

Spectrum prediction in non-minimal supersymmetric models

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Motivation – minimal supersymmetry (MSSM)

MSSM gauge group:

$$SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$$

MSSM particle content:

(S)Quarks:	$Q_i, \bar{U}_i, \bar{D}_i$
(S)Leptons:	L_i, \bar{E}_i
Higgs(inos):	H_1, H_2
Gauge(inos):	V_g^a, \vec{V}^W, V^Y

MSSM superpotential:

$$\mathcal{W}_{\text{MSSM}} = \mu H_1 H_2 - y_u (Q H_2) \bar{U} - y_d (H_1 Q) \bar{D} - y_e (H_1 L) \bar{E}$$

Motivation – μ problem of the MSSM

MSSM superpotential:

$$\mathcal{W}_{\text{MSSM}} = \mu H_1 H_2 - y_u(QH_2)\bar{U} - y_d(H_1 Q)\bar{D} - y_e(H_1 L)\bar{E}$$

Problem:

- $\mathcal{W}_{\text{MSSM}}$ generated at unification scale $M_X \Rightarrow \mu \sim M_X$
- but EWSB conditions imply

$$\frac{1}{2}M_Z^2 = \frac{m_{H_1}^2 - m_{H_2}^2 \tan^2 \beta}{\tan^2 \beta - 1} - \mu^2 \quad \Rightarrow \quad \mu \sim M_Z$$

Solution: introduce gauge singlet $S \rightarrow S + v_s$

$$\mathcal{W}_{\text{NMSSM}} = \mathcal{W}_{\text{MSSM}}(\mu = 0) + \lambda S H_1 H_2 + \kappa S^3$$

Motivation – m_h fine tuning in the MSSM

$$\text{MSSM: } m_h^2 = (M_Z c_{2\beta})^2 + \Delta m_h^2$$

$$\text{NMSSM: } m_h^2 = (M_Z c_{2\beta})^2 + (\lambda v_s s_{2\beta})^2 + \Delta m_h^2$$

$$\text{E}_6\text{SSM: } m_h^2 = (M_Z c_{2\beta})^2 + (\lambda v_s s_{2\beta})^2 + \frac{M_Z^2}{4} \left(1 + \frac{c_{2\beta}}{4}\right)^2 + \Delta m_h^2$$

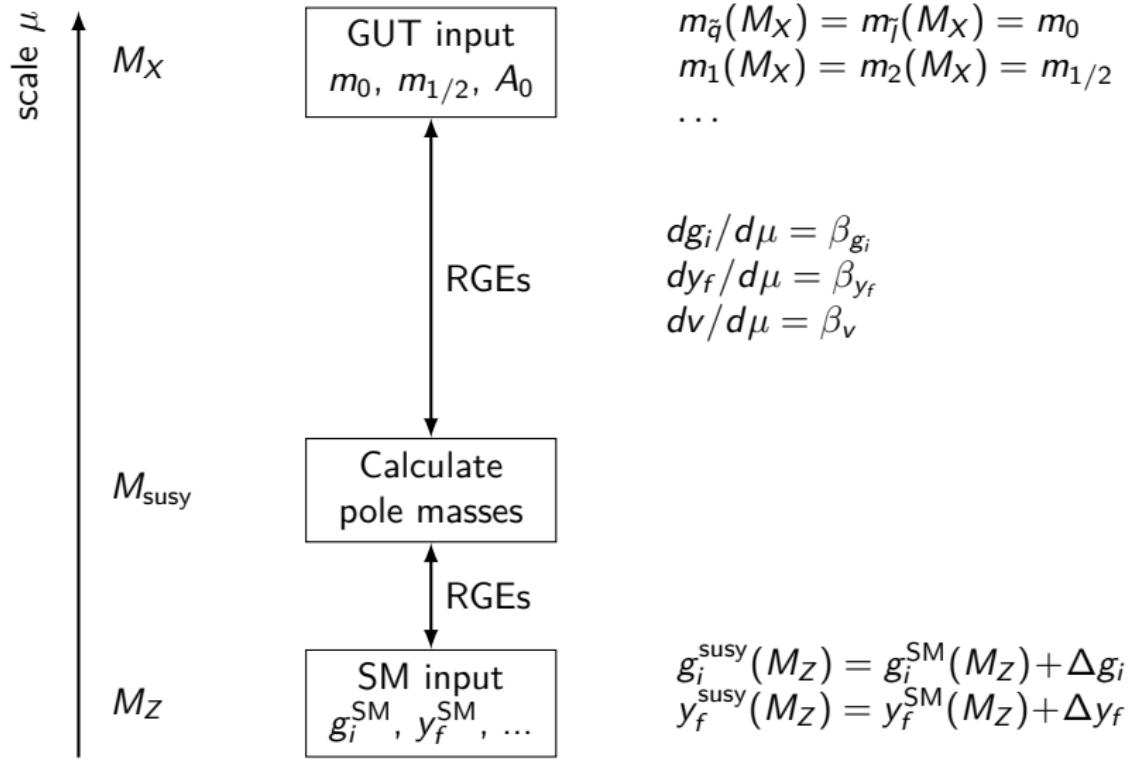
⇒ Non-minimal supersymmetric models can increase the tree-level Higgs mass

