

The Leptonic CP Phases: Dirac vs Majorana

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Jarah Evslin, **SFG**, Kaoru Hagiwara, JHEP **1602** (2016) 137 [arXiv:1506.05023]
SFG, Pedro Pasquini, M. Tortola, J. W. F. Valle, PRD **95** (2017) No.3, 033005 [arXiv:1605.01670]
SFG, Alexei Smirnov, JHEP **1610** (2016) 138 [arXiv:1607.08513]
SFG [arXiv:1704.08518]
SFG, Manfred Lindner, PRD **95** (2017) No.3, 033003 [arXiv:1608.01618]

Why neutrino mass & oscillation?

- Higgs boson \Rightarrow electroweak symmetry breaking & mass.
- Chiral symmetry breaking \Rightarrow majority of mass.
- **The world seems not affected by the tiny neutrino mass?**
 - Neutrino mass \Rightarrow Mixing
 - 3 Neutrino \Rightarrow possible **CP violation**
 - CP violation \Rightarrow **Leptogenesis**
 - **Leptogenesis** \Rightarrow **Matter-Antimatter Asymmetry**
 - There is something left in the Universe.
 - Baryogenesis from quark mixing is not enough.
- Majorana $\nu \Leftrightarrow$ **Lepton Number Violation**
- **Residual \mathbb{Z}_2 Symmetries:** $\cos \delta_D = \frac{(s_s^2 - c_s^2 s_r^2)(c_a^2 - s_a^2)}{4c_a s_a c_s s_s s_r}$

1108.0964

1104.0602

ν Oscillation Data

(for NH)	-1σ	Best Value	$+1\sigma$
$\Delta m_s^2 \equiv \Delta m_{12}^2$ (10^{-5}eV^2)	7.37	7.56	7.75
$ \Delta m_a^2 \equiv \Delta m_{13}^2 $ (10^{-3}eV^2)	2.51	2.55	2.59
$\sin^2 \theta_s$ ($\theta_s \equiv \theta_{12}$)	0.305 (33.5°)	0.321 (34.5°)	0.339 (35.6°)
$\sin^2 \theta_a$ ($\theta_a \equiv \theta_{23}$)	0.412 (39.9°)	0.430 (41.0°)	0.450 (42.1°)
$\sin^2 \theta_r$ ($\theta_r \equiv \theta_{13}$)	0.02080 (8.29°)	0.02155 (8.44°)	0.02245 (8.62°)
δ_D, δ_{Mi}	?, ??	?, ??	?, ??

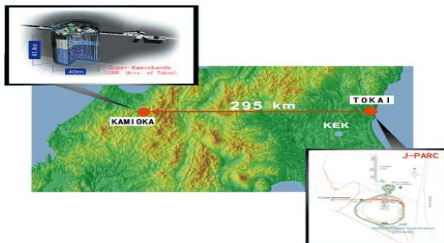
Salas, Forero, Ternes, Tortola & Valle, arXiv:1708.01186

Dirac CP Phase Measurement

Jarah Evslin, **SFG**, Kaoru Hagiwara, JHEP **1602** (2016) 137 [arXiv:1506.05023]
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SFG [arXiv:1704.08518]

CP Measurement @ Accelerator Exps

- T2K



- $\text{NO}\nu\text{A}$



- DUNE/T2KII/T2HK/T2HKK/T2KO; MOMENT/ADS-CI/DAE δ ALUS; Super-PINGU

The Dirac CP Phase δ_D @ Accelerator Exp

- To leading order in $\alpha = \frac{\delta M_{21}^2}{|\delta M_{31}^2|} \sim 3\%$, the oscillation probability relevant to measuring δ_D @ T2(H)K,

$$P_{\nu_\mu \rightarrow \nu_e} \approx 4s_a^2 c_r^2 s_r^2 \sin^2 \phi_{31} - 8c_a s_a c_r^2 s_r c_s s_s \sin \phi_{21} \sin \phi_{31} [\cos \delta_D \cos \phi_{31} \pm \sin \delta_D \sin \phi_{31}]$$

for ν & $\bar{\nu}$, respectively. $[\phi_{ij} \equiv \frac{\delta m_{ij}^2 L}{4E_\nu}]$

- $\nu_\mu \rightarrow \nu_\mu$ Exps measure $\sin^2(2\theta_a)$ precisely, but not $\sin^2 \theta_a$.
- Run both ν & $\bar{\nu}$ modes @ first peak $[\phi_{31} = \frac{\pi}{2}, \phi_{21} = \alpha \frac{\pi}{2}]$,

$$P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e} + P_{\nu_\mu \rightarrow \nu_e} = 2s_a^2 c_r^2 s_r^2,$$

$$P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e} - P_{\nu_\mu \rightarrow \nu_e} = \alpha \pi \sin(2\theta_s) \sin(2\theta_r) \sin(2\theta_a) \cos \theta_r \sin \delta_D.$$

The Dirac CP Phase δ_D @ Accelerator Exp

Accelerator experiment, such as **T2(H)K**, uses off-axis beam to compare ν_e & $\bar{\nu}_e$ appearance @ the oscillation maximum.

- **Disadvantages:**

- **Efficiency:**

- Proton accelerators produce ν more efficiently than $\bar{\nu}$ ($\sigma_\nu > \sigma_{\bar{\nu}}$).
- The $\bar{\nu}$ mode needs more beam time [**$T_{\bar{\nu}} : T_\nu = 2 : 1$**].
- Undercut statistics \Rightarrow Difficult to reduce the uncertainty.

