sum_{n=0}^{\infty} \frac{1}{n!}\left\{\prod _ { i = 1 } ^ { n } \left(\left|\mathscr{M}\left(k_{i}\right)\right|^{2}+\int\left[d k_{a}\right]\left[d k_{b}\right]\left|\tilde{M}\left(k_{a}, k_{b}\right)\right|^{2} \delta^{(2)}\left(\vec{k}_{t a}+\vec{k}_{t b}-\vec{k}_{t i}\right) \delta\left(Y_{a b}-Y_{i}\right)\right.\right. \\
\text { inclusive observable }
\end{array}\right\}
$$

In the soft-collinear limit, $\prod_{i=1}^{n}\left|\mathscr{M}\left(k_{i}\right)\right|$ comes with a factor $\alpha_{s}^{n} \ln { }^{2 n}(v)$ whereas correlated blocks with $n$ emissions $\left|\tilde{\mathscr{M}}\left(k_{1}, \ldots k_{n}\right)\right|$ contribute at most with $\alpha_{s}^{n} \ln ^{n+1}(v)$ thanks to rIRC safety

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$$
\begin{aligned}
& \sum_{n=0}^{\infty}\left|\mathscr{M}\left(\Phi_{B}, k_{1}, \ldots, k_{n}\right)\right|^{2}=\mid \mathscr{M}_{B}\left(\left.\Phi_{B}\right|^{2}\right. \\
& \times \sum_{n=0}^{\infty} \frac{1}{n!}\left\{\prod _ { i = 1 } ^ { n } \left(\left|\mathscr{M}\left(k_{i}\right)\right|^{2}+\int\left[d k_{a}\right]\left[d k_{b}\right]\left|\tilde{M}\left(k_{a}, k_{b}\right)\right|^{2} \delta^{(2)}\left(\vec{k}_{t a}+\vec{k}_{t b}-\vec{k}_{t i}\right) \delta\left(Y_{a b}-Y_{i}\right)\right.\right. \\
& \left.\left.+\int\left[d k_{a}\right]\left[d k_{b}\right]\left[d k_{c}\right]\left|\tilde{M}\left(k_{a}, k_{b}, k_{c}\right)\right|^{2} \delta^{(2)}\left(\vec{k}_{t a}+\vec{k}_{t b}+\vec{k}_{t c}-\vec{k}_{t i}\right) \delta\left(Y_{a b c}-Y_{i}\right)+\ldots\right)\right\} \\
& \equiv\left|\mathscr{M}_{B}\left(\Phi_{B}\right)\right|^{2} \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=1}^{n}\left|\mathscr{M}\left(k_{i}\right)\right|_{\text {inc }}^{2} \\
& \text { *expression valid for } \\
& \text { inclusive observables }
\end{aligned}
$$

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$$
\begin{aligned}
& \sum_{n=0}^{\infty}\left|\mathscr{M}\left(\Phi_{B}, k_{1}, \ldots, k_{n}\right)\right|^{2}=\mid \mathscr{M}_{B}\left(\left.\Phi_{B}\right|^{2}\right. \\
& \times \sum_{n=0}^{\infty} \frac{1}{n!}\left\{\prod _ { i = 1 } ^ { n } \left(\left\lvert\, \begin{array}{c}
\mathbf{L L} \\
\left|\mathscr{M}\left(k_{i}\right)\right|^{2}+\int\left[d k_{a}\right]\left[d k_{b}\right]\left|\tilde{M}\left(k_{a}, k_{b}\right)\right|^{2} \delta^{(2)}\left(\vec{k}_{t a}+\vec{k}_{t b}-\vec{k}_{t i}\right) \delta\left(Y_{a b}-Y_{i}\right)
\end{array}\right.\right.\right. \\
& \left.\left.+\int\left[d k_{a}\right]\left[d k_{b}\right]\left[d k_{c}\right]\left|\tilde{M}\left(k_{a}, k_{b}, k_{c}\right)\right|^{2} \delta^{(2)}\left(\vec{k}_{t a}+\vec{k}_{t b}+\vec{k}_{t c}-\vec{k}_{t i}\right) \delta\left(Y_{a b c}-Y_{i}\right)+\ldots\right)\right\} \\
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\end{aligned}
$$

Upon integration over the phase space, the expansion can be put in a one to one correspondence with the logarithmic structure

$$
\ln \left|\mathscr{M}\left(\Phi_{B}, k_{1}, \ldots, k_{n}\right)\right|^{2} \rightarrow \mathcal{O}\left(\alpha_{s}^{n} \ln (v)^{n+1}\right)+\mathcal{O}\left(\alpha_{s}^{n} \ln (v)^{n}\right)+\mathcal{O}\left(\alpha_{s}^{n} \ln (v)^{n-1}\right) \ldots
$$

Systematic recipe to include terms up to the desired logarithmic accuracy

## Cancellation of the IRC singularities

Exploit rIRC safety of the observable to single out the IRC singularities of the real matrix element and achieve the cancellation of the exponentiated divergences of virtual origin