⇒ avoid m_h fine tuning

Motivation – non-minimal supersymmetric models

	CMSSM	non-minimal SUSY	CE ₆ SSM
solves μ problem of the MSSM	✗	✓	✓ [hep-ph/0510419]
avoid m_h fine tuning	✗	✓	✓ [hep-ph/1302.5291]
connection to specific gravity models	✗	✓	✓ [hep-ph/0510419]
complete repres. of simple GUT gauge group	✗	✓	✓ [hep-ph/0510419]
new particles and phenomena	✓	✓	✓ [hep-ph/1302.5291]
explanation of $(g - 2)_\mu^{\text{exp.}}$	✗	✓	✗

Physics situation



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Current status of β_v calculations

Current status of $\beta_v^{\overline{\text{DR}}/\overline{\text{MS}}}$:

Model	$\beta_v^{(1)}$	$\beta_v^{(2)}$
MSSM	✓ [Chankowski Nucl.Phys. B423]	✓ [Yamada 94] $O(g^2 Y^2)$
E_6 SSM	✓ [Athron et al. 12]	✗
\forall gauge theory	?	✗
\forall SUSY model	?	✗

We calculate:

$\delta v^{(1,2)}$ and $\beta_v^{(1,2)}$ in a general and supersymmetric gauge theory
with R_ξ gauge fixing [Sperling et al. 13]

Renormalization of v

Spontaneously broken gauge theory:

$$\phi \rightarrow \phi + v$$

Most generic renormalization transformation:

$$(\phi + v) \rightarrow \sqrt{Z}\phi + v + \delta v$$

or $(\phi + v) \rightarrow \sqrt{Z}(\phi + v + \delta \bar{v})$

$\delta \bar{v}$ characterizes to what extent v renormalizes differently from ϕ .

With $\sqrt{Z} = 1 + \frac{1}{2}\delta Z \Rightarrow$

$$\delta v = \frac{1}{2}\delta Z v + \delta \bar{v}$$

Influence of global gauge invariance

When does $\delta\bar{v}$ appear?

$$\begin{array}{lll} \text{global gauge invariance} & \Rightarrow & \delta\bar{v} = 0 \\ \text{no global gauge invariance} & \Rightarrow & \delta\bar{v} \neq 0 \end{array}$$

R_ξ gauge fixing:

$$\begin{aligned} \mathcal{L}_{\text{fix,gh}} &= s \left[\bar{c}^A \left(F^A + \xi B^A / 2 \right) \right] \\ F^A &= \partial^\mu V_\mu^A + ig\xi v_a T_{ab}^A \phi_b \end{aligned}$$

R_ξ breaks global gauge invariance for $\xi \neq 0 \Rightarrow \delta\bar{v} \neq 0$.

Background field method – Introducing $\hat{\phi}$

Problem: R_ξ breaks global gauge invariance for $\xi \neq 0 \Rightarrow \delta\bar{v} \neq 0$.

Trick: Keep global gauge invariance in intermediate calculation!