- **Degeneracy:**

- Only **$\sin \delta_D$** appears in $P_{\nu_\mu \rightarrow \nu_e}$ & $P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e}$.
- Cannot distinguish δ_D from $\pi - \delta_D$.

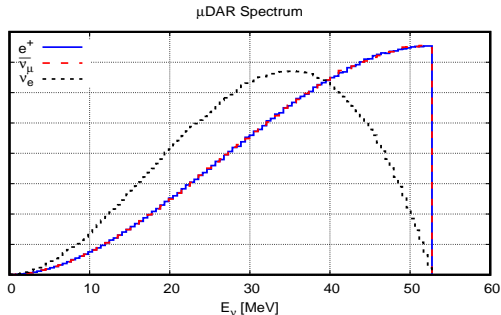
- **CP Uncertainty** $\frac{\partial P_{\mu e}}{\partial \delta_D} \propto \cos \delta_D \Rightarrow \Delta(\delta_D) \propto$ **$1 / \cos \delta_D$** .

- **Solution:**

Measure $\bar{\nu}$ mode with μ^+ decay @ rest (μ DAR)

μ DAR $\bar{\nu}$ Oscillation Experiments

- A cyclotron produces 800 MeV proton beam @ fixed target.
- Produce π^\pm which stops &
 - π^- is absorbed,
 - π^+ decays @ rest: $\pi^+ \rightarrow \mu^+ + \nu_\mu$.
- μ^+ stops & decays @ rest: $\mu^+ \rightarrow e^+ + \bar{\nu}_\mu + \nu_e$.

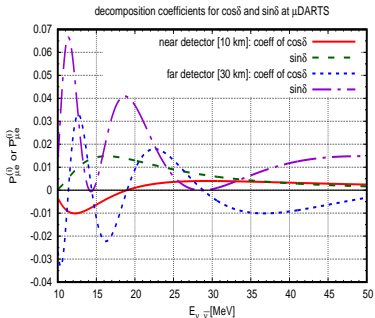
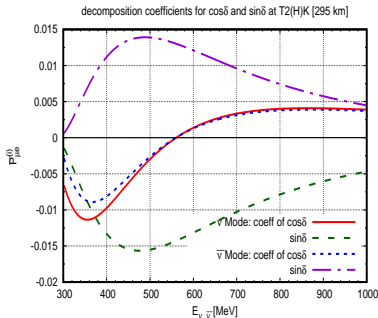


- $\bar{\nu}_\mu$ travel in all directions, oscillating as they go.
- A detector measures the $\bar{\nu}_e$ from $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillation.

Accelerator + μ DAR Experiments

Combining $\nu_\mu \rightarrow \nu_e$ @ accelerator [narrow peak @ 550 MeV] & $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ @ μ DAR [wide peak \sim 45 MeV] solves the 2 problems:

- **Efficiency:**
 - $\bar{\nu}$ @ high intensity, μ DAR is plentiful enough.
 - Accelerator Exps can devote all run time to the ν mode. With same run time, the statistical uncertainty drops by $\sqrt{3}$.
- **Degeneracy:** (**decomposition in propagation basis** [1309.3176])



DAE δ ALUS

- It's the **FIRST** proposal along this line:
 - **3** μ DAR with **3** high-intensity cyclotron complexes.
 - **1** detector.
 - Different baselines: **1.5, 8 & 20** km to break degeneracies.
- **Disadvantages:**
 - The scattering lepton from IBD @ low energy is **isotropic**.
 - **Cannot** distinguish $\bar{\nu}_e$ from different sources
 - Baseline **cannot be measured**.
 - Cyclotrons **cannot** run simultaneously (20~25% duty factor).
 - **Large** statistical uncertainty.
 - **Higher intensity** is necessary.
 - **Expensive** & Technically **challenging**.

New Proposals

1 μ DAR source + 2 detectors

Advantages

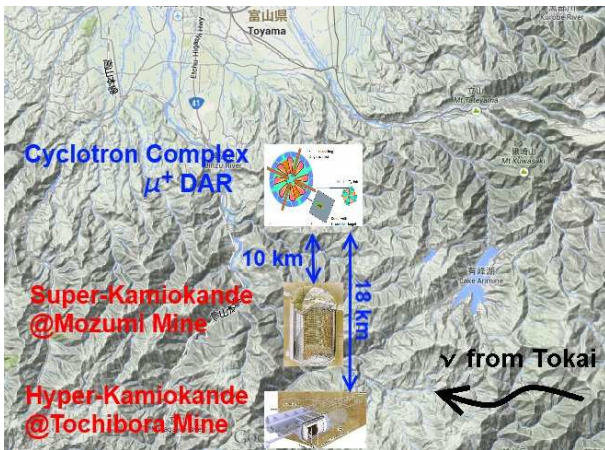
- Full (**100%**) duty factor!
- **Lower** intensity: \sim **9mA** [\sim **4** \times lower than DAE δ ALUS]
- Not far beyond the current state-of-art technology of cyclotron [**2.2mA** @ Paul Scherrer Institute]
- MUCH **cheaper** & technically **easier**.
 - Only one cyclotron.
 - Lower intensity.

Disadvantage?