Introduce a slicing parameter $\epsilon \ll 1$ such that all inclusive blocks with $k_{t, i}<k_{t, 1}, k_{t, 1}$ hardest emission, can be neglected in the computation of the observable

$$
\begin{aligned}
& \Sigma(v)=\int d \Phi_{B}\left|\mathscr{M}_{B}\left(\Phi_{B}\right)\right|^{2} \mathscr{V}\left(\Phi_{B}\right) \\
& \times \int\left[d k_{1}\right]\left|\mathscr{M}\left(k_{1}\right)\right|_{\text {inc }}^{2}\left(\sum_{i=0}^{\infty} \frac{1}{l!} \int \prod_{j=2}^{l+1}\left[d k_{j}\right]\left|\cdot \mathscr{M}\left(k_{j}\right)\right|_{\text {inc }}^{2} \Theta\left(\epsilon V\left(k_{1}\right)-V\left(k_{j}\right)\right)\right) \\
& \times\left(\sum_{m=0}^{\infty} \frac{1}{m!} \int \prod_{i=2}^{m+1}\left[d k_{i}\right]\left|\mathscr{M}\left(k_{i}\right)\right|_{\text {inc }}^{2} \Theta\left(V\left(k_{i}\right)-\epsilon V\left(k_{1}\right)\right) \Theta\left(v-V\left(\Phi_{B}, k_{1}, \ldots, k_{m+1}\right)\right)\right)
\end{aligned}
$$

resolved emissions

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$$
\begin{aligned}
& \Sigma(v)=\int d \Phi_{B}\left|\cdot M_{B}\left(\Phi_{B}\right)\right|^{2} \mathscr{V}\left(\Phi_{B}\right) \quad \text { unresolved emissions } \\
& \left.\times \int\left[d k_{1}\right] \mid \cdot M\left(k_{1}\right)\right)_{\text {inc }}^{2}\left(\sum_{l=0}^{\infty} \frac{1}{l!} \int \prod_{j=2}^{t+1}\left[d k_{j}\right]\left|\cdot M\left(k_{j}\right)\right| l_{\text {inc }}^{2} \Theta\left(\epsilon V\left(k_{1}\right)-V\left(k_{j}\right)\right)\right. \\
& \times\left(\sum_{m=0}^{\infty} \frac{1}{m!} \int \prod_{i=2}^{n+1}\left[d k_{i}\right]\left|\cdot M\left(k_{i}\right)\right|_{\text {inc }}^{2} \Theta\left(V\left(k_{i}\right)-\epsilon V\left(k_{1}\right)\right) \Theta\left(v-V\left(\Phi_{B}, k_{1}, \ldots, k_{m+1}\right)\right)\right)
\end{aligned}
$$

Unresolved emission doesn't contribute to the evaluation of the observable: it can be exponentiated directly and employed to cancel the virtual divergences, giving rise to a Sudakov radiator

$$
\mathscr{V}\left(\Phi_{B}\right) \exp \left\{\int[d k]|\cdot M(k)|_{\text {inc }}^{2} \Theta\left(\epsilon V\left(k_{1}\right)-V(k)\right)\right\} \simeq e^{-R\left(\epsilon V\left(k_{1}\right)\right)}
$$

## Result at NLL accuracy

Result at NLL accuracy can be written as

$$
\begin{aligned}
\Sigma(v)= & \sigma^{(0)} \int \frac{d v_{1}}{v_{1}} \int_{0}^{2 \pi} \frac{d \phi_{1}}{2 \pi} e^{-R\left(\epsilon v_{1}\right)} R^{\prime}\left(v_{1}\right) \\
& \times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^{1} \frac{d \zeta_{i}}{\zeta_{i}} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} R^{\prime}\left(\zeta_{i} v_{1}\right) \Theta\left(v-V\left(\Phi_{B}, k_{1}, \ldots, k_{n+1}\right)\right)
\end{aligned}
$$

Formula can be evaluated with Monte Carlo method; dependence on $\epsilon$ vanishes exactly and result is finite in four dimensions

It contains subleading effect which in the original CAESAR approach are disposed of by expanding $R$ and $R^{\prime}$ around $v$

$$
\begin{aligned}
& R\left(\epsilon v_{1}\right)=R(v)+\frac{d R(v)}{d \ln (1 / v)} \ln \frac{v}{\epsilon v_{1}}+\mathcal{O}\left(\ln ^{2} \frac{v}{\epsilon v_{1}}\right) \\
& R^{\prime}\left(v_{i}\right)=R^{\prime}(v)+\mathcal{O}\left(\ln \frac{v}{v_{i}}\right)
\end{aligned}
$$

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Not possible! valid only if the ratio $v_{i} / v$ remains of order one in the whole emission phase space, but for observables which feature kinematic cancellations there are configurations with $v_{i} \gg v$. Subleading effects necessary