[Kraus,Sibold 95]

Introduce background field $\hat{\phi}$ and shift \hat{v}

$$\phi \rightarrow \phi_{\text{eff}} := \phi + \hat{\phi} + \hat{v}$$

where $\hat{\phi} + \hat{v}$ has same gauge transformation as ϕ .

Modified R_ξ gauge fixing:

$$F^A = \partial^\mu V_\mu^A + ig\xi(\hat{\phi} + \hat{v})_a T_{ab}^A \phi_b$$

\Rightarrow global gauge invariance! $\Rightarrow \delta\bar{v} = 0$

Background field method – Lagrangian

Modified BRS transformations:

$$s\phi_{\text{eff}} = -igT^A c^A \phi_{\text{eff}}, \quad s\hat{\phi} = \hat{q}, \quad s\hat{q} = 0, \quad s\phi = s\phi_{\text{eff}} - \hat{q}$$

Modified Lagrangian:

$$\mathcal{L} = \mathcal{L}_{\text{inv}}|_{\phi \rightarrow \phi_{\text{eff}}} + \mathcal{L}_{\text{fix,gh}} + \mathcal{L}_{\text{ext}}$$

$$\mathcal{L}_{\text{inv}} = -\frac{1}{4}F_{\mu\nu}^A F^{A\mu\nu} + \frac{1}{2}(D_\mu\phi)_a(D^\mu\phi)_a + i\psi_p^\alpha\sigma_{\alpha\dot{\alpha}}^\mu(D_\mu^\dagger\bar{\psi}^{\dot{\alpha}})_p$$

$$-\frac{1}{2!}m_{ab}^2\phi_a\phi_b - \frac{1}{3!}h_{abc}\phi_a\phi_b\phi_c - \frac{1}{4!}\lambda_{abcd}\phi_a\phi_b\phi_c\phi_d$$

$$-\frac{1}{2}\left[(m_f)_{pq}\psi_p^\alpha\psi_{q\alpha} + \text{h.c.}\right] - \frac{1}{2}\left[Y_{pq}^a\psi_p^\alpha\psi_{q\alpha}\phi_a + \text{h.c.}\right]$$

$$\mathcal{L}_{\text{fix,gh}} = s\left[\bar{c}^A\left(F^A + \xi B^A/2\right)\right]$$

$$\mathcal{L}_{\text{ext}} = K_{\phi_a}s\phi_a + K_{V_\mu^A}sV_\mu^A + K_{c^A}sc^A + [K_{\psi_p}s\psi_p + \text{h.c.}]$$

Background field method – Renormalization of ν

Modified renormalization transformation:

$$\phi \rightarrow \sqrt{Z}\phi$$

$$(\hat{\phi} + \hat{\nu}) \rightarrow \sqrt{Z}\sqrt{\hat{Z}}(\hat{\phi} + \hat{\nu})$$

$$\hat{q} \rightarrow \sqrt{Z}\sqrt{\hat{Z}}\hat{q}$$

$$\Rightarrow \phi_{\text{eff}} \rightarrow \sqrt{Z} \left(\phi + \sqrt{\hat{Z}} (\hat{\phi} + \hat{\nu}) \right)$$

Standard approach:

$$(\phi + \nu) \rightarrow \sqrt{Z}\phi + \nu + \delta\nu$$

For $\hat{\phi} = 0 \Rightarrow$

$$\boxed{\delta\nu = \left(\sqrt{Z}\sqrt{\hat{Z}} - 1 \right) \nu = \frac{1}{2} (\delta Z + \delta \hat{Z}) \nu + O(\hbar^2)}$$

Calculation of $\delta\hat{Z}$ – 1 Loop I

$$\mathcal{L}_{\text{ext}} = K_\phi s\phi + \dots$$

$$= -K_\phi igT^A c^A \phi - K_\phi \hat{q} + \dots$$

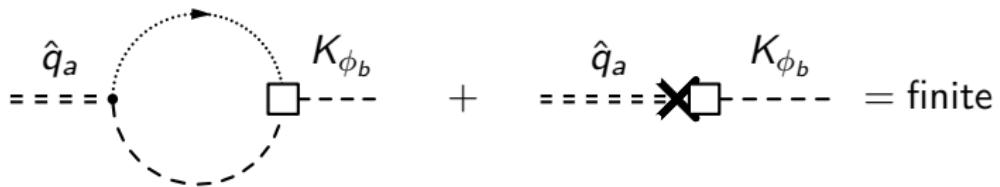
$$\mathcal{L}_{\text{fix,gh}} = s \left[\bar{c}^A \left(F^A + \xi B^A / 2 \right) \right]$$