- A second detector! Tokai 'N Toyama to(2) Kamioka (**TNT2K**) [Evslin, Ge & Hagiwara, 1506.05023]

TNT2K

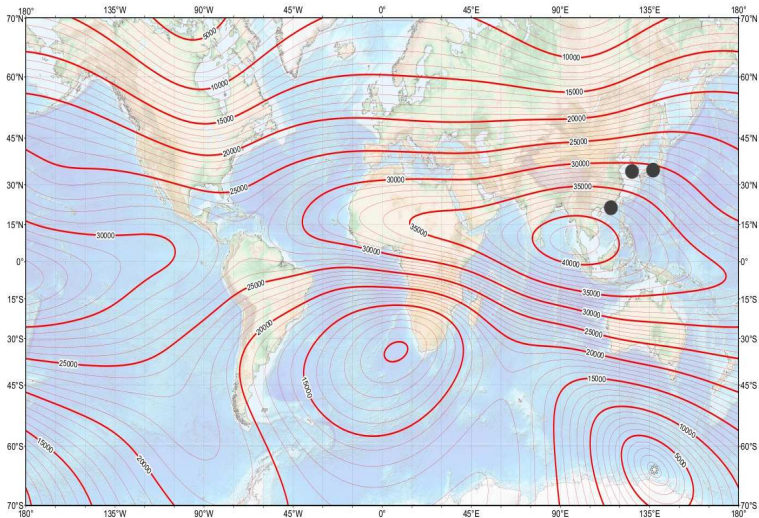
- T2(H)K + μ SK + μ HK



- μ DAR is also useful for **material**, **medicine** industries in Toyama

Lowest Atmospheric Neutrino Background

US/UK World Magnetic Model -- Epoch 2010.0
Main Field Horizontal Intensity (H)

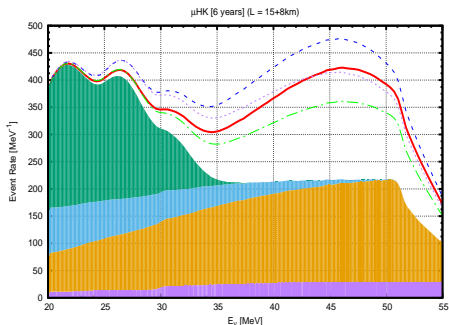
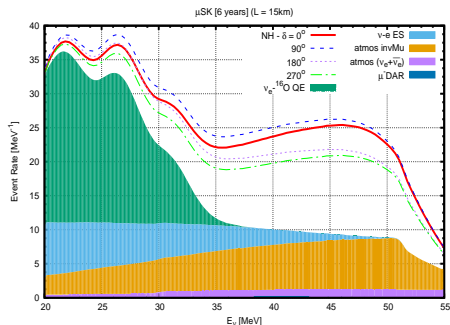


Backgrounds to IBD ($\bar{\nu}_e + p \rightarrow e^+ + n$)

- Reactor $\bar{\nu}_e$: $E_\nu < 10$ MeV
- Accelerator ν_e : $E_\nu > 100$ MeV
- Spallation: $E_\nu \lesssim 20$ MeV
- Supernova Relic Neutrino: $E_\nu \lesssim 20$ MeV

Cut with $30 \text{ MeV} < E_\nu < 55 \text{ MeV}$

- Accelerator $\nu_\mu \rightarrow$ **Invisible muon**
- Atmospheric Neutrino Background
 - **Invisible muon** (below Cherenkov limit)
 - $E_\mu \lesssim 1.5 \times m_\mu$, $\mu^\pm \rightarrow e^\pm$
 - $E_\pi \lesssim 1.5 \times m_\pi$, $\pi^+ \rightarrow \mu^+ \rightarrow e^+$
 - 1 neutron
 - No prompt photon
 - Irreducible $\bar{\nu}_e$: $30 \text{ MeV} \lesssim E_\nu \lesssim 55 \text{ MeV}$
 - Reducible ν_e : $60 \text{ MeV} \lesssim E_\nu \lesssim 100 \text{ MeV}$
 - 1 neutron
 - No prompt photon
 - **Lowest** at μ DARTS & TNT2K sites

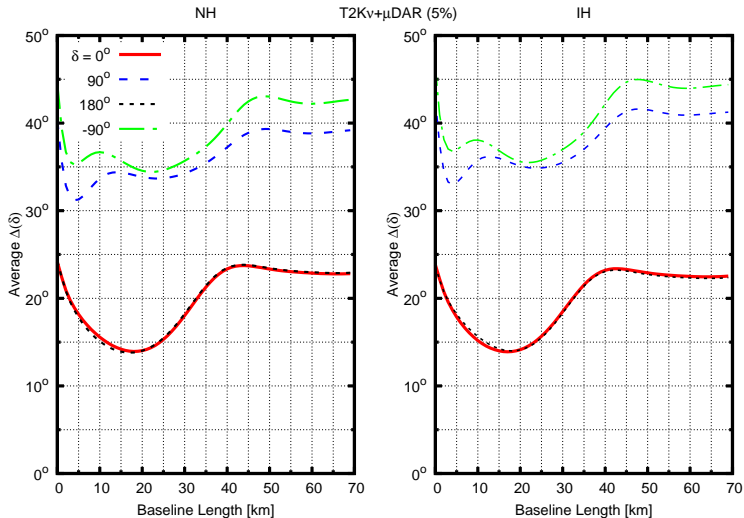


Expected μ DAR IBD signal from 6 yrs of running @ SK (15km) & HK (23km) with NH.