## Result at NLL accuracy

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$$
\begin{aligned}
\Sigma(v)= & \sigma^{(0)} \int \frac{d v_{1}}{v_{1}} \int_{0}^{2 \pi} \frac{d \phi_{1}}{2 \pi} e^{-R\left(\epsilon v_{1}\right)} R^{\prime}\left(v_{1}\right) \quad v_{i}=V\left(k_{i}\right), \quad \zeta_{i}=v_{i} / v_{1} \\
& \times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^{1} \frac{d \zeta_{i}}{\zeta_{i}} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} R^{\prime}\left(\zeta_{i} v_{1}\right) \Theta\left(v-V\left(\Phi_{B}, k_{1}, \ldots, k_{n+1}\right)\right)
\end{aligned}
$$

Formula can be evaluated with Monte Carlo method; dependence on $\epsilon$ vanishes exactly and result is finite in four dimensions

Convenient to perform an expansion around $v_{1}$ (more efficient and simpler implementation)

$$
\begin{aligned}
& R\left(\epsilon v_{1}\right)=R\left(v_{1}\right)+\frac{d R\left(v_{1}\right)}{d \ln \left(1 / v_{1}\right)} \ln \frac{1}{\epsilon}+\mathcal{O}\left(\ln ^{2} \frac{1}{\epsilon}\right) \\
& R^{\prime}\left(v_{i}\right)=R^{\prime}\left(v_{1}\right)+\mathcal{O}\left(\ln \frac{v_{1}}{v_{i}}\right)
\end{aligned}
$$

## Result at NLL accuracy

Final result including parton luminosity

$$
\begin{aligned}
\frac{d \Sigma(v)}{d \Phi_{B}} & =\int \frac{d k_{t, 1}}{k_{t, 1}} \int_{0}^{2 \pi} \frac{d \phi_{1}}{2 \pi} e^{-R\left(k_{t, 1}\right)} e^{R^{\prime}\left(k_{t, 1}\right)} \mathscr{L}_{\mathrm{NLL}}\left(k_{t, 1}\right) R^{\prime}\left(k_{t, 1}\right) \\
& \times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^{1} \frac{d \zeta_{i}}{\zeta_{i}} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} R^{\prime}\left(\zeta_{i} k_{t, 1}\right) \Theta\left(v-V\left(\Phi_{B}, k_{1}, \ldots, k_{n+1}\right)\right)
\end{aligned}
$$

Parton luminosity at NLL reads

$$
\mathscr{L}_{\mathrm{NLL}}\left(k_{t, 1}\right)=\sum_{c} \frac{d\left|M_{B}\right|_{c \bar{c}}^{2}}{d \Phi_{B}} f_{c}\left(x_{1}, k_{t, 1}^{2}\right) f_{\bar{c}}\left(x_{2}, k_{t, 1}^{2}\right)
$$

At higher logarithmic accuracy, it includes coefficient functions and hard-virtual corrections

This formula can be evaluated by means of fast Monte Carlo methods
RadISH (Radiation off Initial State Hadrons)

## Result at N3 ${ }^{3} L$ accuracy

$$
\left.\begin{array}{l}
\frac{d \Sigma(v)}{d \Phi_{B}}=\int \frac{d k_{t 1}}{k_{t 1}} \frac{d \phi_{1}}{2 \pi} \partial_{L}\left(-e^{-R\left(k_{t 1}\right)} \mathcal{L}_{\mathrm{N}^{3} \mathrm{LL}}\left(k_{t 1}\right)\right) \int d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right] \Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}\right)\right) \\
+\int \frac{d k_{t 1}}{k_{t 1}} \frac{d \phi_{1}}{2 \pi} e^{-R\left(k_{t 1}\right)} \int d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right] \int_{0}^{1} \frac{d \zeta_{s}}{\zeta_{s}} \frac{d \phi_{s}}{2 \pi}\left\{\left(R^{\prime}\left(k_{t 1}\right) \mathcal{L}_{\mathrm{NNLL}}\left(k_{t 1}\right)-\partial_{L} \mathcal{L}_{\mathrm{NNLL}}\left(k_{t 1}\right)\right)\right. \\
\times\left(R^{\prime \prime}\left(k_{t 1}\right) \ln \frac{1}{\zeta_{s}}+\frac{1}{2} R^{\prime \prime \prime}\left(k_{t 1}\right) \ln ^{2} \frac{1}{\zeta_{s}}\right)-R^{\prime}\left(k_{t 1}\right)\left(\partial_{L} \mathcal{L}_{\mathrm{NNLL}}\left(k_{t 1}\right)-2 \frac{\beta_{0}}{\pi} \alpha_{s}^{2}\left(k_{t 1}\right) \hat{P}^{(0)} \otimes \mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right) \ln \frac{1}{\zeta_{s}}\right) \\
\left.+\frac{\alpha_{s}^{2}\left(k_{t 1}\right)}{\pi^{2}} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right)\right\}\left\{\Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}, k_{s}\right)\right)-\Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}\right)\right)\right\} \\
+\frac{1}{2} \int \frac{d k_{t 1}}{k_{t 1}} \frac{d \phi_{1}}{2 \pi} e^{-R\left(k_{t 1}\right)} \int d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right] \int_{0}^{1} \frac{d \zeta_{s 1}}{\zeta_{s 1}} \frac{d \phi_{s 1}}{2 \pi} \int_{0}^{1} \frac{d \zeta_{s 2}}{\zeta_{s 2}} \frac{d \phi_{s 2}}{2 \pi} R^{\prime}\left(k_{t 1}\right) \\
\times\left\{\mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right)\left(R^{\prime \prime}\left(k_{t 1}\right)\right)^{2} \ln \frac{1}{\zeta_{s 1}} \ln \frac{1}{\zeta_{s 2}}-\partial_{L} \mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right) R^{\prime \prime}\left(k_{t 1}\right)\left(\ln \frac{1}{\zeta_{s 1}}+\ln \frac{1}{\zeta_{s 2}}\right)\right. \\
\left.+\frac{\alpha_{s}^{2}\left(k_{t 1}\right)}{\pi^{2}} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right)\right\} \\
\times\left\{\Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}, k_{s 1}, k_{s 2}\right)\right)-\Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}, k_{s 1}\right)\right)-\right. \\
\left.\Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}, k_{s 2}\right)\right)+\Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}\right)\right)\right\}+\mathcal{O}\left(\alpha_{s}^{n} \ln 2 n-6\right.  \tag{3.18}\\
v
\end{array}\right),
$$