\Rightarrow

$$\hat{q}_a \times \square \xrightarrow{K_{\phi_b}} = -\frac{i}{2} \delta\hat{Z}_{ba}^{(1)}$$

$$\hat{q}_a \times \square \xrightarrow{\bar{c}^A} = \xi g T_{ab}^A, \quad \square \xrightarrow{K_{\phi_a}} \bar{c}^A \times \square \xrightarrow{\phi_b} = g T_{ab}^A$$

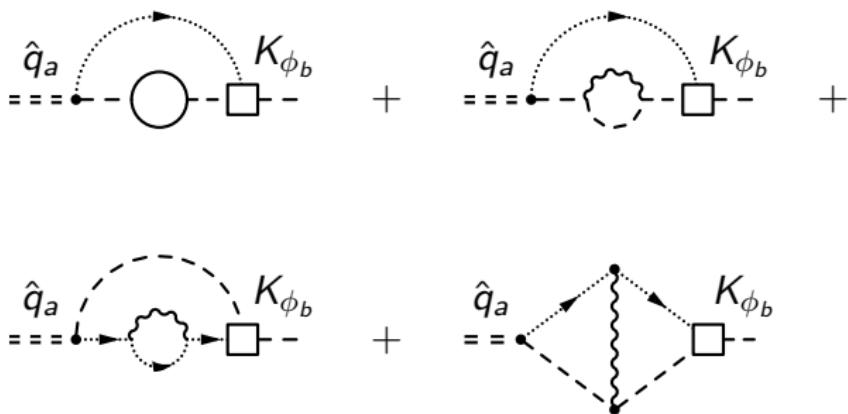
Calculation of $\delta\hat{Z}$ – 1 Loop II



\Rightarrow

$$(4\pi)^2 \delta\hat{Z}^{(1)} = 2g^2\xi C^2(S) \frac{1}{\epsilon}$$

Calculation of $\delta\hat{Z}$ – 2 Loop



\Rightarrow

$$(4\pi)^4 \delta\hat{Z}^{(2)} = g^2 \xi C^2(S) Y^2(S) \left(\frac{1}{\epsilon^2} - \frac{1}{\epsilon} \right) + O(g^4)$$

Result – β_v

$$\delta v = \left(\sqrt{Z} \sqrt{\hat{Z}} - 1 \right) v$$

$$\gamma(S) = \left(\mu \partial_\mu \sqrt{Z}^{-1} \right) \sqrt{Z}, \quad \hat{\gamma}(S) = \left(\mu \partial_\mu \sqrt{\hat{Z}}^{-1} \right) \sqrt{\hat{Z}}$$

\Rightarrow

$$\beta_v^{(n)} = \left[\gamma^{(n)}(S) + \hat{\gamma}^{(n)}(S) \right] v$$

$$\hat{\gamma}^{(1)}(S) = \frac{\xi}{(4\pi)^2} 2g^2 C^2(S)$$

$$\begin{aligned} \hat{\gamma}^{(2)}(S) = & \frac{\xi}{(4\pi)^4} \left\{ g^4 \left[2(1+\xi) C^2(S) C^2(S) + \frac{7-\xi}{2} C_2(G) C^2(S) \right] \right. \\ & \left. - 2g^2 C^2(S) Y^2(S) \right\} \end{aligned}$$

Calculation of β_v – Summary

- In R_ξ gauge fixings (with $\xi \neq 0$), v renormalizes differently from ϕ

$$\phi \rightarrow \sqrt{Z}(\phi + v + \delta\bar{v})$$

- difference $\delta\bar{v}$ can be interpreted as field renormalization $\delta\hat{Z}$ of a background field $\hat{\phi}$

$$\delta\bar{v} = \frac{1}{2}\delta\hat{Z}v$$

- We've calculated $\delta\hat{Z}^{(1,2)}$, $\delta v^{(1,2)}$, $\beta_v^{(1,2)}$, in the $\overline{\text{DR}}$ scheme for arbitrary ξ in a general gauge theory [Sperling et al. 13]