Simulated by [NuPro](http://nupro.hepforge.org/), <http://nupro.hepforge.org/>

δ_D Precision @ TNT2K

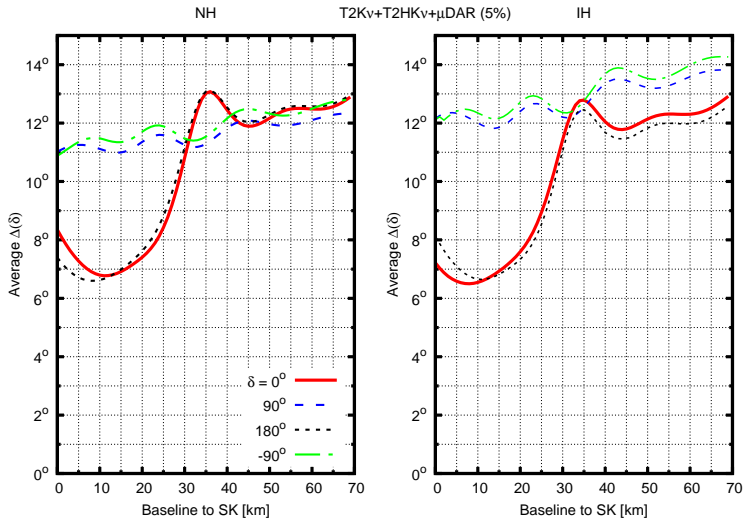
Evslin, Ge & Hagiwara [1506.05023]



Simulated by [NuPro](http://nupro.hepforge.org/), <http://nupro.hepforge.org/>

δ_D Precision @ TNT2K

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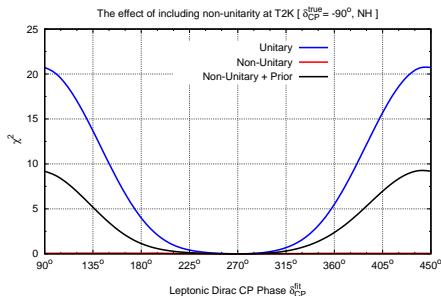
Simulated by [NuPro](http://nupro.hepforge.org/), <http://nupro.hepforge.org/>

Non-Unitarity Mixing (NUM)

Ge, Pasquini, Tortola & Valle [1605.01670]

$$N = N^{NP} U = \begin{pmatrix} \alpha_{11} & 0 & 0 \\ |\alpha_{21}| e^{i\phi} & \alpha_{22} & 0 \\ \alpha_{31} & \alpha_{32} & \alpha_{33} \end{pmatrix} U.$$

$$P_{\mu e}^{NP} = \alpha_{11}^2 \left\{ \alpha_{22}^2 \left[c_a^2 |S'_{12}|^2 + s_a^2 |S'_{13}|^2 + 2c_a s_a (\cos \delta_D \mathbb{R} - \sin \delta_D \mathbb{I})(S'_{12} S'_{13}^*) \right] + |\alpha_{21}|^2 P_{ee} \right. \\ \left. + 2\alpha_{22} |\alpha_{21}| \left[c_a (c_\phi \mathbb{R} - s_\phi \mathbb{I})(S'_{11} S'_{12}^*) + s_a (c_{\phi+\delta_D} \mathbb{R} - s_{\phi+\delta_D} \mathbb{I})(S'_{11} S'_{13}^*) \right] \right\}.$$



NUM vs Seesaw Mechanism

- Heavy neutrinos

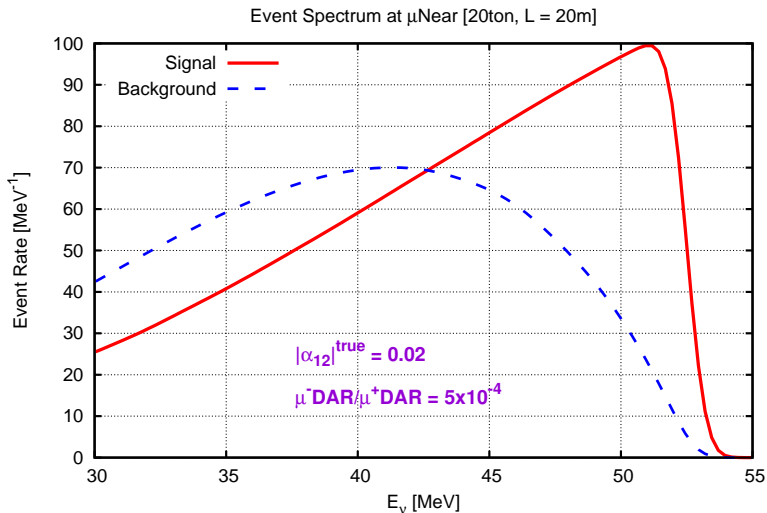
$$\bar{\nu} M_D \mathcal{N} + h.c. + \bar{\mathcal{N}} M_N \mathcal{N} = \begin{pmatrix} \bar{\nu} & \bar{\mathcal{N}} \end{pmatrix} \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix} \begin{pmatrix} \nu \\ \mathcal{N} \end{pmatrix}$$