Result formally equivalent to the $b$-space formulation
[Parisi, Petronzio '78][Collins, Soper, Sterman '85]

## Implementation: matching to fixed order

$$
\Sigma\left(v, \phi_{B}\right)=\int_{0}^{v} d v^{\prime} \frac{d \sigma}{d v^{\prime} d \phi_{B}} \quad \begin{aligned}
& \begin{array}{l}
\text { Cumulative cross } \\
\text { section should } \\
\text { reduce to the fixed } \\
\text { order at large } v
\end{array}
\end{aligned} \rightarrow \Sigma_{\text {res }} \quad \rightarrow \Sigma_{\text {f.o. }} \quad p_{t} \gtrsim<M_{B}
$$

Additive matching

$$
\Sigma_{\text {add }}^{\text {matched }}(v)=\Sigma^{\text {res }}(v)+\Sigma^{\text {fo. }}(v)-\Sigma^{\text {expanded }}(v) \quad \Sigma_{\text {matched }}^{\text {mult }}(v)=\Sigma_{\text {res }}(v)\left[\frac{\Sigma_{\text {f.o. }}(v)}{\Sigma_{\text {res }}(v)}\right]
$$

- it allows one to extract the relative $O\left(\alpha^{3}\right)$ constant terms from the fixedorder whenever the $\mathrm{N}^{3} \mathrm{LO}$ total cross section is known, e.g. Higgs
- only viable solution till constant terms are not known analytically to consistently match to NNLO
- numerically more stable as the physical suppression at small $v$ cures potential instabilities


## Implementation: matching to fixed order

$$
\Sigma\left(v, \phi_{B}\right)=\int_{0}^{v} d v^{\prime} \frac{d \sigma}{d v^{\prime} d \phi_{B}}
$$

Cumulative cross
section should $\quad \rightarrow \Sigma_{\text {res }} \quad p_{t} \ll M_{B}$
reduce to the fixed
order at large $v$

## Additive matching

$\Sigma_{\text {add }}^{\text {matched }}(v)=\Sigma^{\text {res }}(v)+\Sigma^{\text {f.o. }}(v)-\Sigma^{\operatorname{expanded}}(v)$

- perhaps more natural, simpler
- numerically delicate in the very small $v$ limit as f.o. can be unstable


## Multiplicative matching

$\Sigma_{\text {matched }}^{\text {mult }}(v)=\Sigma_{\text {res }}(v)\left[\frac{\Sigma_{\text {f.o. }}(v)}{\Sigma_{\text {res }}(v)}\right]$

- it allows one to extract the relative $O\left(\alpha^{3}\right)$ constant terms from the fixedorder whenever the $\mathrm{N}^{3} \mathrm{LO}$ total cross section is known, e.g. Higgs
- only viable solution till constant terms are not known analytically to consistently match to NNLO
- numerically more stable as the physical suppression at small $v$ cures potential instabilities


## Multiplicative matching

$$
\Sigma_{\text {matched }}^{\text {mult }}(v)=\Sigma_{\text {res }}(v)\left[\frac{\Sigma_{\text {f.o. }}(v)}{\Sigma_{\text {res }}(v)}\right]_{\text {expanded }}
$$

Drawback: the fixed-order result at large $v$ receives spurious contributions; e.g. at $\mathrm{N}^{3} \mathrm{LO}$

$$
\Sigma_{\text {mult }}^{\text {matched }}(v) \sim \Sigma^{\mathrm{N}^{3} \mathrm{LO}}(v)\left(1+\mathcal{O}\left(\alpha_{s}^{4}\right)\right)
$$

Reason: when logarithms $L$ tend to zero, $\Sigma_{\text {res }}(v)$ tends to

$$
\Sigma_{\text {asym. }}^{\mathrm{res}}=\int_{\text {with cuts }} d \Phi_{B}\left(\lim _{L \rightarrow 0} \mathscr{L}_{\mathrm{N}^{k} \mathrm{LL}}\right)
$$