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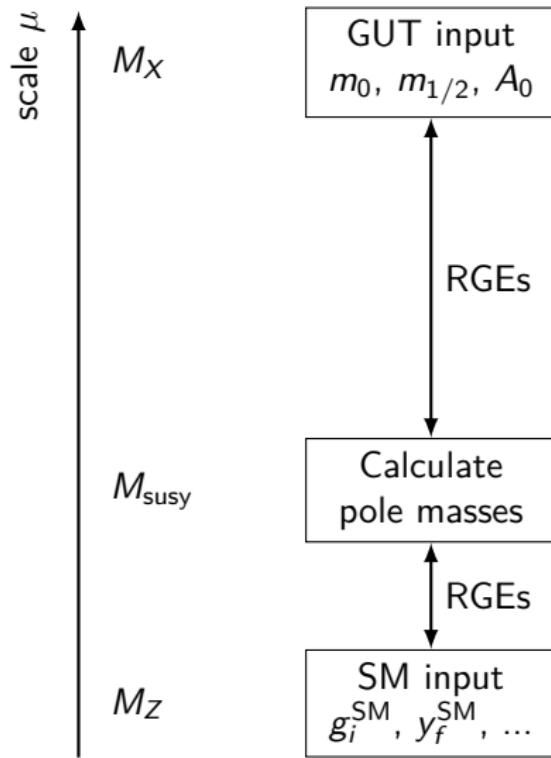
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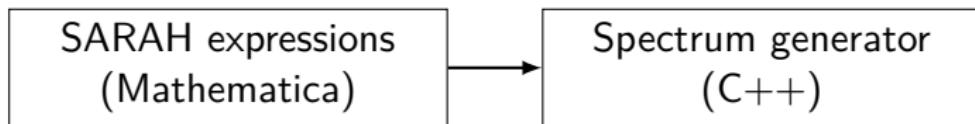
$$m_{\tilde{q}}(M_X) = m_{\tilde{l}}(M_X) = m_0$$
$$m_1(M_X) = m_2(M_X) = m_{1/2}$$
$$\dots$$

$$dg_i/d\mu = \beta_{g_i}$$
$$dy_f/d\mu = \beta_{y_f}$$
$$dv/d\mu = \beta_v$$

$$g_i^{\text{susy}}(M_Z) = g_i^{\text{SM}}(M_Z) + \Delta g_i$$
$$y_f^{\text{susy}}(M_Z) = y_f^{\text{SM}}(M_Z) + \Delta y_f$$
$$\dots$$

Motivation – Why creating a new spectrum generator?

FlexibleSUSY provides Mathematica meta code which creates a spectrum generator

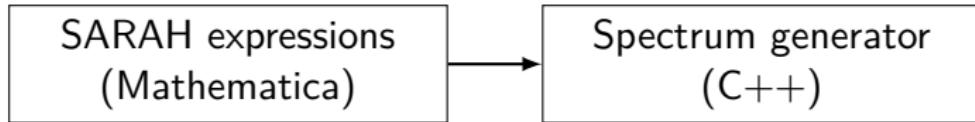


Already existing: SARAH → SPheno (Fortran)

Motivation for FlexibleSUSY:

- large variety of supersymmetric models
 - user customization desired
- convergence problems in certain parameter regions
 - provide alternative RG solvers

Design goals



Design goals:

- modular, object oriented C++ code ⇒ easy to hack! ✓
- fast (smart linear algebra, multithreading) ✓
- multiple RGE solvers:
 - two-scale running (adaptive Runge-Kutta) ✓
 - lattice method + variants (Jae-hyeon Park) ✓
- SARAH-like user interface ✓
- tower of effective field theories X

Usage – 0. Get the source code

Get the source code:

```
$ git clone https://github.com/Expander/FlexibleSUSY  
$ cd FlexibleSUSY
```