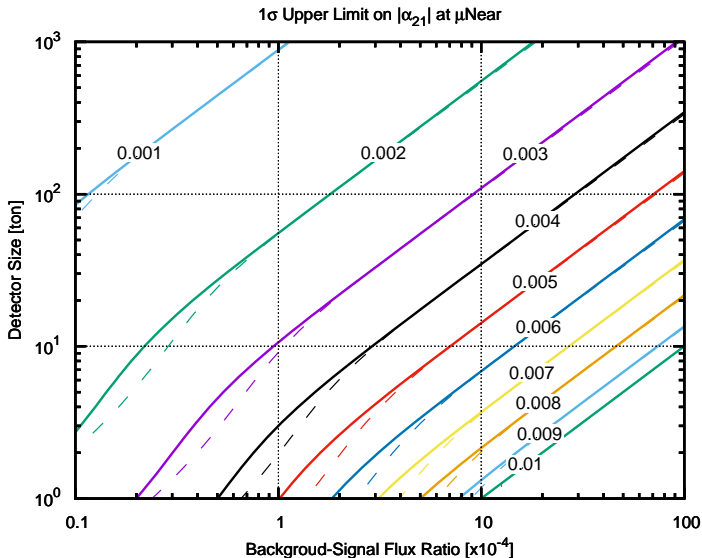
- Seesaw Mechanism

$$M_\nu = -M_D M_N^{-1} M_D^T, \quad \nu' = \nu + M_D M_N^{-1} \mathcal{N}$$

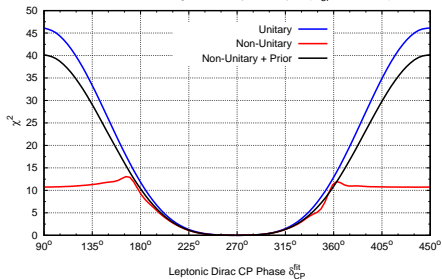


$$P_{\mu e}^{NP}(L \rightarrow 0) = \alpha_{11}^2 |\alpha_{21}|^2 P_{ee} \approx \alpha_{11}^2 |\alpha_{21}|^2 \approx |\alpha_{21}|^2$$

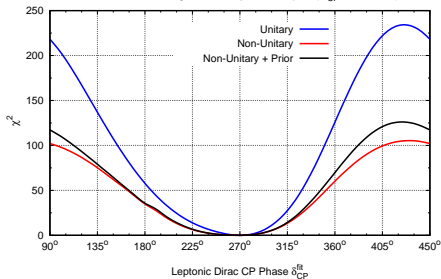




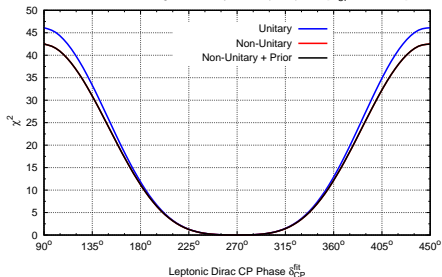
The effect of including non-unitarity at T2K+ μ SK [$\delta_{CP}^{true} = -90^\circ$, NH]



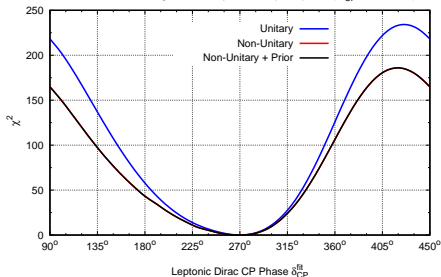
The effect of including non-unitarity at T2HK+ μ HK [$\delta_{CP}^{true} = -90^\circ$, NH]



The effect of including non-unitarity at T2K+ μ SK+ μ Near [$\delta_{CP}^{true} = -90^\circ$, NH]



The effect of including non-unitarity at T2HK+ μ HK+ μ Near [$\delta_{CP}^{true} = -90^\circ$, NH]

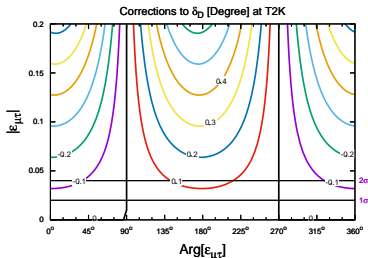
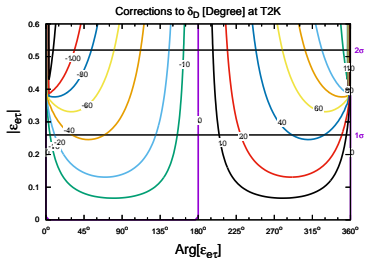
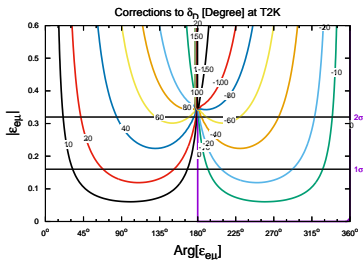
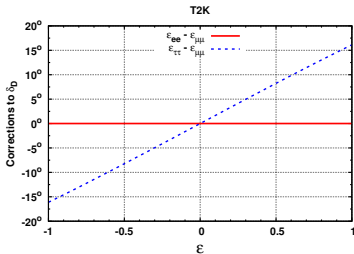


$$\mathcal{H} \equiv \frac{1}{2\mathbf{E}_\nu} U \begin{pmatrix} 0 & & \\ & \Delta m_s^2 & \\ & & \Delta m_a^2 \end{pmatrix} U^\dagger + V_{cc} \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix}$$