Solution: normalize to the asymptotic value

$$
\Sigma_{\text {mult }}^{\operatorname{matched}}(v)=\frac{\Sigma^{\mathrm{res}}(v)}{\Sigma_{\text {asym. }}^{\mathrm{res}}}\left[\sum_{\text {asym. }}^{\mathrm{res}} \frac{\Sigma^{\mathrm{f.o}}(v)}{\sum^{\exp }(v)}\right]_{\text {expanded }}
$$

## Higgs transverse momentum at NNLO+N3LL: inclusive

[Bizon, Monni, Re, LR, Torrielli + NNLOJET '18]
$\mathbf{N}^{3}$ LL vs NNLL


- $\mathrm{N}^{3}$ LL corrections moderate in size ( $\sim 5 \%$ at low $\mathrm{p}_{\text {т }}$ ) and contained in the NNLO+NNLL band
- Reduction of the perturbative uncertainty by a factor of 2 for $p_{t} \leqslant 10 \mathrm{GeV}$


## Higgs transverse momentum at NNLO+N3LL: inclusive



- Effect of resummation starts to be increasingly important for $p_{t} \leqslant 40 \mathrm{GeV}$
- Resummation effects are progressively less important above 50 GeV
- Heavy-quark mass effects start to be relevant at this level of precision


## Higgs transverse momentum at NNLO+N3LL: fiducial cuts

- Effect of resummation starts to be increasingly important for $p_{t} \leqslant 40 \mathrm{GeV}$
- Resummation effects are progressively less important above 50 GeV
- Heavy-quark mass effects start to be relevant at this level of precision
- Similar results for fiducial region


## Drell-Yan



- State-of-the-art QCD prediction do not match the precision of the data
- LO MC are used, tuned on $Z$ data
- Would be preferable to use more accurate theoretical predictions

Extreme precision is needed for $W$ mass extraction


- Template fits to lepton observables
- Modelling of $p_{t, w}$ is crucial. Fit predictions to $Z$ data, apply to $W$


## Drell-Yan transverse momentum

## Comparison with ATLAS data @ 8 TeV [1512.02192]



- Matched results offer a good description of the data in the low-medium $p_{\text {T }}$ range, in all fiducial regions
- Perturbative uncertainty at the few percent level, still does not match the precision of the data
- Estimate of nonperturbative effects may start to be relevant


## Drell-Yan $\varphi^{*}$

Approach can be used for resumming other transverse observables; e.g $\varphi^{*}$
Comparison with ATLAS data @ 8 TeV [1512.02192]


$$
\phi^{*}=\tan \left(\frac{\pi-\Delta \phi}{2}\right) \sin \theta^{*} \quad \phi^{*} \sim \frac{p_{t}}{2 M}
$$

angle between electron and beam axis, in $Z$ boson rest frame

- Similar situation as $p_{t}$, with perturbative uncertainty at the few percent level but with experimental errors at the sub-percent level


## Conclusion

- No sign of NP at the LHC so far - necessary to perform detailed theory/ experimental comparisons, to look for deviations from SM. Perturbation theory must be pushed at its limit
- New formalism formulated in momentum space for all-order resummation up to $\mathbf{N}^{3}$ LL accuracy for inclusive, transverse observables.
- Access to multi-differential information. This is effectively similar to a semi-inclusive parton shower, but with higher-order logarithms, and control on formal accuracy
- Method allows for an efficient implementation in a computer code. Towards a single generator able to resum entire classes of observables at high accuracy.
- Results at NNLO+N3LL for Higgs and DY differential distributions
- Good description of the data in the fiducial distributions, with uncertainties at the few percent level


## Backup

## Parton luminosities

Consider configurations in which emissions are ordered in $k_{t, i,} k_{t, 1}$ hardest emission Phase space for each secondary emission can be depicted in the Lund diagram
$\qquad$

DGLAP evolution governs the radiation in the strictly collinear limit

resolved emissions live in this strip
rapidity in the centre-of-mass frame of the incoming partons

- DGLAP evolution can be performed inclusively up to $\epsilon k_{t, 1}$ thanks to rIRC safety
- In the overlapping region hard-collinear emissions modify the observable's value: the evolution should be performed exclusively (unintegrated in $k_{t}$ )
- At NLL the real radiation can be approximated with its soft limit: DGLAP can be performed inclusively up to $k_{t, 1}$ (i.e. one can evaluate $\mu_{\mathrm{F}}=k_{t, 1}$ )


## Beyond NLL

Extension to NNLL and beyond requires the systematic inclusion of the correlated blocks necessary to achieve the desired logarithmic accuracy

Moreover, one needs to relax a series of assumptions which give rise to subleading corrections neglected at NLL (for instance, exact rapidity bounds). These corrections can be included systematically by including additional terms in the expansion

$$
R\left(\epsilon v_{1}\right)=R\left(v_{1}\right)+\frac{d R\left(v_{1}\right)}{d \ln \left(1 / v_{1}\right)} \ln \frac{1}{\epsilon}+\mathcal{O}\left(\ln ^{2} \frac{1}{\epsilon}\right)
$$