Usage – 1. Create FlexibleSUSY model file I

```
FSModelName = "NMSSM";  
  
MINPAR = { {1, m0},  
           {2, m12},  
           {3, TanBeta},  
           {5, Azero} };  
  
EXTPAR = { {61, LambdaInput} };  
  
EWSBOutputParameters = { Kappa, vS, ms2 };  
  
SUSYScale = Sqrt[M[Su[1]]*M[Su[6]]];  
  
HighScale = g1 == g2;  
  
HighScaleInput = {  
    {mHd2, m0^2}, {mHu2, m0^2}, {mq2, UNITMATRIX[3] m0^2},  
    ...  
};  
  
LowScale = SM[MZ];  
  
LowScaleInput = { ... };
```

Usage – 1. Create FlexibleSUSY model file II

Minimizer (iterative):

```
LowScaleInput = {  
    FSMinimize [{vd,vu},  
                (SM[MZ] - Pole[M[VZ]])^2 /  
                STANDARDDEVIATION[MZ]^2  
                + (SM[MH] - Pole[M[hh[1]]])^2 /  
                STANDARDDEVIATION[MH]^2 ]  
};
```

Root finder (iterative):

```
LowScaleInput = {  
    FSFindRoot [{vd,vu}, {SM[MZ] - Pole[M[VZ]],  
                 SM[MH] - Pole[M[hh[1]]]} ]  
};
```

Usage – 2. Configure and make

Create a NMSSM spectrum generator:

```
$ ./createmodel --models=NMSSM  
$ ./configure --with-models=NMSSM  
$ make
```

Run it:

```
$ ./models/NMSSM/run_NMSSM.x
```

Run with an SLHA input file:

```
$ ./models/NMSSM/run_NMSSM.x \  
--slha-input-file=input.slha \  
--slha-output-file=output.slha
```

Modularity – generated C++ code

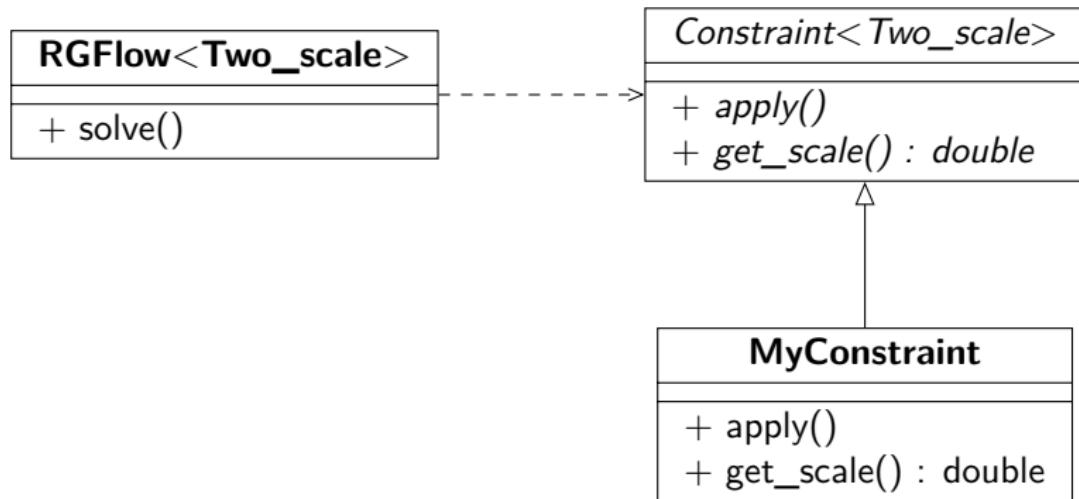
```
typedef Two_scale T; // or Lattice
NMSSM<T> nmssm;
NMSSM_input_parameters input;

std::vector<Constraint<T>*> constraints = {
    new NMSSM_low_scale_constraint<T>(input),
    new NMSSM_susy_scale_constraint<T>(input),
    new NMSSM_high_scale_constraint<T>(input)
};

// solve RG eqs. with the above constraints
RGFlow<T> solver;
solver.add_model(&nmssm, constraints);
solver.solve();

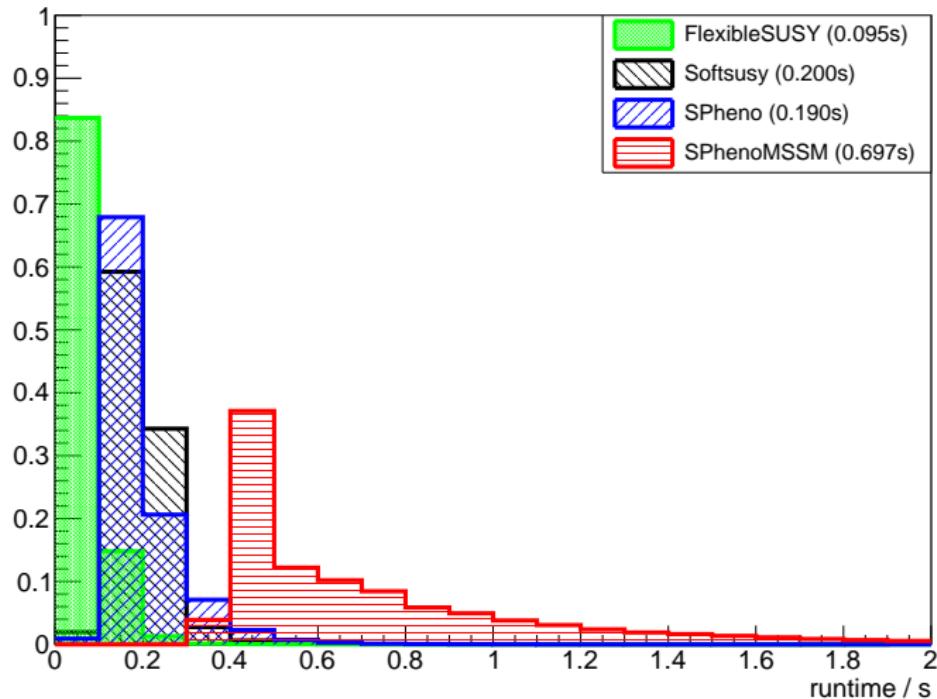
nmssm.calculate_spectrum();
```