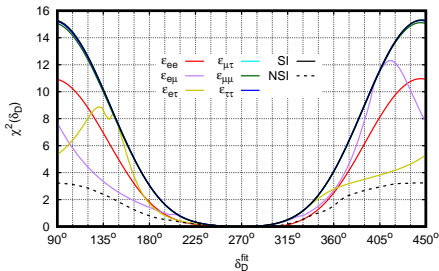
- Standard Interaction – V_{cc} (also V_{nc})
- Non-Standard Interaction – $\epsilon_{\alpha\beta}$
 - **Diagonal** $\epsilon_{\alpha\alpha}$ are real
 - **Off-diagonal** $\epsilon_{\alpha\neq\beta}$ are complex
 - **Both can fake CP**
- Z' in LMA-Dark model with $L_\mu - L_\tau$ gauged as $U(1)$
 - $M_{Z'} \sim \mathcal{O}(10)\text{MeV}$
 - $g_{Z'} \sim 10^{-5}$

Faked CP with NSI

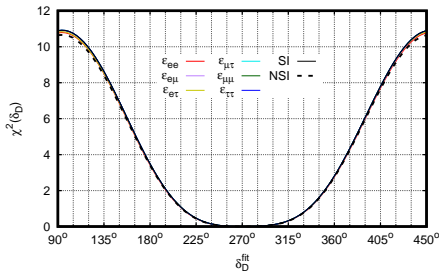
SFG & Alexei Smirnov [arXiv:1607.08513]



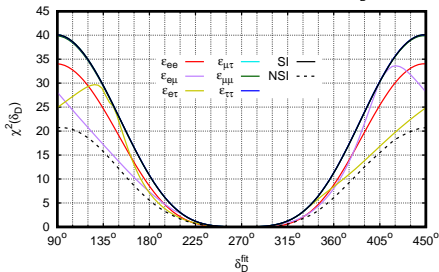
The effect of NSI on the CP sensitivity at T2K [$\delta_D^{\text{true}} = -90^\circ$]



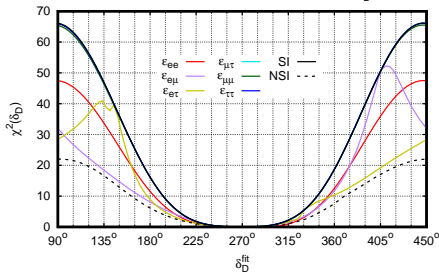
The effect of NSI on the CP sensitivity at μ SK [$\delta_D^{\text{true}} = -90^\circ$]



The effect of NSI on the CP sensitivity at T2K+ μ SK [$\delta_D^{\text{true}} = -90^\circ$]



The effect of NSI on the CP sensitivity at ν T2K+ μ SK [$\delta_D^{\text{true}} = -90^\circ$]



Summary (1)

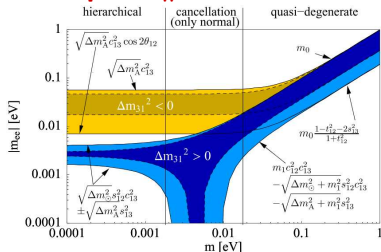
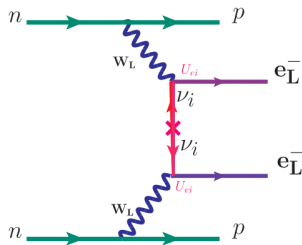
- **Better CP measurement than T2K**
 - Much larger event numbers
 - Much better CP sensitivity around maximal CP
 - Solve degeneracy between δ_D & $\pi - \delta_D$
 - Guarantee CP sensitivity against NUM
 - Guarantee CP sensitivity against NSI
- **Better configuration than DAE δ ALUS**
 - Only one cyclotron
 - 100% duty factor
 - Much lower flux intensity
 - Much easier
 - Much cheaper
 - Single near detector

Majorana CP Phase Measurement

SFG, Manfred Lindner, PRD **95** (2017) No.3, 033003 [arXiv:1608.01618]

$0\nu 2\beta$ Decay

- Mediated by **Majorana Neutrino** + **Lepton # Violation**



- Helicity Suppression** → **Mass Suppression**

$$\mathcal{M} \propto \sum_i U_{ei} \frac{i}{\not{p} - m_i} U_{ei} \approx \sum_i U_{ei} \frac{m_i}{p^2} U_{ei}$$

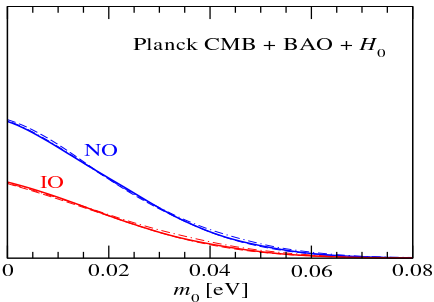
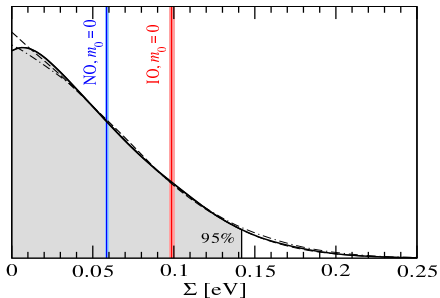
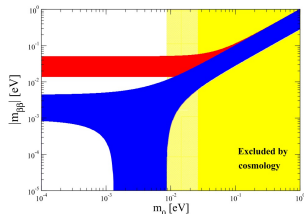
- Effective Electron Neutrino Mass**

$$\langle m \rangle_{ee} \equiv \left| \sum_i m_i U_{ei}^2 \right| = \left| c_s^2 c_r^2 m_1 e^{i\delta_{M1}} + s_s^2 c_r^2 m_2 + s_r^2 m_3 e^{i\delta_{M3}} \right|$$