Finally, one needs to specify a complete treatment for hard-collinear radiation. Starting at NNLL one or more real emissions can be hard and collinear to the emitting leg, and the available phase space for subsequent real emissions changes
Two classes of contributions:

- one soft by construction and which is analogous to the $\mathrm{R}^{\prime}$ contribution

$$
R^{\prime}\left(v_{i}\right)=R^{\prime}\left(v_{1}\right)+\mathcal{O}\left(\ln \frac{v_{1}}{v_{i}}\right)
$$

- another hard and collinear (exclusive DGLAP step): last step of DGLAP evolution must be performed unintegrated in $k_{t}$


## Logarithmic counting

Necessary to establish a well defined logarithmic counting: possibile to do that by decomposing the squared amplitude in terms of n-particle correlated blocks (nPC)
e.g. $p p \rightarrow H+$ emission of up to 2 (soft) gluons $O\left(\alpha_{s}{ }^{2}\right)$


Logarithmic counting defined in terms of nPC blocks (owing to rIRC safety of the observable)

## Logarithmic counting: correlated blocks



Thanks to P. Monni

## Equivalence with $b$-space formulation

$$
\frac{d \Sigma(v)}{d \Phi_{B}}=\int_{\mathscr{C}_{1}} \frac{d N_{1}}{2 \pi i} \int_{\mathscr{C}_{2}} \frac{d N_{2}}{2 \pi i} x_{1}^{-N_{1}} x_{2}^{-N_{2}} \sum_{c_{1}, c_{2}} \frac{d\left|M_{B}\right|_{c_{1} c_{2}}^{2}}{d \Phi_{B}} \mathbf{f}_{N_{1}}^{T}\left(\mu_{0}\right) \hat{\boldsymbol{\Sigma}}_{N_{1}, N_{2}}^{c_{1}, c_{2}}(v) \mathbf{f}_{N_{2}}\left(\mu_{0}\right)
$$

unresolved emission + virtual corrections

Result valid for all inclusive observables (e.g. $p_{t,} \varphi^{*}$ )

$$
\begin{aligned}
& \hat{\boldsymbol{\Sigma}}_{N_{1}, N_{2}}^{c_{1}, c_{2}}(v)=\left[\mathbf{C}_{N_{1}}^{c_{1}, T}\left(\alpha_{s}\left(\mu_{0}\right)\right) H\left(\mu_{R}\right) \mathbf{C}_{N_{2}}^{c_{2}}\left(\alpha_{s}\left(\mu_{0}\right)\right)\right] \int_{0}^{M} \frac{d k_{t 1}}{k_{t 1}} \int_{0}^{2 \pi} \frac{d k_{1}}{2 \pi} \exp \left\{-\sum_{t=1}^{2}\left(\int_{e k_{1}}^{\mu_{0}} \frac{d k_{t}}{k_{t}} \frac{\alpha_{s}\left(k_{t}\right)}{\pi} \boldsymbol{\Gamma}_{N_{t}}\left(\alpha_{s}\left(k_{t}\right)\right)+\int_{\epsilon k_{1}}^{\mu_{1}} \frac{d k_{t}}{k_{t}} \boldsymbol{\Gamma}_{N_{t}}^{(C)}\left(\alpha_{s}\left(k_{t}\right)\right)\right)\right\}
\end{aligned}
$$

resolved

$$
\sum_{\ell_{1}=1}^{2}\left(\mathbf{R}_{\ell_{1}}^{\prime}\left(k_{t 1}\right)+\frac{\alpha_{s}\left(k_{t 1}\right)}{\pi} \boldsymbol{\Gamma}_{N_{1}}\left(\alpha_{s}\left(k_{t 1}\right)\right)+\Gamma_{N_{\ell_{1}}}^{(C)}\left(\alpha_{s}\left(k_{t 1}\right)\right)\right)
$$ emission

$$
\times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^{1} \frac{d \zeta_{i}}{\zeta_{i}} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} \sum_{\ell_{i}=1}^{2}\left(\mathbf{R}_{\ell_{i}}^{\prime}\left(k_{t i}\right)+\frac{\alpha_{s}\left(k_{t i}\right)}{\pi} \boldsymbol{\Gamma}_{N_{i}}\left(\alpha_{s}\left(k_{t i}\right)\right)+\boldsymbol{\Gamma}_{N_{t_{i}}}^{(C)}\left(\alpha_{s}\left(k_{t i}\right)\right)\right)
$$

$$
\times \Theta\left(v-V\left(\{\tilde{p}\}, k_{1}, \ldots, k_{n+1}\right)\right)
$$