Modularity – Constraint interface



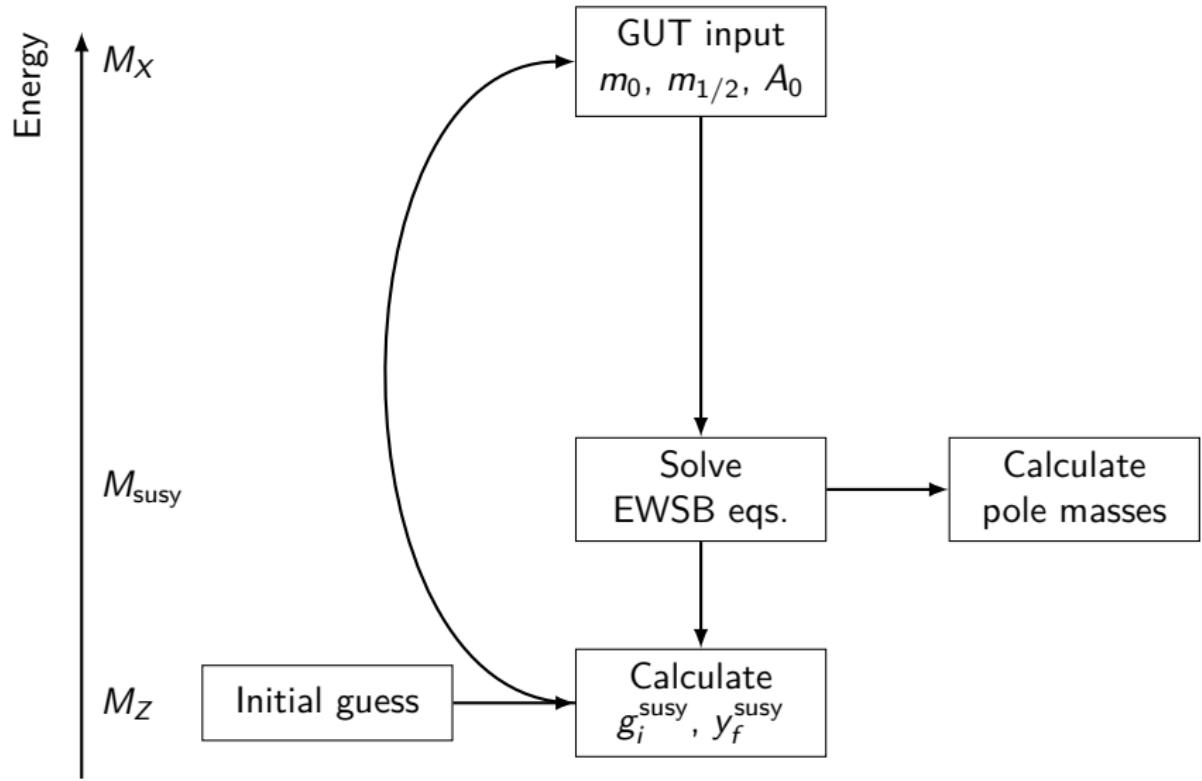
Speed – runtime comparison

CMSSM spectrum generator runtime



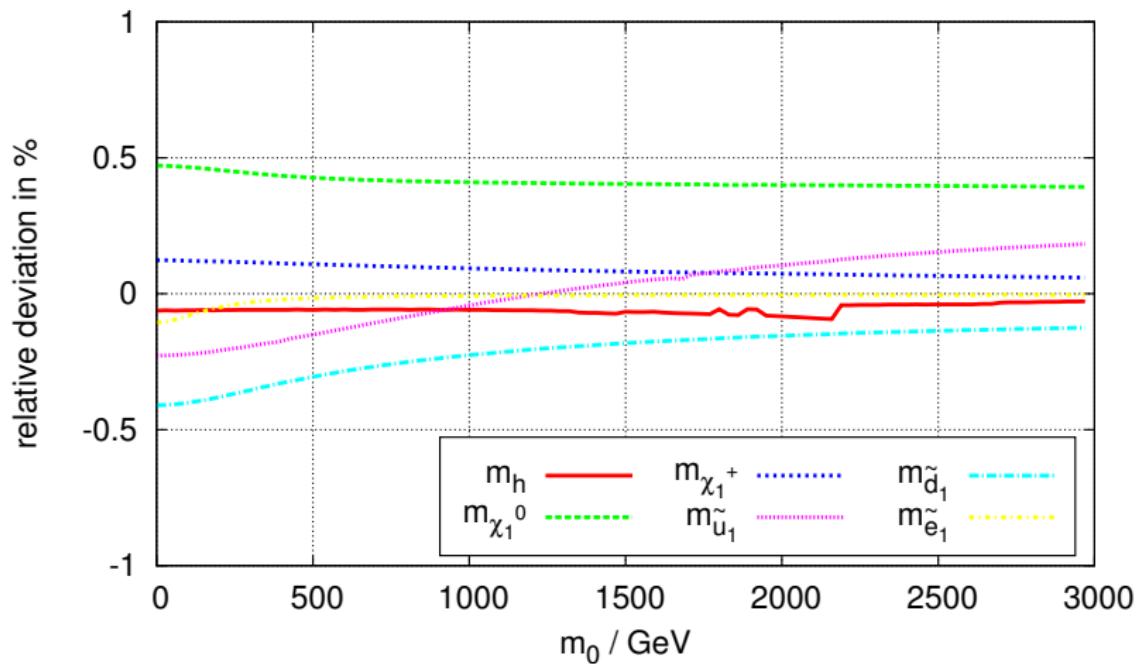
g++ 4.8.0, gfortran 4.8.0

RGE solvers – two-scale algorithm



Spectrum comparison FlexibleSUSY vs. Softsusy

CMSSM spectrum FlexibleSUSY vs. Softsusy
 $\tan(\beta) = 5$, $m_{1/2} = 500$ GeV, $A_0 = 0$, $\text{sign}(\mu) = +$



Conclusions

FlexibleSUSY

- is **Modular** (C++ classes, easy to modify and extend)
- is **Fast** (CMSSM runtime $\approx 0.1s$)
- Provides **different RGE solvers**
 - two-scale running (adaptive Runge-Kutta)
 - lattice method + variants

Currently supported models:

- MSSM, NMSSM, SMSSM, UMSSM, E6SSM, MRSSM

Future plans:

- determination of $\sin \theta_W$ from G_μ
- two-loop leading log Higgs mass corrections
- tower of effective field theories