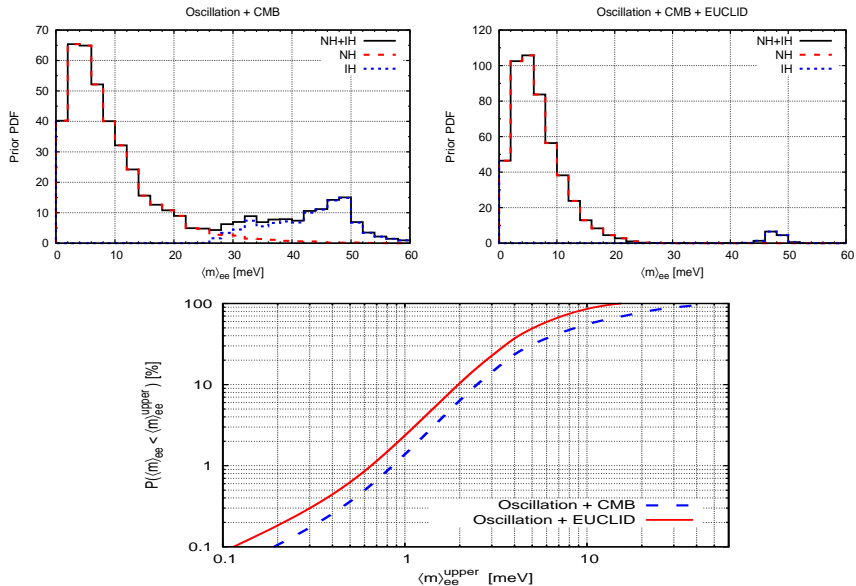
see also SFG & Werner Rodejohann [1507.05514]

Cosmological Data on Mass Sum

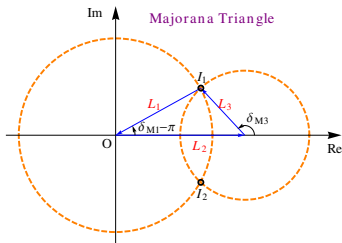
$$\Sigma \equiv m_1 + m_2 + m_3$$



Preference of NH \Rightarrow Non-Observation of $0\nu 2\beta$?



Any chance of obtaining some information?



$$\langle m \rangle_{ee} \equiv \vec{L}_1 + \vec{L}_2 + \vec{L}_3,$$

with

$$\vec{L}_1 \equiv m_1 U_{e1}^2 = m_1 c_r^2 c_s^2 e^{i\delta_{M1}},$$

$$\vec{L}_2 \equiv m_2 U_{e2}^2 = \sqrt{m_1^2 + \Delta m_s^2} c_r^2 s_s^2,$$

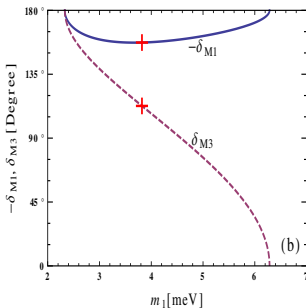
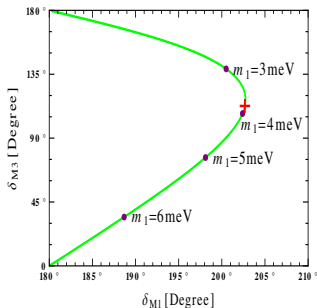
$$\vec{L}_3 \equiv m_3 U_{e3}^2 = \sqrt{m_1^2 + \Delta m_a^2} s_r^2 e^{i\delta_{M3}}.$$

Determine 2 Majorana Phases Simultaneously

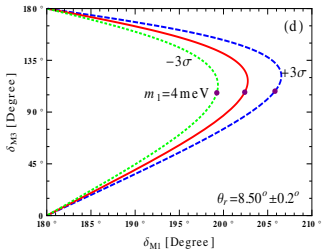
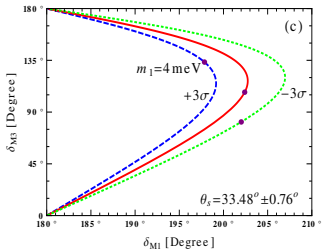
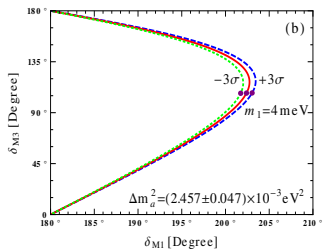
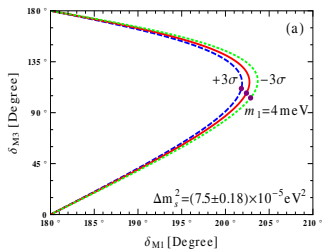
$$|L_1 - L_3| \leq L_2 \leq L_1 + L_3.$$

$$\cos \delta_{M1} = -\frac{L_1^2 + L_2^2 - L_3^2}{2L_1L_2} = -\frac{m_1^2 c_r^4 c_s^4 + m_2^2 c_r^4 s_s^4 - m_3^2 s_r^4}{2m_1 m_2 c_r^4 c_s^2 s_s^2},$$