Formulation equivalent to $\boldsymbol{b}$-space result (up to a scheme change in the anomalous dimensions)

$$
\begin{aligned}
\frac{d^{2} \Sigma(v)}{d \Phi_{B} d p_{t}}= & \sum_{c_{1}, c_{2}} \frac{d\left|M_{B}\right|_{c_{1} c_{2}}^{2}}{d \Phi_{B}} \int b d b p_{t} J_{0}\left(p_{t} b\right) \mathbf{f}^{T}\left(b_{0} / b\right) \mathbf{C}_{N_{1}}^{c_{1} ; T}\left(\alpha_{s}\left(b_{0} / b\right)\right) H(M) \mathbf{C}_{N_{2}}^{c_{2}}\left(\alpha_{s}\left(b_{0} / b\right)\right) \mathbf{f}\left(b_{0} / b\right) \\
& \times \exp \left\{-\sum_{\ell=1}^{2} \int_{0}^{M} \frac{d k_{t}}{k_{t}} \mathbf{R}_{\ell}^{\prime}\left(k_{t}\right)\left(1-J_{0}\left(b k_{t}\right)\right)\right\} \quad\left(1-J_{0}\left(b k_{t}\right)\right) \simeq \Theta\left(k_{t}-\frac{b_{0}}{b}\right)+\frac{\zeta_{3}}{12} \frac{\partial^{3}}{\partial \ln \left(M b / b_{0}\right)^{3}} \Theta\left(k_{t}-\frac{b_{0}}{b}\right)
\end{aligned}
$$

$\mathrm{N}^{3}$ LL effect: absorbed in the definition of $H_{2}, B_{3}, A_{4}$ coefficients wrt to CSS

## Behaviour at small $\boldsymbol{p}_{\boldsymbol{t}}$

## pT vs. ET: dependence on the first emission



## Behaviour at small $\boldsymbol{p}_{t}$

Explicit evaluation shows that the Parisi-Petronzio perturbative scaling at small $p_{t}$ is reproduced. At NLL, Drell-Yan pair production, $n_{f}=4$

$$
\frac{d^{2} \Sigma(v)}{d p_{t} d \Phi_{B}}=4 \sigma^{(0)}\left(\Phi_{B}\right) p_{t} \int_{\Lambda_{\text {ecD }}}^{M} \frac{d k_{t 1}}{k_{t 1}^{3}} e^{-R\left(k_{t 1}\right)} \simeq 2 \sigma^{(0)}\left(\Phi_{B}\right) p_{t}\left(\frac{\Lambda_{\mathrm{CCD}}^{2}}{M^{2}}\right)^{\frac{16}{25} \ln \frac{4 t}{16}}
$$

As now higher logarithmic terms (up to $\mathrm{N}^{3} \mathrm{LL}$ ) are under control, the coefficient of this scaling can be systematically improved in perturbation theory (non-perturbative effects - of the same order - not considered)
$\mathrm{N}^{3} \mathrm{LL}$ calculation allows one to have control over the terms of relative order $O\left(\alpha_{\mathrm{s}}{ }^{2}\right)$. Scaling $L \sim 1 / \alpha_{s}$ valid in the deep infrared regime.

## Numerical implementation

$$
\begin{aligned}
\frac{d \Sigma\left(p_{t}\right)}{d \Phi_{B}}= & \int_{0}^{M} \frac{d k_{t 1}}{k_{t 1}} \int_{0}^{2 \pi} \frac{d \phi_{1}}{2 \pi} \partial_{L}\left(-e^{-R^{\prime}\left(k_{t 1}\right)} \mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right)\right) \times \\
& \times \underbrace{\epsilon^{R^{\prime}\left(k_{t 1}\right)} \sum_{n=0}^{\infty} \frac{1}{n!}\left(\prod_{i=2}^{n+1} \int_{\epsilon k_{t 1}}^{k_{t 1}} \frac{d k_{t i}}{k_{t i}} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} R^{\prime}\left(k_{t 1}\right)\right) \Theta\left(p_{t}-\left|\vec{k}_{t 1}+\ldots+\vec{k}_{t(n+1)}\right|\right)}_{\equiv \int d \mathcal{Z}\left\{\left\{R^{\prime}, k_{i}\right\}\right\} \Theta\left(p_{t}-\left|\vec{k}_{t 1}+\ldots+\vec{k}_{t(n+1)}\right|\right)} .
\end{aligned}
$$

- $L=\ln \left(M / k_{t 1}\right) ;$ luminosity $\mathcal{L}_{\mathrm{NLL}}\left(k_{t 1}\right)=\sum_{c_{1}, c_{2}} \frac{d\left|M_{B}\right|_{c_{1} c_{2}}^{2}}{d \Phi_{B}} f_{c_{1}}\left(x_{1}, k_{t 1}\right) f_{c_{2}}\left(x_{2}, k_{t 1}\right)$.
- $\int d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right] \Theta$ finite as $\epsilon \rightarrow 0$ :