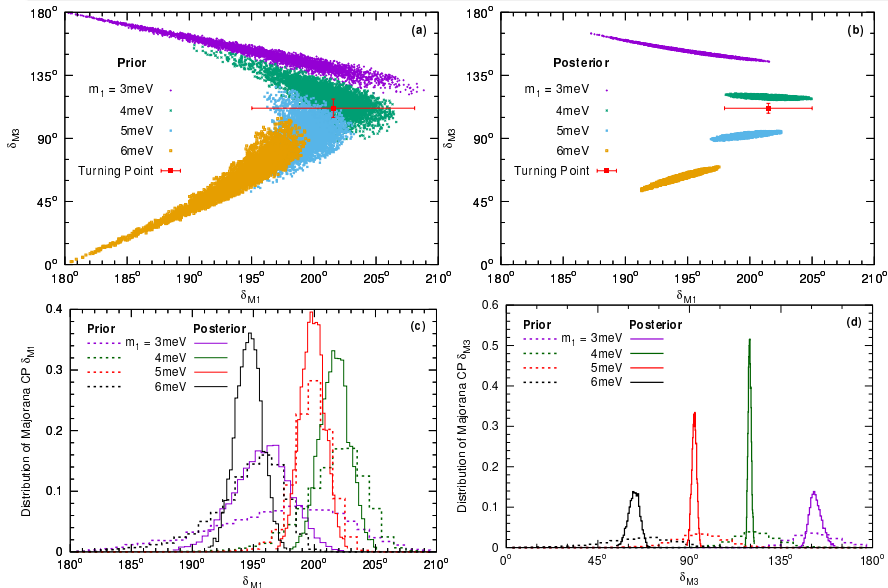
$$\cos \delta_{M3} = +\frac{L_1^2 - L_2^2 - L_3^2}{2L_2L_3} = +\frac{m_1^2 c_r^4 c_s^4 - m_2^2 c_r^4 s_s^4 - m_3^2 s_r^4}{2m_2 m_3 c_r^2 s_r^2 s_s^2}.$$



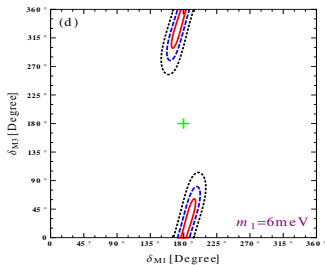
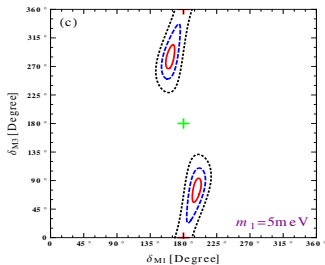
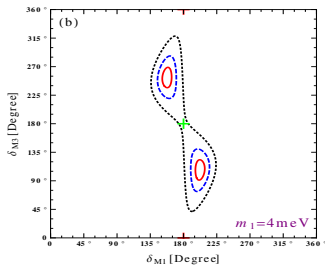
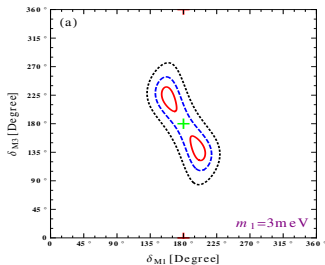
Uncertainties from Oscillation Parameters



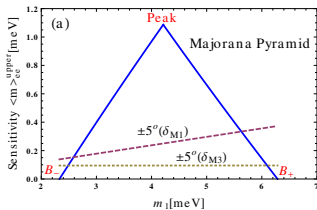
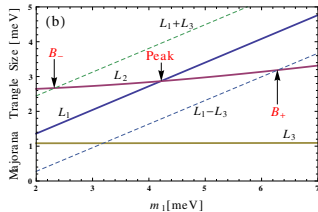
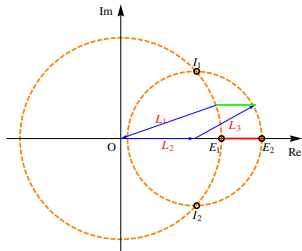
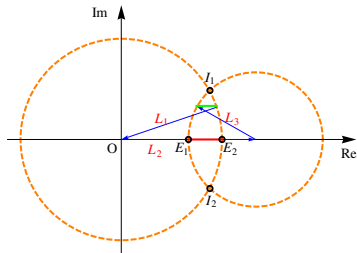
Uncertainties from Oscillation Parameters



Uncertainties from $\langle m \rangle_{ee}$



Majorana Pyramid



Prey of Leptonic CP Phases



Dirac



Majorana 1



Majorana 2



Majorana Pyramid

- **CP is fundamentally important**
 - Matter-Antimatter Asymmetry
 - Test of Discrete Symmetries
- **Better Dirac CP measurement at TNT2K**
 - Much larger event numbers
 - Much better CP sensitivity around maximal CP
 - Solve degeneracy between δ_D & $\pi - \delta_D$
 - Guarantee CP sensitivity against NUM
 - Guarantee CP sensitivity against NSI
- **Elusive Majorana CP Phases**
 - Observing nothing is not bad at all
 - Even better: fixing two Majorana CP phases simultaneously

Thank You!