$$
\begin{gathered}
\epsilon^{R^{\prime}\left(k_{t 1}\right)}=1-R^{\prime}\left(k_{t 1}\right) \ln (1 / \epsilon)+\ldots=1-\int_{\epsilon k_{t 1}}^{k_{t 1}} R^{\prime}\left(k_{t 1}\right)+\ldots, \\
\int d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right] \Theta=\left[1-\int_{\epsilon k_{t 1}}^{k_{t 1}} R^{\prime}\left(k_{t 1}\right)+\ldots\right]\left[\Theta\left(p_{t}-\left|\vec{k}_{t 1}\right|\right)+\int_{\epsilon k_{t 1}}^{k_{t 1}} R^{\prime}\left(k_{t 1}\right) \Theta\left(p_{t}-\left|\vec{k}_{t 1}+\vec{k}_{t 2}\right|\right)+\ldots\right] \\
= \\
\Theta\left(p_{t}-\left|\vec{k}_{t 1}\right|\right)+\underbrace{\int_{0}^{k_{t 1}}}_{\epsilon \rightarrow 0} R^{\prime}\left(k_{t 1}\right) \underbrace{\left[\Theta\left(p_{t}-\left|\vec{k}_{t 1}+\vec{k}_{t 2}\right|\right)-\Theta\left(p_{t}-\left|\vec{k}_{t 1}\right|\right)\right]}_{\text {finite: real-virtual cancellation }}+\ldots
\end{gathered}
$$

- Evaluated with Monte Carlo techniques: $\int d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right]$ is generated as a parton shower over secondary emissions.


## Numerical implementation

- Secondary radiation:

$$
\begin{aligned}
d \mathcal{Z}\left[\left\{R^{\prime}, k_{i}\right\}\right] & =\sum_{n=0}^{\infty} \frac{1}{n!}\left(\prod_{i=2}^{n+1} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} \int_{\epsilon k_{t 1}}^{k_{t 1}} \frac{d k_{t i}}{k_{t i}} R^{\prime}\left(k_{t 1}\right)\right) \epsilon^{R^{\prime}\left(k_{t 1}\right)} \\
& =\sum_{n=0}^{\infty}\left(\prod_{i=2}^{n+1} \int_{0}^{2 \pi} \frac{d \phi_{i}}{2 \pi} \int_{\epsilon k_{t 1}}^{k_{t(i-1)}} \frac{d k_{t i}}{k_{t i}} R^{\prime}\left(k_{t 1}\right)\right) \epsilon^{R^{\prime}\left(k_{t 1}\right)} \\
\epsilon^{R^{\prime}\left(k_{t 1}\right)} & =e^{-R^{\prime}\left(k_{t 1}\right) \ln 1 / \epsilon}=\prod_{i=2}^{n+2} e^{-R^{\prime}\left(k_{t 1}\right) \ln k_{t(i-1)} / k_{t i}}
\end{aligned}
$$

with $k_{t(n+2)}=\epsilon k_{t 1}$.

- Each secondary emissions has differential probability

$$
d w_{i}=\frac{d \phi_{i}}{2 \pi} \frac{d k_{t i}}{k_{t i}} R^{\prime}\left(k_{t 1}\right) e^{-R^{\prime}\left(k_{t 1}\right) \ln k_{t(i-1)} / k_{t i}}=\frac{d \phi_{i}}{2 \pi} d\left(e^{-R^{\prime}\left(k_{t 1}\right) \ln k_{t(i-1)} / k_{t i}}\right) .
$$

- $k_{t(i-1)} \geq k_{t i}$. Scale $k_{t i}$ extracted by solving $e^{-R^{\prime}\left(k_{t 1}\right) \ln k_{t(i-1)} / k_{t i}}=r$, with $r$ random number extracted uniformly in $[0,1]$. Shower ordered in $k_{t i}$.
- Extract $\phi_{i}$ randomly in $[0,2 \pi]$.


## Factorization and resummation in conjugate space

Phase-space constraints do not usually factorize in direct space
Resummation usually performed in impact-parameter (b) space where the two competing mechanisms are handled trough a Fourier transform. Transversemomentum conservation is respected

$$
\delta\left(\vec{p}_{t}-\sum_{i=1}^{n} \vec{k}_{t, i}\right)=\int d^{2} b \frac{1}{4 \pi^{2}} e^{i \vec{b} \cdot \vec{p}_{t}} \prod_{i=1}^{n} e^{-i \vec{b} \cdot \vec{k}_{t, i}}
$$

[Parisi, Petronzio '78; Collins, Soper, Sterman '85]

Extremely successful approach
Some limitations

- resummation tied to the existence of a resummation theorem for the observable
- process-dependent, must be performed manually and analytically in each case (error prone)
- inverse transform sometime causes numerical instabilities


## Resummation in direct space: the $\boldsymbol{p}_{\boldsymbol{t}}$ case

Non-trivial problem: not possible to find a closed analytic expression in direct space which is both
a) free of logarithmically subleading corrections
b) free of singularities at finite $p_{t}$ values

A naive logarithmic counting at small $p_{t}$ is not sensible, as one loses the correct power-suppressed scaling if only logarithms are retained

Resummation in direct space now possible

[Monni, Re, Torrielli '16, Bizon, Monni, Re, LR, Torielli '17]
[Ebert, Tackmann '16]
see also [Kang,Lee,Vaidya '17]

